

# Synthesis of scalar-wave envelopes in two-dimensional weakly anisotropic random media by using the Markov approximation

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## SUMMARY

Due to the Earth's inhomogeneity, seismogram envelopes increase in duration and decrease in amplitude with increasing travel distance. Those features have been explained on the basis of the forward scattering in random media using the Markov approximation. Since the conventional studies assumed isotropic random media, they were not realistic enough to represent the anisotropic lithosphere characterized by long horizontal and short vertical correlation distances. Using the Markov approximation, the present study formulates the envelope synthesis in 2-D weakly anisotropic random media characterized by Gaussian and von Kármán-type autocorrelation functions (ACF). Also, wave propagation is numerically simulated with the finite-difference (FD) method; 2 Hz Ricker wavelets propagate through the random media characterized by 4 km s<sup>-1</sup> average velocity and the Gaussian ACF with 5 km horizontal correlation distance, 2.5 km vertical correlation distance, and 5 per cent rms fractional velocity fluctuation. We find a good coincidence between the envelopes of the Markov approximation and those of the FD simulations, which supports the reliability of the synthesis method using the Markov approximation. The envelopes of the Markov approximation are scaled by using a characteristic time, which is a function of propagation direction and other model parameters. It predicts that envelopes increase in duration and decrease in maximum amplitude more rapidly in the horizontal propagation than in the vertical when the media are characterized by long horizontal and short vertical correlation distances. In the case of the vertical wave propagation, the envelope of the anisotropic random media has shorter duration and larger maximum amplitude than those of the isotropic random media. It means that the intensity of the inhomogeneity is underestimated when one analyses seismogram envelopes of deep events considering anisotropic lithosphere as isotropic random media.

**Key words:** anisotropy, random media, scattering, seismic waves.

## 1 INTRODUCTION

The concept of random media, fluctuation from the average medium property is random in space, is widely accepted for modelling velocity inhomogeneity distributed from the very shallow crust to the upper mantle. The inhomogeneity at the very shallow crust can be directly measured in boreholes; fluctuations of *P*- and *S*-wave velocities can be considered as random (e.g. Shiomi *et al.* 1997), and are considered to be anisotropic with long horizontal and short vertical correlation distances (e.g. Wu *et al.* 1994; Dolan *et al.* 1998). Holliger & Levander (1992) reported the anisotropic inhomogeneity in the lower crust from a surface exposure at the Ivrea Zone, Italy. Analyses of seismic array records also detect the anisotropic inhomogeneity in the lower crust (e.g. Hestholm *et al.* 1994; Matsumoto *et al.* 2001). Ryberg *et al.* (2000) employed the anisotropic random media in the upper mantle to explain the *P<sub>n</sub>*-wave trains appearing in the observed seismograms of peaceful nuclear explosion. Furumura & Kennett (2005) distributed anisotropic inhomogeneity in the subducting plate to simulate regional seismograms observed in northern Japan. Kamei *et al.* (2005) proposed the anisotropic random media for modelling methane hydrate-bearing zones. In this study, we use the term 'anisotropic random media' to represent the random media where the characteristic scale length of the inhomogeneity depends on direction. It should be noted that some papers used the term 'anisomeric media' to distinguish the anisotropy of elastic tensor (e.g. Kravtsov *et al.* 2003). This study does not consider the anisotropy of the elastic tensor.

For interpretations of observed seismograms, it is useful to simulate wave propagation in anisotropic random media and to investigate its effects on wavefield. Iooss (1998) and Müller & Shapiro (2003) theoretically investigated traveltimes and amplitude fluctuations in anisotropic random media on the bases of the Rytov method (e.g. Chernov 1960; Rytov *et al.* 1989). In those papers, they compared the theoretical

predictions using the Rytov method with those from finite-difference (FD) wave simulations, and found a good coincidence between them. Supposing the geometry of reflection seismic survey, Kravtsov *et al.* (2003) calculated traveltime fluctuation in anisotropic random media as a function of the incident angle. Scattered waves also strongly depend on the anisotropic inhomogeneity. Numerical simulations of wavefields with the FD method show that the anisotropy significantly affects the excitation of later phases (e.g. Ikelle *et al.* 1993). Jannaud *et al.* (1992) investigated coda-wave excitation in anisotropic random media on the basis of the single-back scattering model with the Born approximation; they also mentioned a possibility of the estimation of the anisotropy from coda-wave analysis.

Due to the wave scattering and the excitation of later phases, wave envelopes usually have long duration compared to source duration. Waves impulsively radiated from a source fluctuate due to the inhomogeneous media and the envelopes show increases in duration with travel distances. This is one of the prominent features of wave envelopes observed in inhomogeneous media. Using the parabolic approximation and the Markov approximation, we can theoretically synthesize wave envelopes without numerical wave propagation simulations (e.g. Ishimaru 1978; Sato & Fehler 1998). The method considers small-angle scattering around the forward direction and neglects large-angle scattering; small-angle scattering dominates over large-angle scattering when the characteristic scale of the inhomogeneity is larger than the wavelength. We refer to the envelopes synthesized by using the Markov approximation as Markov envelopes, hereinafter. Fehler *et al.* (2000) confirmed the validity of this method by comparing the Markov envelopes with the envelopes of wave traces from FD simulations in 2-D case. Sato (1989) first used the Markov envelopes to explain the increase of seismogram-envelope duration with travel distance, which is referred to as envelope broadening. Later, Saito *et al.* (2002) extended the method under more realistic assumptions. They explained not only envelope broadening but also maximum-amplitude decay of envelopes (Saito *et al.* 2005). Based on the envelope modelling, one can estimate the power spectral density function (PSDF) of velocity inhomogeneity in the lithosphere (e.g. Scherbaum & Sato 1991; Obara & Sato 1995). Also, one can roughly estimate the hypocentral distance analysing the initial part of a *P*-wave envelope observed at a single station (Tsukada *et al.* 2004). Recently, the Markov envelope was incorporated in the radiative transfer theory to simulate not only small-angle scattering but also large-angle scattering (Saito *et al.* 2003; Saito *et al.* 2004). Furthermore, Korn & Sato (2005) extended the Markov envelope to the elastic wave propagation case; they showed that the scalar-wave equation can be used for the envelope synthesis even in elastic media when the conversion scattering is negligibly small. We may say that the envelope modelling using the Markov approximation is one of the useful methods for interpreting seismogram envelopes. However, the conventional studies have supposed only isotropic inhomogeneity.

The present study formulates a method of scalar-wave envelope synthesis in anisotropic 2-D random media using the Markov approximation. In Section 2, we mention 2-D wave propagation in anisotropic random media and statistical characterization of the media. In Section 3, we numerically calculate the wave propagation with the FD method to obtain reference envelopes (FD envelopes) that will be compared with the Markov envelopes. In Section 4, the envelope synthesis method is formulated using the Markov approximation. We compare the Markov envelopes and the FD envelopes to examine the reliability of the theoretical calculations. In Section 5, we discuss the dependence of envelopes on the propagation direction in the anisotropic random media. Also, we compare the envelopes in anisotropic random media and those in isotropic random media. Finally, Section 6 presents some conclusions.

## 2 WAVE PROPAGATION IN 2-D ANISOTROPIC RANDOM MEDIA

Let us consider propagation of scalar wave  $u(\mathbf{x}, t)$  through 2-D space governed by

$$\left(\Delta - \frac{1}{V(\mathbf{x})^2} \frac{\partial^2}{\partial t^2}\right)u(\mathbf{x}, t) = 0, \quad (1)$$

where  $\mathbf{x} = (x, z)$  and  $\Delta$  is the Laplacian in 2-D. In randomly inhomogeneous media, wave velocity is written as  $V(\mathbf{x}) = V_0\{1 + \xi(\mathbf{x})\}$ , where  $V_0$  is the average velocity and  $\xi$  the fractional velocity fluctuation. Throughout this study, weak fluctuation  $|\xi| \ll 1$  is assumed. To obtain statistical properties of the media and wavefield, we consider an ensemble of the fractional fluctuation  $\{\xi(\mathbf{x})\}$  such that  $\langle \xi(\mathbf{x}) \rangle = 0$ , where the angular brackets mean the ensemble average. The media are referred to as random media, which are statistically characterized by the autocorrelation function (ACF) of the fractional fluctuation  $R(\mathbf{x}_d) \equiv \langle \xi(\mathbf{x}_d + \mathbf{x})\xi(\mathbf{x}) \rangle$  or the power spectral density function (PSDF),  $P(\mathbf{m})$ , where  $\mathbf{m}$  is the wavenumber. We suppose that  $\xi$  is statistically homogeneous random function in space; the ACF depends on difference vector  $\mathbf{x}_d$  but is independent of location  $\mathbf{x}$ . The ACF and the PSDF are generally characterized by rms value of the fractional velocity fluctuation  $\varepsilon$  and two correlation distances  $a_x$  and  $a_z$ , which are the characteristic scale lengths of the inhomogeneity in the  $x$  and  $z$  directions, respectively.

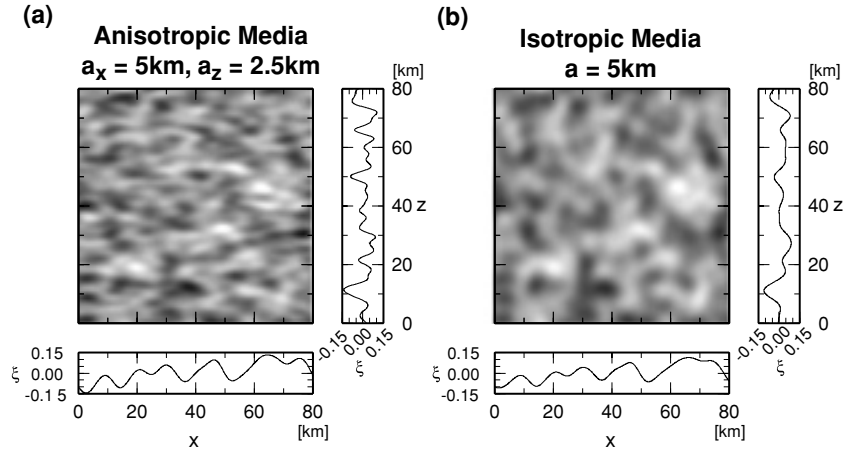
A Gaussian ACF for 2-D anisotropic random media is given by

$$R(x_d, z_d) = \varepsilon^2 \exp\left(-\frac{x_d^2}{a_x^2} - \frac{z_d^2}{a_z^2}\right), \quad (2)$$

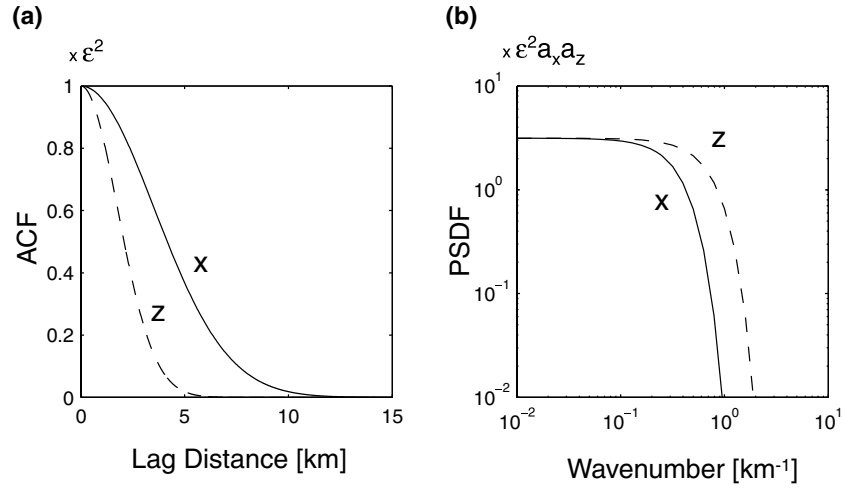
and the corresponding PSDF is given by

$$\begin{aligned} P(m_x, m_z) &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} R(x_d, z_d) \exp(-im_x x_d - im_z z_d) dx_d dz_d \\ &= \varepsilon^2 \pi a_x a_z \exp\left(-\frac{a_x^2 m_x^2 + a_z^2 m_z^2}{4}\right), \end{aligned} \quad (3)$$

where  $(m_x, m_z)$  is the wavenumber vector of the inhomogeneity. Fig. 1(a) shows an example of anisotropic random media characterized by the Gaussian ACF. In Fig. 1(b), an example of isotropic random media is also shown for comparison. The ACF and PSDF of the anisotropic



**Figure 1.** The random media characterized by the Gaussian autocorrelation function (ACF) with  $\varepsilon = 0.05$ . (a) Anisotropic random media with  $a_x = 5$  km,  $a_z = 2.5$  km and (b) isotropic random media with  $a_x = a_z = 5$  km. The fractional velocity fluctuation is imaged by grey scale and the fluctuations along  $x = 40$  km and  $z = 40$  km are also plotted beside the image.



**Figure 2.** (a) The Gaussian ACF with  $a_x = 5$  km and  $a_z = 2.5$  km. A cross section along  $x$ -axis is plotted by a solid curve and along  $z$ -axis by a dashed curve. (b) The corresponding power spectral density function (PSDF). A cross section along  $x$ -axis is plotted by a solid curve and along  $z$ -axis by a dashed curve.

random media are shown in Fig. 2. The correlation distances in the  $x$ - and  $z$ -axes are 5.0 and 2.5 km, respectively. The ACF decreases more rapidly in the  $z$ -axis than in the  $x$ -axis (Fig. 2a). The PSDF along the  $z$ -axis is richer in the short-wavelength components than that along the  $x$ -axis (Fig. 2b).

A von Kármán type ACF is given by

$$R(x_d, z_d) = \frac{\varepsilon^2 2^{1-\kappa}}{\Gamma(\kappa)} \left( \sqrt{(x_d/a_x)^2 + (z_d/a_z)^2} \right)^\kappa K_\kappa \left( \sqrt{(x_d/a_x)^2 + (z_d/a_z)^2} \right), \quad (4)$$

and the corresponding PSDF is given by

$$P(m_x, m_z) = \frac{4\pi\kappa\varepsilon^2 a_x a_z}{(1 + a_x^2 m_x^2 + a_z^2 m_z^2)^{\kappa+1}}, \quad (5)$$

where  $\Gamma$  and  $K_\kappa$  are the Gamma function and the modified Bessel function of the second kind of the order  $\kappa$ , respectively.

In general, the ACFs for 2-D anisotropic random media, including Gaussian and von Kármán type, can be represented by using a normalized ACF  $\bar{R}(\bar{r})$ ,

$$R(x_d, z_d) = \varepsilon^2 \bar{R} \left( \sqrt{\frac{x_d^2}{a_x^2} + \frac{z_d^2}{a_z^2}} \right) = \varepsilon^2 \bar{R}(\bar{r}), \quad (6)$$

where the normalized difference distance is defined as

$$\bar{r} \equiv \sqrt{\frac{x_d^2}{a_x^2} + \frac{z_d^2}{a_z^2}}. \quad (7)$$

The corresponding PSDF in 2-D is given by

$$\begin{aligned}
 P(m_x, m_z) &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} R(x_d, z_d) \exp(-im_x x_d - im_z z_d) dx_d dz_d \\
 &= 2\pi \varepsilon^2 a_x a_z \int_0^{\infty} J_0(\bar{m}\bar{r}) \bar{r} \bar{R}(\bar{r}) d\bar{r} \\
 &= \varepsilon^2 a_x a_z \bar{P}(\bar{m}),
 \end{aligned} \tag{8}$$

and  $J_0$  is the Bessel function of zero order. The normalized PSDF is defined as

$$\bar{P}(\bar{m}) \equiv 2\pi \int_0^{\infty} J_0(\bar{m}\bar{r}) \bar{r} \bar{R}(\bar{r}) d\bar{r}, \tag{9}$$

where the normalized wavenumber is defined as

$$\bar{m} \equiv \sqrt{a_x^2 m_x^2 + a_z^2 m_z^2}. \tag{10}$$

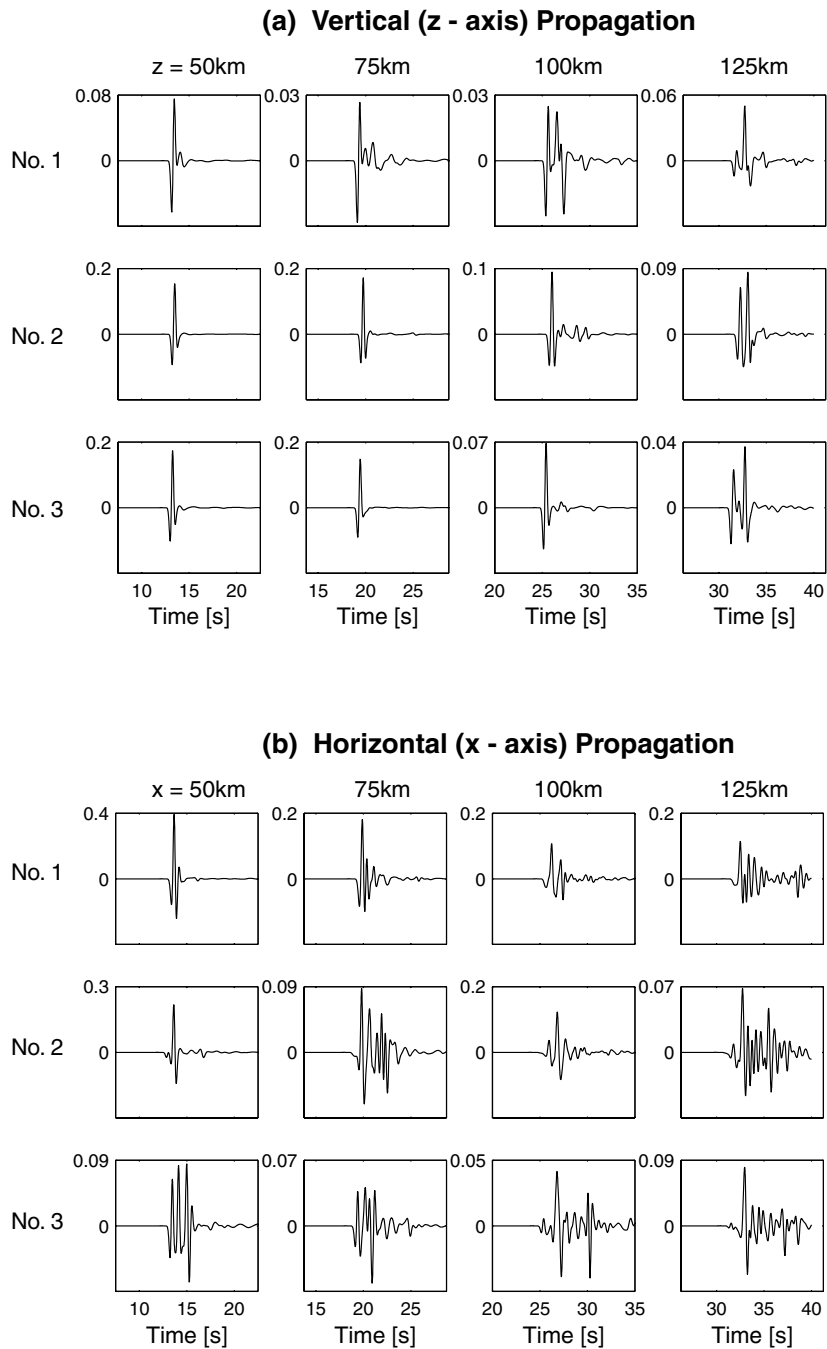
Using the above expressions of the ACF (eq. 6) and PSDF (eq. 8) with the normalized functions, we can represent general type of 2-D anisotropic random media.

### 3 FD ENVELOPES IN 2-D ANISOTROPIC RANDOM MEDIA

Wave propagation in 2-D anisotropic random media is numerically simulated before the derivation of the Markov envelopes. Those numerical simulations would be helpful for intuitively recognizing wave propagation in anisotropic random media. Also, the envelopes obtained from the numerically calculated wavefield will be compared with the Markov envelopes in order to examine the reliability of the synthesis using the Markov approximation. We here consider the random media characterized by the Gaussian ACF with  $\varepsilon = 0.05$ ,  $a_x = 5.0$  km,  $a_z = 2.5$  km and  $V_0 = 4$  km s<sup>-1</sup>. The orders of those values are consistent with the observed *S*-wave random inhomogeneity (e.g. Saito *et al.* 2005) and the parameters used in the previous works (e.g. Sato *et al.* 2004). The ratio  $a_x/a_z = 2$  is consistent with that of the KTB borehole data obtained by Wu *et al.* (1994). In the medium with dimensions of 200 km square surrounded by absorbing boundaries, wave propagation from a point source with isotropic radiation is numerically simulated with the FD method. We use a Ricker wavelet with its dominant frequency  $f_c = 2$  Hz as a source time function. The dominant frequency corresponds to  $ka_x \approx 16$  and  $ka_z \approx 8$  where  $k$  is the wavenumber given by  $k = 2\pi f_c/V_0$ . The FD simulations are accomplished with fourth-order accuracy in space and second-order in time with grid spacing of 50 m and time step of 4 ms.

To investigate statistical properties of wavefield, we make the above numerical simulations with respect to 50 realizations of the random media. As examples, Fig. 3 shows wave traces calculated by the FD simulations with respect to 3 realizations of the random media. The wave records after the arrival of reflected waves from the boundaries are excluded. Figs 3(a) and (b) show wave traces along the  $z$ - and  $x$ -axes, respectively. We consider Figs 3(a) and (b) as the vertical wave propagation and the horizontal propagation, respectively. Each row corresponds to each realization of the random media and the four panels in each row correspond to the distances, 50, 75, 100 and 125 km apart from the source. Wave traces, on the whole, become more complex with increasing propagation distance; later arrivals lengthen the duration of wave trains. This tendency is more prominent in the horizontal propagation cases, Fig. 3(b), than in the vertical propagation cases, Fig. 3(a).

From wave traces calculated with respect to 50 realizations of random media, we make mean-square (MS) envelopes and rms envelopes, as follows. We calculate 50 envelopes at each station by using the Hilbert transform (p. 250, Shearer 1999). At each station, squaring those 50 realized envelopes and averaging the squared envelopes, we obtain an ensemble-averaged squared envelope, which is referred to as MS envelope. Taking the square root of the MS envelope, we obtain rms envelope. The rms envelopes from FD simulations are referred to as FD envelopes, hereinafter. Fig. 4 shows the FD envelopes at stations located 50, 75, 100 and 125 km apart from the source. Each bin corresponds to each propagation direction  $\phi$ ; the directions,  $\phi = 0, 30, 45, 60$  and  $90$  degree, are measured as the angle between the  $z$ -axis and the global ray direction. Solid black curves are the FD envelopes or ensemble-averaged envelopes made from 50 realized envelopes. Grey area around the solid curve indicates the range of standard deviation of the ensemble-averaged envelope. The range is estimated with bootstrap method; we estimate the standard deviation over 30 ensemble-averaged envelopes, which are calculated from artificial data sets. Each artificial data set consists of 40 envelopes, which are randomly sampled from the 50 realized envelopes. Fine grey curves indicate the standard deviation of the each realized envelope. Note that the large standard deviation of each realized envelope (fine grey curves in Fig. 4) means that the shapes of a single-realized envelope and the ensemble-averaged envelope are quite different. Although the standard deviation of each realized envelope is large, we see that the shape of the ensemble-averaged envelope is stable (grey areas in Fig. 4) when we use more than 40 realized envelopes for the average. In this study, we are interested in the ensemble-averaged envelope. The FD envelopes, or ensemble-averaged envelopes, increase in their duration and decrease in their maximum amplitude with increasing travel distance. This tendency is more prominent with increasing the angle  $\phi$ .



**Figure 3.** Wave traces calculated by the finite-difference simulations with respect to 3 realizations of random media characterized by the Gaussian ACF with  $\varepsilon = 0.05$ ,  $a_x = 5$  km,  $a_z = 2.5$  km and  $V_0 = 4$  km s<sup>-1</sup>. (a) Wave traces at stations located at 50, 75, 100 and 125 km on the z-axis apart from a source. A set of four panels in each row is calculated for a single realization of the random media. (b) Wave traces at stations located at 50, 75, 100 and 125 km on the x-axis apart from a source.

## 4 MARKOV ENVELOPE IN 2-D RANDOM MEDIA

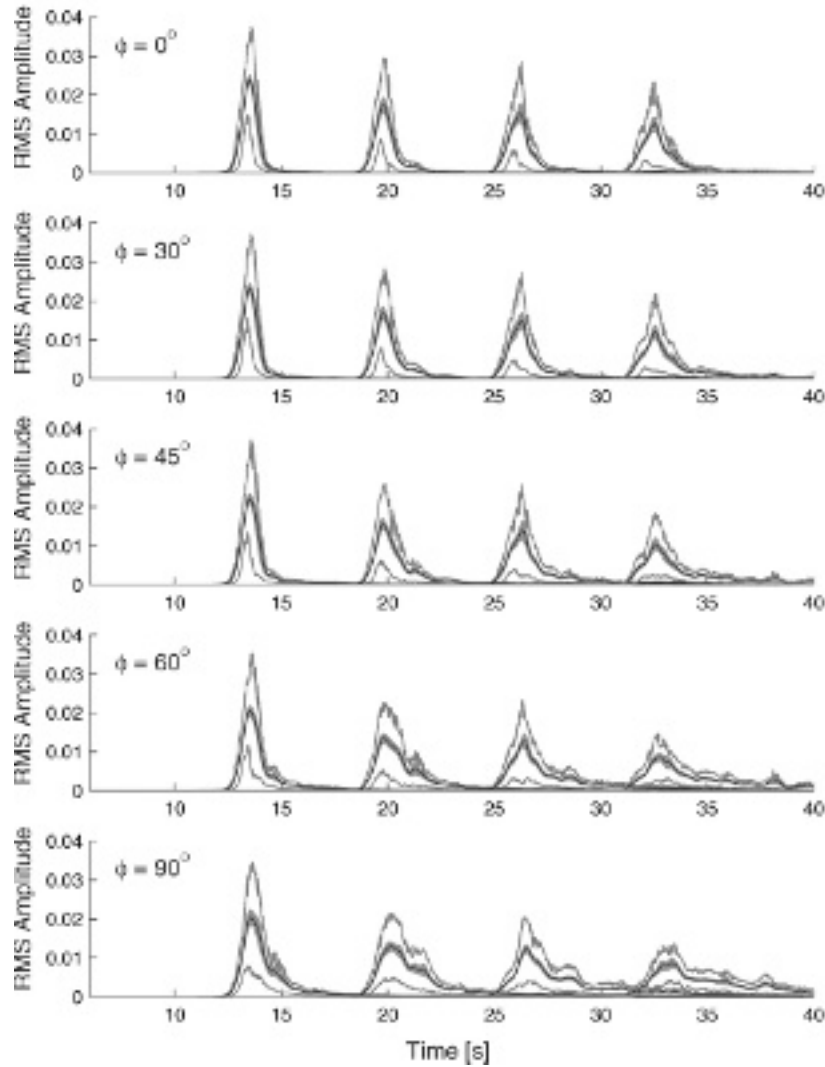
### 4.1 Formulation

#### 4.1.1 General formulation for 2-D random media

Using 2-D polar coordinates  $(r, \phi)$  shown in Fig. 5(a), we write wavefield as a superposition of harmonic cylindrical waves of angular frequency  $\omega$ ,

$$u(\mathbf{x}, t) = \frac{1}{2\pi} \int_{-\infty}^{\infty} \frac{U(r, \phi, \omega)}{\sqrt{kr}} \exp\{i(kr - \omega t)\} d\omega, \quad (11)$$

**RMS Envelopes in 2-D Anisotropic Random Media**  
**Gaussian PSDF ( $a_x = 5\text{km}$ ,  $a_z = 2.5\text{km}$ ,  $\varepsilon = 0.05$   $V_0 = 4\text{km s}^{-1}$ )**



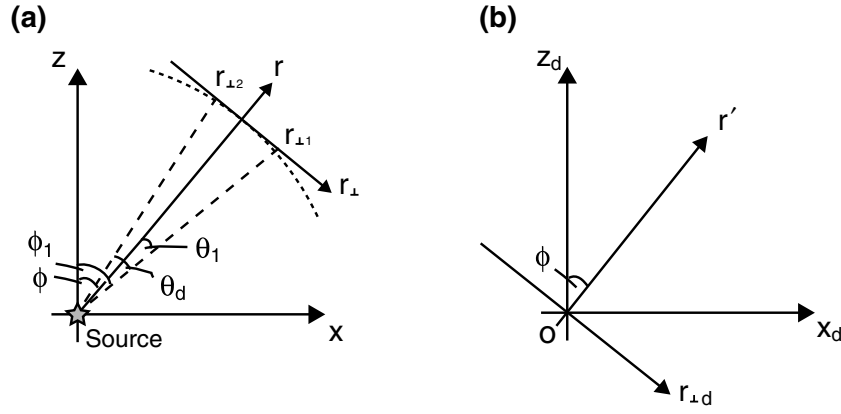
**Figure 4.** The rms envelopes of finite-difference wave simulations. The envelopes at travel distance 50, 75, 100 and 125 km are plotted. Numerals at upper left side in each bin show the propagation direction measured from the  $z$ -axis. Wave propagations from a 2 Hz Ricker wavelet source are numerically simulated for 50 realizations of the random media characterized by the Gaussian ACF with  $\varepsilon = 0.05$ ,  $a_x = 5$  km,  $a_z = 2.5$  km and  $V_0 = 4$  km  $\text{s}^{-1}$ . Solid black curves are the FD envelopes or ensemble-averaged envelopes over 50 realized envelopes. Grey area around the solid curve indicates the range of standard deviation of the FD envelope. Fine grey curves indicate the standard deviation of the each realized envelopes.

where the wavenumber  $k = \omega/V_0$  (e.g. Saito *et al.* 2003). When the correlation distance of the medium inhomogeneity is much larger than the wavelength  $ak \gg 1$ , the second derivative of  $U(r, \phi, \omega)$  with respect to radial direction  $r$  may be neglected. In addition, when wave travels much longer than the wavelength  $rk \gg 1$ , the parabolic wave equation is obtained from eqs (1) and (11),

$$2ik \frac{\partial}{\partial r} U(r, \phi, \omega) + \frac{1}{r^2} \frac{\partial^2}{\partial \phi^2} U(r, \phi, \omega) - 2k^2 \xi(r, \phi) U(r, \phi, \omega) = 0. \quad (12)$$

For the case of  $ak \gg 1$ , small-angle scattering dominates. Hence, we consider wave propagation around the global ray direction  $\phi$ . In Fig. 5(a), the angle  $\phi_i$  is represented by small angle  $\theta_i$  as  $\phi_i = \theta_i + \phi$  where  $|\theta_i| \ll 1$ . Furthermore, a local coordinate  $r_{\perp}$  is introduced at a large distance  $r$  from the source ( $r \gg a$ ); the axis  $r_{\perp}$ , which is referred to as transverse axis, is tangential to the circle of radius  $r$  (Fig. 5a). In order to derive the envelopes, we introduce a two-frequency mutual coherence function (TFMCF) in the local coordinate. The TFMCF is defined at two adjacent points  $r_{\perp 1}$  and  $r_{\perp 2}$  on the transverse axis and two angular frequencies  $\omega_1$  and  $\omega_2$  as

$$\Gamma_2(r_{\perp 1}, r_{\perp 2}, r, \omega_1, \omega_2) \equiv \langle U(r_{\perp 1}, r, \omega_1) U^*(r_{\perp 2}, r, \omega_2) \rangle, \quad (13)$$



**Figure 5.** Coordinates used in the formulation of 2-D anisotropic media ( $x$ - $z$  space). (a) The global ray direction from a source is in the radial direction  $r$ . A transverse axis  $r_{\perp}$ , which is tangential to the circle of radius  $r$  is taken at large distance  $r$  from the source. (b) Difference coordinates between two points  $(r_{\perp 1}, r)$  and  $(r_{\perp 2}, r)$ ; the origin is chosen at  $(r_{\perp 2}, r)$ . An axis  $r'$  is parallel to the global ray direction and an axis  $r_{\perp d}$  is in the transverse axis  $r_{\perp}$ . Axes  $x_d$  and  $z_d$  are parallel to the axes  $x$  and  $z$ , respectively.

where asterisk denotes complex conjugate and  $r_{\perp i} \approx r\theta_i$ . The equation of  $\Gamma_2$  is derived from eq. (12) as

$$2i \frac{\partial}{\partial r} \Gamma_2(r_{\perp 1}, r_{\perp 2}, r, \omega_1, \omega_2) + \frac{1}{r^2} \left( \frac{1}{k_1} \frac{\partial^2}{\partial \phi_1^2} - \frac{1}{k_2} \frac{\partial^2}{\partial \phi_2^2} \right) \Gamma_2 - 2 \{ k_1 \xi(r_{\perp 1}, r) - k_2 \xi(r_{\perp 2}, r) \} U(r_{\perp 1}, r, \omega_1) U^*(r_{\perp 2}, r, \omega_2) = 0. \quad (14)$$

When we consider the range such that  $|r_{\perp 1} - r_{\perp 2}| \ll r$  in homogeneous random media,  $\Gamma_2$  strongly depends on the difference angle  $\theta_d \equiv \theta_1 - \theta_2$ , while it is almost independent of the centre of mass coordinate  $\theta_c \equiv (\theta_1 + \theta_2)/2$  within the local coordinates. Hence, the derivatives with respect to angles  $\phi_i$  are approximated as  $\partial^2 / \partial \phi_1^2 \approx \partial^2 / \partial \phi_2^2 \approx \partial^2 / \partial \theta_d^2$ . Considering the two points  $(r_{\perp 1}, r)$  and  $(r_{\perp 2}, r)$  on the transverse axis, we introduce difference coordinates  $x_d = x_1 - x_2$  and  $z_d = z_1 - z_2$  at distance  $r$  where  $x_i$  and  $z_i$  indicate the position of  $(r_{\perp i}, r)$  in the  $x$ - $z$  coordinates (Figs 5a and b). Introducing difference coordinates  $(r_{\perp d}, r')$ , where the  $r_{\perp d}$ -axis is parallel to the  $r_{\perp}$ -axis and the  $r'$ -axis is parallel to the  $r$ -axis, we define the longitudinal integral of the ACF along the global ray direction as

$$A(r_{\perp d}, \phi) \equiv \int_{-\infty}^{\infty} R(\mathbf{x}_d) dr' = \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} dr' \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} P(\mathbf{m}) \exp\{i\mathbf{m} \cdot \mathbf{x}_d\} d\mathbf{m} \\ = \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} dr' \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} P(\mathbf{m}) \exp\{i\mathbf{m} \cdot (r_{\perp d} \mathbf{e}_{\perp} + r' \mathbf{e}_r)\} d\mathbf{m}, \quad (15)$$

where  $r_{\perp d} = r_{\perp 1} - r_{\perp 2}$ , the vectors  $\mathbf{e}_r$  and  $\mathbf{e}_{\perp}$  are unit vectors in the  $r'$ - and  $r_{\perp d}$ -axes, respectively. The difference vector  $\mathbf{x}_d = (x_1 - x_2, z_1 - z_2)$  is given by  $\mathbf{x}_d = r_{\perp d} \mathbf{e}_{\perp} + r' \mathbf{e}_r$ . The third term in eq. (14) is approximated as

$$\langle (k_1 \xi_1 - k_2 \xi_2) U_1 U_2^* \rangle \approx -\frac{i}{2} [(k_1^2 + k_2^2) A(0, \phi) - 2k_1 k_2 A(r_{\perp d}, \phi)] \Gamma_2, \quad (16)$$

where  $U_i$  means that its argument is  $(r_{\perp i}, r, \omega_i)$  and  $\xi_i$  means that its argument is  $(r_{\perp i}, r)$  (see Appendix A). During the derivation, large-angle scattering, or backward scattering, is neglected. Also, we suppose weak anisotropy and the existence of an intermediate scale  $\Delta r$ , which is larger than the correlation distance of inhomogeneity but smaller than the scale of variation of  $U$ . The approximation is called the Markov approximation (e.g. Sato & Fehler 1998). From eqs (14) and (16), the master equation of the TFMCF, for quasi-monochromatic waves  $\omega_1 \approx \omega_2$ , is written as

$$\frac{\partial}{\partial r} \Gamma_2(\theta_d, r, \omega_c, \omega_d; \phi) + i \frac{k_d}{2k_c^2} \frac{1}{r^2} \frac{\partial^2}{\partial \theta_d^2} \Gamma_2 + k_c^2 [A(0, \phi) - A(r_{\perp d}, \phi)] \Gamma_2 + \frac{k_d^2}{2} A(0, \phi) = 0, \quad (17)$$

where we use  $r_{\perp d} \approx r\theta_d$ . Also, the centre of mass and the difference wavenumbers are  $k_c = (k_1 + k_2)/2$  and  $k_d = k_1 - k_2$ , respectively. Corresponding coordinates for the angular frequency are also used. We introduced  ${}_0\Gamma_2$  as  $\Gamma_2 = {}_0\Gamma_2 \exp\{-k_d^2 A(0, \phi) r/2\}$ . The term  $\exp\{-k_d^2 A(0, \phi) r/2\}$  is referred to as the wandering effect term, which corresponds to the travel time fluctuation over each realization of the envelopes (e.g. Lee & Jokipii 1975b). From eq. (17), we obtain the master equation for  ${}_0\Gamma_2$  as,

$$\frac{\partial}{\partial r} {}_0\Gamma_2(\theta_d, r, \omega_c, \omega_d; \phi) + i \frac{k_d}{2k_c^2} \frac{1}{r^2} \frac{\partial^2}{\partial \theta_d^2} {}_0\Gamma_2 + k_c^2 [A(0, \phi) - A(r_{\perp d}, \phi)] {}_0\Gamma_2 = 0. \quad (18)$$

The squared envelope in a narrow-frequency band  $\Delta\omega_c$  around the centre angular frequency  $\omega_c$  is written as

$$I(r, t, \omega_c, \phi) \equiv \langle |u|^2 \rangle \\ \approx \frac{\Delta\omega_c}{(2\pi)^2 k_c r} \int_{\omega_c - \Delta\omega_c/2}^{\omega_c + \Delta\omega_c/2} {}_0\Gamma_2(\theta_d = 0, r, \omega_d, \omega_c; \phi) \exp\{-i\omega_d(t - r/V_0)\} d\omega_d. \quad (19)$$

By solving the master equation of TFMCF (18) and inserting the solved TFMCF into eq. (19), we can synthesize the MS envelope without the wandering effect. Square root of the MS envelope is the rms envelope. We refer to the rms envelope derived from the Markov approximation as Markov envelope, hereinafter.

#### 4.1.2 Anisotropic random media

We derive the master equation of TFMCF and the Markov envelope in the case of anisotropic random media. From Fig. 5(b), the relation between  $(x_d, z_d)$  and  $(r_{\perp d}, r')$  is given by

$$x_d = r_{\perp d} \cos \phi + r' \sin \phi \quad \text{and} \quad z_d = -r_{\perp d} \sin \phi + r' \cos \phi. \quad (20)$$

Using eq. (20), we calculate eq. (15) as

$$\begin{aligned} A(r_{\perp d}, \phi) &\equiv \int_{-\infty}^{\infty} R(x_d, z_d) dr' \\ &= \int_{-\infty}^{\infty} dr' \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} P(m_x, m_z) \exp(im_x x_d + im_z z_d) dm_x dm_z \\ &= \frac{1}{(2\pi)^2} \int_{-\infty}^{\infty} dr' \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} P(m_x(m_r, m_{\perp}, \phi), m_z(m_r, m_{\perp}, \phi)) \times \exp(im_r r' + im_{\perp} r_{\perp d}) dm_r dm_{\perp} \\ &= \frac{1}{2\pi} \int_{-\infty}^{\infty} P(m_x(m_r = 0, m_{\perp}, \phi), m_z(m_r = 0, m_{\perp}, \phi)) \exp(im_{\perp} r_{\perp d}) dm_{\perp}, \end{aligned} \quad (21)$$

where

$$m_{\perp} \equiv m_x \cos \phi - m_z \sin \phi \quad \text{and} \quad m_r \equiv m_x \sin \phi + m_z \cos \phi. \quad (22)$$

Substituting eq. (8) into eq. (21) and considering the coefficient of  $m_{\perp}$  in the PSDF as an correlation distance  $a_t$  along the  $r_{\perp d}$ -axis, we obtain

$$\begin{aligned} A(r_{\perp d}, \phi) &= \frac{1}{2\pi} \varepsilon^2 a_x a_z \int_{-\infty}^{\infty} \bar{P}(a_t m_{\perp}) \exp(im_{\perp} r_{\perp d}) dm_{\perp} \\ &= \frac{1}{2\pi} \varepsilon^2 a_r \int_{-\infty}^{\infty} \bar{P}(\bar{m}_{\perp}) \exp(i\bar{m}_{\perp} r_{\perp d} / a_t) d\bar{m}_{\perp}, \end{aligned} \quad (23)$$

with the normalized PSDF (eq. 9) where  $\bar{m}_{\perp} \equiv a_t m_{\perp}$ . Here, the values of  $a_t$  and  $a_r$  are effective correlation distances in the  $r_{\perp}$ - and  $r$ -axes, which are defined as

$$a_t \equiv \sqrt{a_x^2 \cos^2 \phi + a_z^2 \sin^2 \phi} \quad \text{and} \quad a_r \equiv a_x a_z / \sqrt{a_x^2 \cos^2 \phi + a_z^2 \sin^2 \phi}, \quad (24)$$

respectively. The effective correlation distances are the functions of the angle  $\phi$ ;  $a_t = a_x$  and  $a_r = a_z$  when  $\phi = 0$ , and  $a_t = a_z$  and  $a_r = a_x$  when  $\phi = \pi/2$ .

The non-dimensional PSDF's of the Gaussian and von Kármán-type ACFs are given by

$$\bar{P}(\bar{m}) = \begin{cases} \pi \exp(-\bar{m}^2/4) & \text{Gaussian.} \\ \frac{4\pi\kappa}{(1 + \bar{m}^2)^{\kappa+1}}, \quad \kappa > 0 & \text{von Karman.} \end{cases} \quad (25)$$

For the case of the Gaussian ACF, eq. (23) is given by

$$A(r_{\perp d}, \phi) = \sqrt{\pi} \varepsilon^2 a_r \exp\left(-\frac{r_{\perp d}^2}{a_t^2}\right). \quad (26)$$

At a long distance from the source, the correlation of the wavefield at two points spatially separated on the transverse axis rapidly decreases to zero with increasing the lag distance (see p. 245, Sato & Fehler 1998). Hence, the value of  $A(0, \phi) - A(r_{\perp d}, \phi)$  at small transverse distance  $r_{\perp d} \ll a_t$  is dominant in eq. (18). The value is approximated as

$$A(0, \phi) - A(r_{\perp d}, \phi) \approx \sqrt{\pi} \varepsilon^2 a_r \left(\frac{r_{\perp d}}{a_t}\right)^2. \quad (27)$$

For the case of the von Kármán type ACF, eq. (23) is given by

$$A(r_{\perp d}, \phi) = \frac{2^{3/2-\kappa} \sqrt{\pi} \varepsilon^2 a_r}{\Gamma(\kappa + 1)} \left(\frac{r_{\perp d}}{a_t}\right)^{\kappa+1/2} K_{\kappa+1/2}\left(\frac{r_{\perp d}}{a_t}\right), \quad (28)$$

and the value of  $A(0, \phi) - A(r_{\perp d}, \phi)$  is approximated as

$$A(0, \phi) - A(r_{\perp d}, \phi) \approx C(\kappa) \varepsilon^2 a_r \left(\frac{r_{\perp d}}{a_t}\right)^{p(\kappa)}, \quad (29)$$

where  $C(\kappa)$  and  $p(\kappa)$  are numerically calculated in Saito *et al.* (2002). From eqs (27), (29) and (18), the master equation of TFMCF is given by



**Table 1.** Parameters  $C$  and  $p$  for different type of random media.

Random media type	$C$	$p$
Gaussian ACF	$\sqrt{\pi}$	2
$\kappa$ (von Kármán-type ACF)		
0.1	0.56	1.19
0.2	1.06	1.38
0.3	1.56	1.56
0.4	2.00	1.71
0.5	2.28	1.83
0.6	2.31	1.91
0.7	2.14	1.95
0.8	1.90	1.98
0.9	1.68	1.99
1.0	1.50	1.99

$$\frac{\partial}{\partial r} {}_0\Gamma_2(\theta_d, r, \omega_c, \omega_d; \phi) + i \frac{k_d}{2k_c^2} \frac{1}{r^2} \frac{\partial^2}{\partial \theta_d^2} {}_0\Gamma_2 + k_c^2 C \varepsilon^2 a_r \left( \frac{r\theta_d}{a_t} \right)^p {}_0\Gamma_2 = 0, \quad (30)$$

for the cases of the Gaussian and von Kármán-type ACFs. The values of  $C$  and  $p$  depend on the type of random media. For the Gaussian ACF,  $p = 2$  and  $C = \sqrt{\pi}$ . For the von Kármán-type ACF, the values of  $C$  and  $p$  depend on the order  $\kappa$  (Table 1). Eq. (30) is the same as that in isotropic random media except the use of  $a_t$  and  $a_r$  (Fehler *et al.* 2000; Saito *et al.* 2003; Sato *et al.* 2004). Introducing the characteristic time for angular frequency  $\omega_c$  at travel distance  $r_0$  as

$$t_M \equiv \frac{V_0 r_0}{2\omega_c^2 a_t^2} \left( \frac{C \varepsilon^2 \omega_c^2 a_r r_0}{V_0^2} \right)^{2/p}, \quad (31)$$

and the non-dimensional travel distance and difference angle as,

$$\tau \equiv r/r_0 \quad \text{and} \quad \chi \equiv \sqrt{2r_0 \omega_c^2 t_M / V_0} \theta_d, \quad (32)$$

respectively, we rewrite eq. (30) as

$$\frac{\partial}{\partial \tau} {}_0\Gamma_2 + i t_M \omega_d \frac{1}{\tau^2} \frac{\partial^2}{\partial \chi^2} {}_0\Gamma_2 + \tau^p \chi^p {}_0\Gamma_2 = 0. \quad (33)$$

When  $p = 2$ , we can use the analytic solution of  ${}_0\Gamma_2$  derived by Fehler *et al.* (2000) since eq. (33) is the same as eq. (30) in Fehler *et al.* (2000). When  $p \neq 2$ , we numerically solve eq. (33) with the FD method (see Saito *et al.* 2002). Inserting the solved  ${}_0\Gamma_2$  into eq. (19), we can obtain squared-amplitude Markov envelopes without the wandering effect. Here, we consider the case of an isotropic source radiation of the delta-function source time function,  $I(r \rightarrow 0) = 1/(2\pi r) \delta(t - r/V_0)$ . For the corresponding conditions, we take  ${}_0\Gamma_2$  to be  ${}_0\Gamma_2 = k_c/\Delta\omega_c$ , and suppose the integration range of  $\omega_d$  in eq. (19) to be infinity at the source,  $r = 0$ . In practice, the integration range of  $\omega_d$  is finite at a travel distance  $r$  since  ${}_0\Gamma_2$  decreases to zero with increasing  $\omega_d$ . Fig. 6 plots the squared-amplitude Markov envelopes for the radiation condition  $I(r \rightarrow 0) = 1/(2\pi r) \delta(t - r/V_0)$ , where the square amplitude is normalized by the geometrical spreading factor and the characteristic time (eq. 31), and the reduced time is normalized by the characteristic time. We may consider the characteristic time as the duration time of the envelopes. As the characteristic time increases, the duration increases and the amplitude decreases.

#### 4.1.3 Validity range of the Markov envelopes

When we derive eqs (12) and (16), we assume small-angle scattering around the forward direction and neglect large-angle scattering. We evaluate the large-angle scattering to discuss the validity range of the Markov envelopes as follows. By using the Born approximation, the scattering coefficient in 2-D anisotropic random media is calculated as a function of scattering angle  $\theta$  for the case of incident angle  $\phi$  measured from the  $z$ -axis (Jannaud *et al.* 1992):

$$g(\theta) = \varepsilon^2 k^3 a_x a_z \bar{P}(\bar{k}), \quad (34)$$

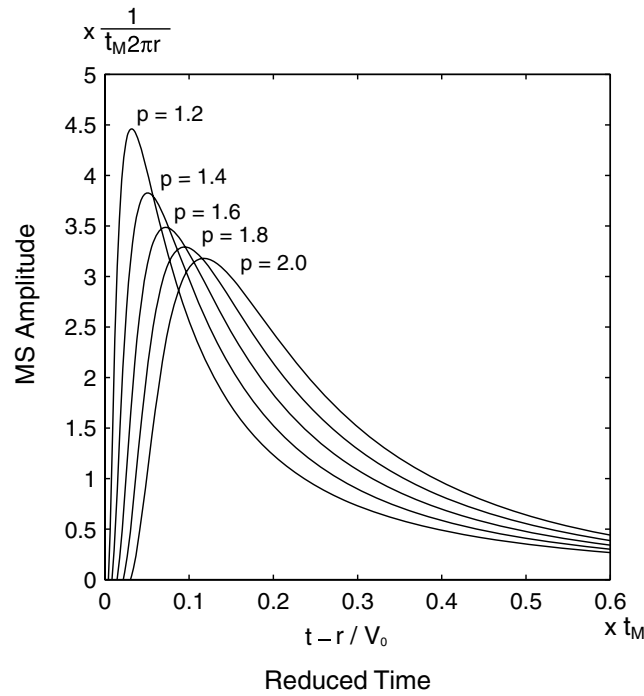
where

$$\bar{k}^2 = k^2 a_x^2 \{\sin(\theta + \phi) - \sin \phi\}^2 + k^2 a_z^2 \{\cos(\theta + \phi) - \cos \phi\}^2. \quad (35)$$

It should be noted that the scattering coefficient (eq. 34) is not always a symmetric function with respect to  $\theta$ , while that of isotropic random media is a symmetric function. Using a cut-off angle  $\theta_c$ , we introduce a value:

$$r_0 g_L \equiv r_0 \frac{1}{2\pi} \int_{|\theta| > \theta_c} g(\theta) d\theta = \varepsilon^2 (r_0 k)(k a_x)(k a_z) \frac{1}{2\pi} \int_{|\theta| > \theta_c} \bar{P}(\bar{k}(\theta)) d\theta, \quad (36)$$

as a measure of large-angle scattering. This value is related to the scattering attenuation for the main part of envelope, or around the maximum amplitude, at travel distance  $r_0$  (Sato 1982; Wu 1982). Also, the value is related to coda-wave excitation since the attenuated wave energy



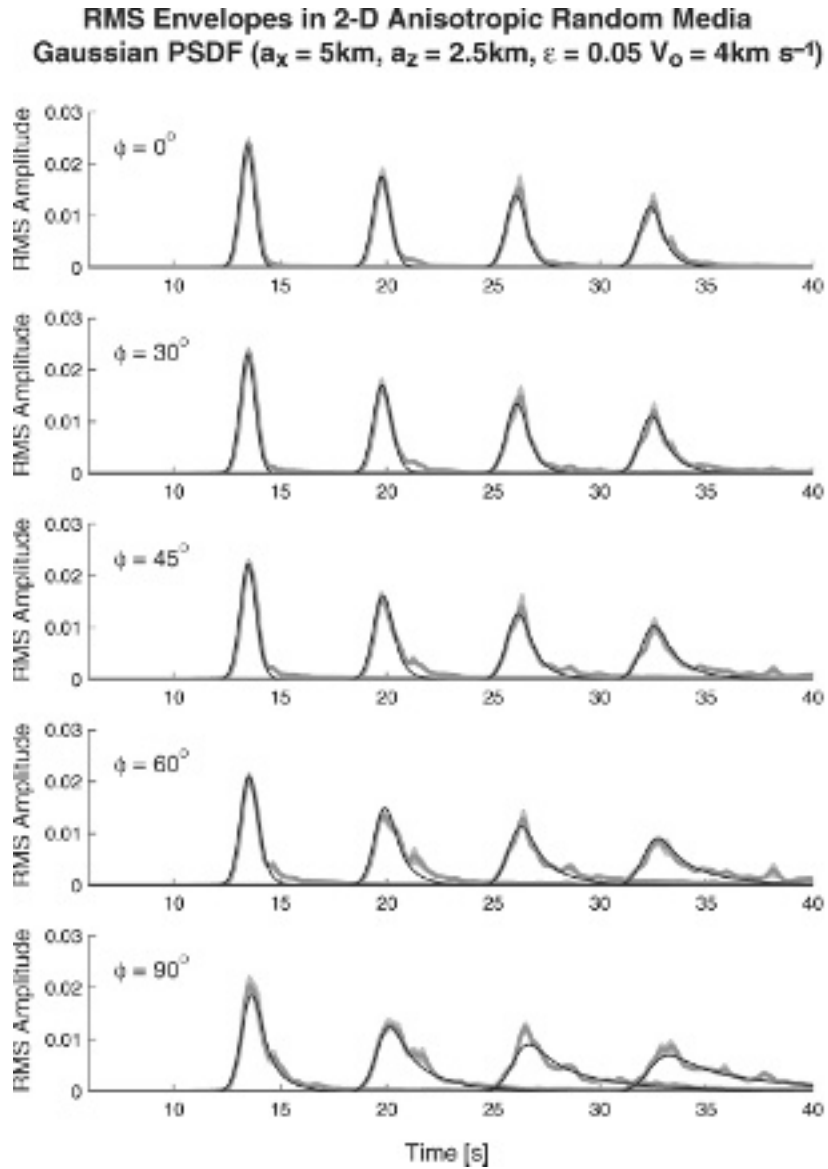
**Figure 6.** Theoretical mean-square (MS) envelopes in 2-D random media calculated by using the Markov approximation. Square amplitude is scaled by using the characteristic time  $t_M$  and geometrical spreading factor. Reduced time is scaled by  $t_M$ . The value of  $p$  depends on the type of random media. For the Gaussian ACF,  $p = 2$ . For the von Kármán-type ACF, the value of  $p$  is a function of the order  $\kappa$  (see Table 1).

appears again as coda-wave energy. Hence, the value of eq. (36) would be one of the appropriate parameters for judging the applicability of the envelope synthesis using the small-angle scattering approximation. When  $r_0 g_L \ll 1$ , we may consider that the effects of the large-angle scattering are small enough within travel distance  $r_0$ . On the value of the cut-off angle  $\theta_c$ , we have no consensus at present. Recently, Kawahara (2002) theoretically predicted  $\theta_c \approx 65^\circ$  by using the Kramers-Krönig relation (e.g. Aki & Richards 1980). On the other hand, this study uses  $\theta_c = 30^\circ$  referring to Sato (1982) since the case of  $\theta_c = 30^\circ$  predicting larger  $g_L$  than  $\theta_c = 65^\circ$  is preferable for judging the condition  $r_0 g_L \ll 1$ . It should be noted that the condition  $r_0 g_L \ll 1$  would be satisfied when  $ka_x \gg 1$  and  $ka_z \gg 1$  for the case of the Gaussian ACF since the order of  $\bar{k}$  is the same as  $ka_x$  and  $ka_z$ .

Another limitation of the Markov envelope would come from the approximation used in eq. (16); we suppose the existence of an intermediate scale  $\Delta r$ , which is larger than the correlation distance in  $r$ -axis, but smaller than the scale of the wavefield variation in the  $r$ -axis (see p. 243 and 246, Sato & Fehler 1998). It is difficult to discuss the existence of  $\Delta r$  quantitatively. However, we expect that the existence of the intermediate scale  $\Delta r$  would be easily assured for the wave propagation along the  $z$ -axis when  $a_x > a_z$  since the correlation distance  $a_z$  is small. On the other hand, when the anisotropy is strong  $a_x \gg a_z$ , the existence of  $\Delta r$  would not always be assured for the wave propagation along the  $x$ -axis.

#### 4.2 Comparisons between Markov envelopes and FD envelopes

For the FD envelopes plotted in Fig. 4, the condition (eq. 36) is satisfied; the values of  $r_0 g_L$  with  $\theta_c = 30$  degree are estimated as the order of  $10^{-7}$ ,  $10^{-5}$ ,  $10^{-3}$ ,  $10^{-2}$  and  $10^{-2}$  for the angles  $\phi = 0, 30, 45, 60$  and  $90$  degree, respectively. In those cases, the Markov envelopes are expected to be synthesized properly. In Fig. 7, the Markov envelopes are plotted by solid black curves, where the Markov envelopes in different incident angles are calculated with the same source radiation energy. The Markov envelopes are convoluted with the wandering effect, in other words,  $\Gamma_2$  is used instead of  ${}_0\Gamma_2$ . Additionally, the temporal change in the power of a 2 Hz Ricker wavelet is convoluted with the MS envelopes since the Ricker wavelet is used as the source time function in the FD simulations. The Markov envelopes (solid curves) are in good agreement with the FD envelopes (grey curves) for all the travel distances for all the incident angles. This supports the reliability of the envelope synthesis using the Markov approximation. There are some slight discrepancies. As example, for the envelopes at the distance  $r = 75$  km with the incident angle  $\phi = 60$  degree, the later phase excitation of the FD envelope at the time around 22 s cannot be explained by the Markov envelope. Such larger later-phase excitation in the FD envelopes would be caused by large-angle scattering, which is neglected in the formulation of the Markov envelope.



**Figure 7.** The rms envelopes for a 2 Hz Ricker wavelet source in 2-D random media characterized by the Gaussian ACF with  $\varepsilon = 0.05$ ,  $a_x = 5$  km,  $a_z = 2.5$  km and  $V_0 = 4$  km s<sup>-1</sup>. The envelopes at travel distance 50, 75, 100 and 125 km are plotted. Numerals at upper left side in each bin show incident angle measured from the  $z$ -axis. Solid-black curves show rms envelopes based on the Markov approximation (Markov envelopes) Bold-grey curves and grey areas around the curves show rms envelopes by FD simulations (FD envelopes) and their standard deviations, respectively.

## 5 DISCUSSION

### 5.1 Dependence of wave propagation on the propagation direction

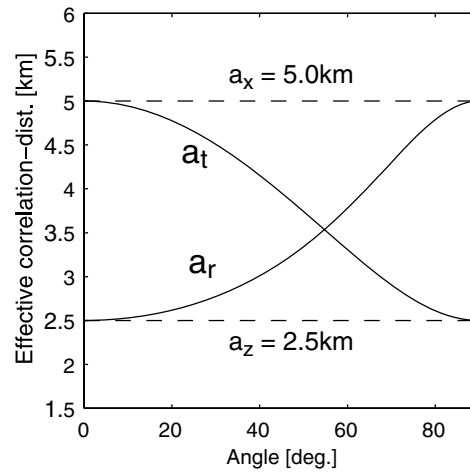
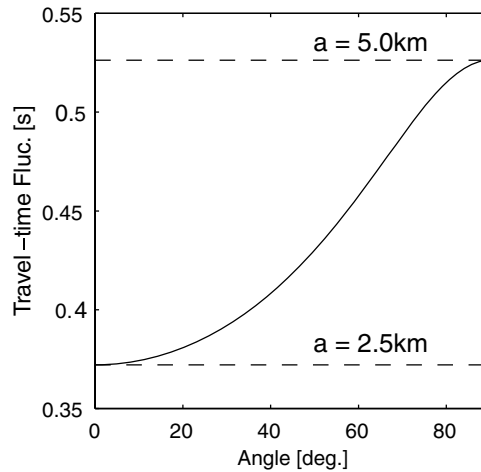
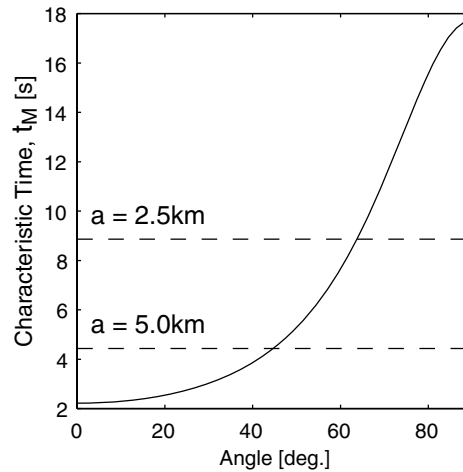
The characteristic time  $t_M$  (eq. 31), which can be considered as the envelope duration or a measure for distortion of envelopes, is almost the same as that of isotropic random media (see eq. 16 in Sato *et al.* 2004). Incident-angle dependence is additionally included in eq. (31) through the effective correlation distances,  $a_t$  and  $a_r$ . Fig. 8(a) shows the effective correlation distances (eq. 24) against the incident angle  $\phi$ . As the incident angle changes from  $\phi = 0$  to  $\phi = \pi/2$ ,  $a_t$  changes from  $a_x$  to  $a_z$  and  $a_r$  changes from  $a_z$  to  $a_x$ . Fig. 8(b) shows the traveltime fluctuation at a travel distance  $r = 200$  km for the Gaussian ACF with  $\varepsilon = 0.05$ ,  $a_x = 5$  km,  $a_z = 2.5$  km and  $V_0 = 4$  km s<sup>-1</sup>. This is calculated on the basis of the geometrical optic approximation (p. 234, Sato & Fehler 1998) as

$$\sqrt{\left\langle \left( \int_0^r \frac{dr'}{V(\mathbf{x})} - \frac{r}{V_0} \right)^2 \right\rangle} = \frac{\sqrt{r} A(0, \phi)}{V_0}. \quad (37)$$

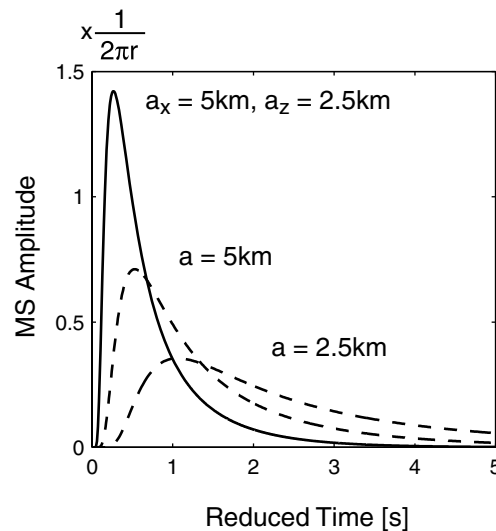
It should be noted that the geometrical optic approximation does not include the finite-wavelength effect on the traveltime fluctuation. The traveltime fluctuation increases with increasing the incident angle. When wave propagates along the  $x$ -axis ( $\phi = \pi/2$ ), the traveltime fluctuation

**Gaussian PSDF**

$a_x = 5\text{km}$ ,  $a_z = 2.5\text{km}$ ,  $\varepsilon = 0.05$ ,  
 $V_0 = 4\text{km s}^{-1}$ ,  $r = 200\text{km}$

**(a)****(b)****(c)**

**Figure 8.** Incident angle dependence of (a) effective correlation distances  $a_t$  and  $a_r$ , (b) traveltime fluctuation predicted by the geometrical optics (eq. 37) and (c) characteristic time (eq. 31). Those values are calculated at travel distance of 200 km in the random media characterized by the Gaussian ACF with  $\varepsilon = 0.05$ ,  $a_x = 5\text{ km}$ ,  $a_z = 2.5\text{ km}$  and  $V_0 = 4\text{ km s}^{-1}$ . Broken lines show the values in the case of isotropic random media with  $a_x = a_z = 5\text{ km}$  and  $2.5\text{ km}$ .



**Figure 9.** Squared-amplitude Markov envelopes at travel distance of 200 km for vertical (along the  $z$ -axis) wave propagation through 2-D random media characterized by the Gaussian ACF with  $\varepsilon = 0.05$ . The envelope in anisotropic media with  $a_x = 5$  and  $a_z = 2.5$  km is plotted by a solid curve. The envelopes in isotropic media with  $a_x = a_z = 5$  and 2.5 km by dashed curves.

takes the maximum value, which is the same as that calculated from the isotropic random media of a correlation distance  $a_x$ . The traveltime fluctuation at  $\phi = 0$  is the same as that from the isotropic random media of a correlation distance  $a_z$ . This behaviour is easily predicted since the value of  $A(0, \phi)$  is proportional to  $a_r$  (see eq. 23). Fig. 8(c) shows the characteristic time (eq. 31) against the incident angle. The characteristic time also increases with increasing the incident angle. However, the behaviour is not so simple as the traveltime fluctuation predicted by the geometrical optic approximation; neither of the minimum value of  $t_M$  nor the maximum value corresponds to that predicted from the isotropic random media. It is because that  $t_M$  is proportional to a combination of the effective correlation distances; the combination is given by  $a_r/a_i^2$  for the case of the Gaussian ACF. This makes the variation range of  $t_M$  wider than the range calculated from the isotropic random media of correlation distance  $a_x$  and  $a_z$ .

## 5.2 Envelopes in anisotropic lithosphere

Fig. 9 shows squared-amplitude Markov envelopes in anisotropic random media and isotropic random media for the case of vertical wave incidence ( $\phi = 0$  in Fig. 8). The envelope of the anisotropic random media has larger maximum amplitude and shorter duration than those of the isotropic random media with a correlation distance  $a_x$  or  $a_z$ . This discrepancy between the anisotropic and isotropic cases causes serious problems when estimating the random inhomogeneity from observed seismogram envelopes. We shall consider a case as follows. One analyses wave envelopes of the vertical wave incidence, or of deep events, to estimate the random inhomogeneity assuming isotropic random media, even though those envelopes actually propagate through anisotropic random media. The estimated correlation distance of isotropic random media would be larger than  $a_x$  and  $a_z$ . In other words, one would underestimate the spectral amplitude of the inhomogeneity. Analysing seismogram envelopes of the intermediate-depth events on the basis of the Markov envelopes of isotropic random media, Saito *et al.* (2005) estimated the PSDF of the inhomogeneous lithosphere in northeastern Honshu, Japan. They discussed that the estimated PSDF would not give enough coda excitation observed in past studies. The anisotropic inhomogeneity in the lithosphere may be one of the keys to solve this discrepancy. Although this study is limited to 2-D case, the formulation for the 3-D case is necessary for more rigorous and quantitative discussions of the anisotropic lithosphere.

## 6 CONCLUSIONS

The present study formulates a method of envelope synthesis using the Markov approximation in 2-D weakly anisotropic random media characterized by the Gaussian and von Kármán-type ACFs. The reliability of the formulation is examined by FD numerical simulations; wave propagation of a 2 Hz Ricker wavelet source through the random media characterized by the Gaussian ACF with  $\varepsilon = 0.05$ ,  $a_x = 5$  km,  $a_z = 2.5$  km and  $V_0 = 4$  km s<sup>-1</sup> is numerically solved in the travel distance ranging from 50 to 125 km. The envelopes of the Markov approximation are in good agreement with the envelopes obtained from wave traces of the FD simulations. This supports the reliability of the envelope synthesis using the Markov approximation. As in the case of isotropic random media, the Markov envelopes are scaled by using the characteristic time. The characteristic time is represented as a function of the propagation direction, the travel distance, the wave frequency and the parameters of random media. It predicts that envelopes increase in duration and decrease in maximum amplitude more rapidly in the horizontal propagation than in the vertical when the media are characterized by long horizontal and short vertical correlation distances.

Especially, in the case of the vertical wave propagation, the envelopes of the anisotropic random media have shorter duration and larger maximum amplitude than those of the isotropic random media. This implies that the intensity of the inhomogeneity is underestimated when one consider anisotropic lithosphere as isotropic random media.

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**APPENDIX A: DERIVATION OF EQ. (16)**

In this appendix, we derive the eq. (16) following the derivations of Lee & Jokipii (1975a) and Sato & Fehler (1998). From the parabolic eq. (12), we can write the wavefield at  $(r_{\perp j}, r)$  by using the wavefield at  $(r_{\perp j}, r - \Delta r)$ ,

$$\begin{aligned} U(r_{\perp j}, r, \omega_j) &\approx U(r_{\perp j}, r - \Delta r, \omega_j) + \frac{i}{2k_j} \int_{r-\Delta r}^r \frac{1}{r'^2} \frac{\partial^2}{\partial \phi^2} U(r_{\perp j}, r', \omega_j) dr' - ik_j \int_{r-\Delta r}^r \xi(r_{\perp j}, r') U(r_{\perp j}, r', \omega_j) dr' \\ &\approx U_j(r - \Delta r) + \frac{i}{2k_j} \frac{1}{r^2} \frac{\partial^2}{\partial \phi^2} U_j(r - \Delta r) \Delta r - ik_j U_j(r - \Delta r) \int_{r-\Delta r}^r \xi_j(r') dr', \end{aligned} \quad (\text{A1})$$

where  $U_j(r - \Delta r)$  and  $\xi_j(r')$  mean  $U(r_{\perp j}, r - \Delta r)$  and  $\xi(r_{\perp j}, r')$ , respectively. In the approximation used in eq. (A1), we suppose the existence of an intermediate scale  $\Delta r$ , which is larger than the scale of inhomogeneity along the  $r$ -axis, or the correlation distance, but smaller than the scale of variation of  $U$  along the  $r$ -axis. It should be noted that the second and third terms in the last equation are small compared to the first term. Neglecting smaller terms, we calculate an ensemble-averaged value as follows,

$$\begin{aligned} \langle \xi_1(r) U_1(r) U_2^*(r) \rangle &\approx \langle ik_2 \xi_1(r) U_1(r - \Delta r) U_2^*(r - \Delta r) \int_{r-\Delta r}^r \xi_2(r'') dr'' \rangle - \langle ik_1 \xi_1(r) U_1(r - \Delta r) U_2^*(r - \Delta r) \int_{r-\Delta r}^r \xi_1(r') dr' \rangle \\ &\quad - \left\langle \frac{i \xi_1(r) \Delta r}{2k_2 r^2} U_1(r - \Delta r) \frac{\partial^2}{\partial \phi^2} U_2^*(r - \Delta r) \right\rangle + \left\langle \frac{i \xi_1(r) \Delta r}{2k_1 r^2} U_2^*(r - \Delta r) \frac{\partial^2}{\partial \phi_1^2} U_1(r - \Delta r) \right\rangle \\ &\quad + \langle \xi_1(r) U_1(r - \Delta r) U_2(r - \Delta r) \rangle \\ &\approx ik_2 \langle U_1(r - \Delta r) U_2^*(r - \Delta r) \rangle \left\langle \int_{r-\Delta r}^r \xi_1(r) \xi_2(r'') dr'' \right\rangle - ik_1 \langle U_1(r - \Delta r) U_2^*(r - \Delta r) \rangle \left\langle \int_{r-\Delta r}^r \xi_1(r) \xi_1(r') dr' \right\rangle \\ &= ik_2 \langle U_1(r - \Delta r) U_2^*(r - \Delta r) \rangle \int_0^{\Delta r} R(r_{\perp 1} - r_{\perp 2}, r') dr' - ik_1 \langle U_1(r - \Delta r) U_2^*(r - \Delta r) \rangle \int_0^{\Delta r} R(0, r') dr' \\ &\approx \frac{1}{2} ik_2 \langle U_1(r - \Delta r) U_2^*(r - \Delta r) \rangle \int_{-\infty}^{\infty} R(r_{\perp 1} - r_{\perp 2}, r') dr' - \frac{1}{2} ik_1 \langle U_1(r - \Delta r) U_2^*(r - \Delta r) \rangle \int_{-\infty}^{\infty} R(0, r') dr' \\ &\approx \frac{i}{2} \{k_2 A(r_{\perp 1} - r_{\perp 2}, \phi) - k_1 A(0, \phi)\} \Gamma_2. \end{aligned} \quad (\text{A2})$$

At the derivation of the third equation, the causality deduced from the forward scattering approximation is used; there is no contribution of the inhomogeneity at  $r$  to the wavefield at  $r - \Delta r$ . At the derivation of the fifth equation, the integration at the fourth equation is approximated by the infinite integration since we supposed  $\Delta r$  is larger than the correlation distance. Also, we supposed  $R(r_{\perp k} - r_{\perp m}, -r'') \approx R(r_{\perp k} - r_{\perp m}, r'')$  in weakly anisotropic random media considering that the contribution of  $R(r_{\perp k} - r_{\perp m}, r'')$  is larger as  $|r_{\perp k} - r_{\perp m}|$  become smaller. At the derivation of the last equation, we suppose  $\langle U_1(r - \Delta r) U_2^*(r - \Delta r) \rangle \approx \langle U_1(r) U_2^*(r) \rangle$  considering the existence of small  $\Delta r$  which is smaller than the scale of variation of  $U$  along the  $r$ -axis. By using the eq. (A2), we obtain the following relation,

$$\langle (k_1 \xi_1 - k_2 \xi_2) U_1 U_2^* \rangle \approx -\frac{i}{2} [(k_1^2 + k_2^2) A(0, \phi) - 2k_1 k_2 A(r_{\perp 1} - r_{\perp 2}, \phi)] \Gamma_2. \quad (\text{A3})$$