# ORIGINAL PAPER

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# **Redox state of eclogites and peridotites from sub-cratonic upper mantle and a connection with diamond genesis**

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Abstract Pressure–Temperature-  $f_{O_2}$  conditions and fluid compositions estimated for mineral parageneses from inclusions in diamonds, diamond-bearing and diamondfree xenoliths using a garnet–clinopyroxene–silica oxygen barometer data indicate that the upper mantle is zoned, with a relatively oxidized lithosphere and a reduced asthenosphere. Calculations in the C–O–H system indicate that eclogite inclusions within diamonds and xenoliths have formed mainly in equilibrium with water-rich fluids.

**Keywords** Oxygen fugacity · Barometry · Mantle · Diamond genesis

### Introduction

The oxidation state of the mantle has been a source of recent controversy. Thermodynamic calculations based on the olivine-orthopyroxene-spinel (O'Neill and Wall 1987) and olivine-orthopyroxene-garnet (Luth et al. 1990) equilibria along with measurement of  $Fe^{+3}/Fe^{+2}$ in basaltic glasses (Christie et al. 1986) support a relatively oxidized mantle, characterized by values of oxygen fugacity between fayalite-magnetite-quartz oxygen (FMQ) buffer and wustite-magnetite (WM) buffer. In contrast, the intrinsic oxygen-fugacity measurement (IOF) indicates more reduced values of  $f_{\Omega_2}$ , near iron-wustite (IW) buffer (Ulmer et al. 1987). It has been suggested (Haggerty and Tompkins 1983) that the upper mantle is zoned with the depleted lithosphere being more reduced than the deeper, fertile portions of the asthenosphere. Taylor and Green (1989) and Ballhaus and Frost (1994) suggested an opposite view that the mantle consists of a relatively oxidized lithosphere

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S. K. Simakov Geological Department, St Petersburg University, 7/9 Universitetskaya Nab., St. Petersburg 199034, Russia E-mail: simakov@ap1250.spb.edu and a reduced asthenosphere. Comparatively little attention has been paid to the role of oxygen fugacity in the diamond formation, which determines whether carbon exists either in the elemental form or remains complexed with oxygen in carbonates or in a gaseous species. Diamonds have formed in association with metal-silicate melts, traces of which have been noted in natural diamonds (Bulanova and Zayakina 1991). Such melts form at oxygen fugacities that corresponded to the IW buffer. From previous calculations based on the olivine-orthopyroxene-garnet equilibria, Simakov (1998) argued that diamonds with peridotite inclusions were formed under the reducing conditions corresponded to IW buffer. On the other hand, the presence of carbonate inclusions and solid CO<sub>2</sub> in diamonds would indicate a more oxidizing environment (Bulanova and Pavlova 1987; Schrauder and Navon 1993), corresponding to the QFM buffer. Schrauder and Navon (1993) suggested that the carbon for the formation of the diamond might have come from a partial reduction of CO<sub>2</sub> itself. From the composition of phases in spinel peridotites it follows that  $f_{O_2}$  may vary by five log units in the lithosphere under the ancient cratons (Ballhaus 1993). The oxidation state of garnet lherzolites and eclogites has not been as extensively studied as that of spinel lherzolites.

## Calibration of garnet-clinopyroxene-silica oxygen barometer

The oxidation state of the mantle can be estimated by redox reactions involving iron-bearing minerals. Commonly, the phase containing ferric iron is the spinel, but it is possible to calibrate a garnet–clinopyroxene–silica barometer based on the solubility of Fe<sup>+3</sup> in garnets. Zhang and Saxena (1991) proposed to use the reactions of the garnet–clinopyroxene–quartz equilibrium for skarns with coexisting andradite and hedenbergite. Another reaction was proposed by Simakov (1993) for mantle eclogites:

$$\begin{array}{ccc} Ca_3Fe_2Si_3O_{12} & +Fe_2Si_2O_6 & +4SiO_2 & \Rightarrow 6CaFeSi_2O_6 + O_2 \\ Andr & Fs & Qu(Cs) & Hed \end{array}$$
(

For this reaction, the oxygen-fugacity value is expressed as:

$$\log f_{\rm O_2} = -\Delta G_T^P / 2.303 RT - \alpha_{\rm Hed}^6 / (\alpha_{\rm Andr}^2 \alpha_{\rm SiO_2}^4 \alpha_{\rm Fs}), \qquad (1)$$

 $\Delta G^{P}_{T}$  was calculated from the thermochemical data compiled by Kuskov and Fabrichnaya (1990) and Moecher et al. (1988) (Table 1). Clinopyroxene-activity models are calculated in accordance with the formulas:

$$\begin{aligned} \alpha_{\text{Hed}} &= \gamma_{\text{Hed}} \text{Ca}^{\text{M2}} \text{Fe}^{\text{M1}} \text{Si}^2, \end{aligned} (2) \\ \alpha_{\text{Fs}} &= \gamma_{\text{Fs}} \text{Fe}^{\text{M2}} \text{Fe}^{\text{M1}} \text{Si}^2, \end{aligned} (3)$$

where:

 $\mathbf{2}$ 

$$\begin{split} M1 &= Fe^{M1} + Mg^{M1} + Ti + Cr + Fe^{+3} + Al^{M1}, \\ M2 &= Fe^{M2} + Mg^{M2} + Ca + Mn + Na + K, \\ Al^{M1} &= X_{Al} - Al^{IV}, \\ Al^{IV} &= 2 - X_{Si}, \\ Si &= X_{Si}/2. \end{split}$$

this clinopyroxene model was tested by comparing it with the experimental data of omphacitic clinopyroxenes (Simakov 2005) and hedenbergite–jadeite solutions (Fig. 1) (Perchuk and Aranovich 1991). The andradite activity ( $\alpha_{(Andr)}$ ) is calculated using the multi-component solid-solution model in garnet (Simakov 1998):

$$\log \alpha_{\text{Andr}} = (G^{e}_{\text{Andr}_{\text{I}}} + 3G^{e}_{\text{Andr}_{\text{I}}} + 2G^{e}_{\text{Andr}_{\text{II}}})/RT + \log X^{3}_{\text{Ca}}X^{2}_{\text{Fe}} + 3.$$
(4)

Non-ideal mixing on the dodecahedral sites  $(G_{Andr_{I}}^{e})$  is treated by an asymmetric sub-regular solution model with the formulation given by Moecher et al. (1988). Non-ideal mixing on the octahedral sites  $(G_{Andr_{I}}^{e})$  was treated with a symmetric regular solution model with the interaction parameters listed in Table 4. Reciprocal solid-solution effects in the garnets (Wood and Nicols 1978) were accounted for with the exchange reactions:

$$\begin{array}{ccc} Mg_3Al_2Si_3O_{12} & +Ca_3Fe_2Si_3O_{12} & =Mg_3Fe_2Si_3O_{12} & +Ca_3Al_2Si_3O_{12} \\ Pyr & Andr & Gross \end{array}$$
(II)

$$\begin{array}{ccc} Ca_{3}Fe_{2}Si_{3}O_{12} & +Fe_{3}Al_{2}Si_{3}O_{12} & = Fe_{3}Fe_{2}Si_{3}O_{12} & +Ca_{3}Al_{2}Si_{3}O_{12} \\ Andr & Alm & Ski & Gross \end{array} \tag{III}$$

$$\begin{array}{cccc} Mg_3Al_2Si_3O_{12} & +Ca_3Cr_2Si_3O_{12} & = Mg_3Cr_2Si_3O_{12} & +Ca_3Al_2Si_3O_{12} \\ Pyr & Uvar & Knor & Gross \end{array} \tag{$IV$}$$

$$\begin{array}{ccc} Ca_3Cr_2Si_3O_{12} & +Fe_3Al_2Si_3O_{12} & =Fe_3Cr_2Si_3O_{12} & +Ca_3Al_2Si_3O_{12} \\ Uvar & Alm & Fe - Knor & Gross \end{array} \tag{$V$}$$

The scheme of element distribution from M1 and M2 position is taken from Wood and Banno (1973). The activity coefficients of the clinopyroxene minerals was modeled by assuming the clinopyroxene as an asymmetrical solid-solution model. On the M1 and M2 positions, the parameters of interactions were taken from the papers listed in Table. 2 and 3. The precision of

The Gibbs free energy values of these reactions are listed in Table 4. This model can be applied to the peridotite parageneses by calculating the activity of silica in accordance with the thermodynamic model of olivine–orthopyroxene equilibrium in the upper mantle (O'Neill and Wall 1987) by the reaction:

Table 1 Gibbs free energy, entropy, molar volume and entropy coefficients of the phases involved in the thermodynamic calculations

Mineral	$\Delta G^{\circ}_{298}$ (kJ/mol)	$S^{\circ}_{298}$ (J/mol K)	V°298 (cc/mol)	А	В	С	D	Е
Andradite Ferrosilite Hedenbergite Quartz ( $\alpha$ ) Oxygen (O <sub>2</sub> )	-5,413.2 -2,232.4 -2,677.7 -856.6 0.0	316.82 191.64 173.59 41.46 205.04	131.67 63.98 67.85 22.69	470.395 232.99 214.116 73.448 20.15	46.903 30.402 48.359 0.782 28.078	63.743 32.318 24.698 15.376 2.065	-2,765.6 -1372.6 -1,262 436.1 -1.772	4.172

The thermochemical data were taken from: Moecher et al. (1988). The  $\Delta G_T^P$  of  $\alpha$ -Quartz  $\Rightarrow$  Coesite transition is calculated in accordance with the equation given in Kuskov and Fabrichnaya (1990)

I)

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**Table 2** Data of interaction parameters on M2 clinopyroxene position (in J/mole,  $T^{\circ}$  K, P—kbar)

(M2)	Ca–Mg	Mg–Ca	Ca–Fe	Fe–Ca	Mg–Fe	Ca–Na	Na–Ca	Mg(Fe)–Na
a b c d	31,216–6.1 <i>P</i>	25,484+8.12 P	20,697–2.35 P	16,940 + 5.9 <i>P</i>	-2,172	31,120	16,707	-24,000
The da	ta were taken from	1:						

Lindsley (1981)

<sup>b</sup> Lindsley et al. (1981)

<sup>c</sup> Mukhopadhyay (1991)

<sup>d</sup> Simakov (2005)

**Table 3** The gata of interaction parameters on  $M_1$  clinopyroxene position (in J/mole,  $T^{\circ}$  K, P—kbar)

(M1)	Mg–Fe	Fe–Mg	Mg(Fe)–Cr	Al–Cr	Mg(Fe)–Al	Al-Mg(Fe)
c d	3,978				42,000–10.63 T	-1,000
e			18,481	20,000	,	1,000

<sup>e</sup> Girnis and Grutter (2003), another reference the same as in Table 2

$$\begin{array}{rl} Mg_2SiO_4 & +SiO_2 & \Rightarrow 2MgSiO_3 \\ Fo & Qu(Cs) & Enst \end{array} \tag{$V$}$$

This model can be also applied to the garnet-orthopyroxene-clinopyroxene parageneses. We can calculate the composition of fictive olivine in accordance with the thermodynamic model of olivine-orthopyroxene equilibrium in the upper mantle (Seckendorff and O'Neill 1993) by the reaction:

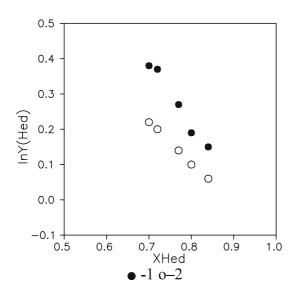
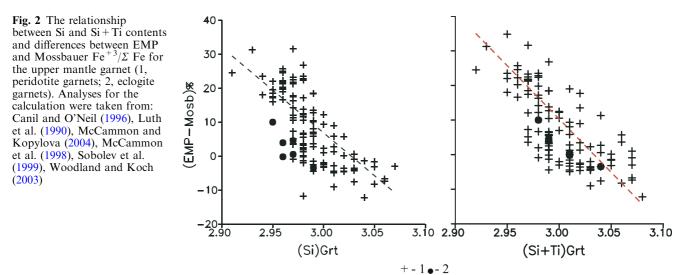


Fig. 1 A comparison of the experimental data of Hed–Jd solution (Perchuk and Aranovich 1991) (1) with calculated ones (2) by our Cpx model

Careful studies of eclogites by Sobolev et al. (1999a, 2000) show the presence of coesite inclusions in eclogite diamonds from kimberlites and lamproites. This implies that we could apply the reaction (I) to mantle eclogite parageneses by taking the activity of SiO<sub>2</sub> to be 1 (Nakamura and Banno 1997). This model can be also applied to the Si-undersaturated eclogite parageneses. We can use the fictive minal of  $SiO_2$  with activity equal to 1. Because ferric iron is so low in garnet, the garnetclinopyroxene-silica-oxygen barometer requires a careful measurement of Fe<sup>+3</sup>/ $\Sigma$  Fe in garnet and clinopyroxene. There are only two analytical methods commonly used to determine  $Fe^{+3}/\Sigma$  Fe directly, Mossbauer spectroscopy, and wet chemistry. It has been common practice to use the crystal chemistry of silicate phases analyzed by the electron microprobe (EMP) to calculate the  $Fe^{+3}$  of a mineral. Several attempts have been made to compare and correlate the measured and calculated  $Fe^{+3}/\Sigma$  Fe contents in silicates from upper mantle xenoliths. Luth et al. (1990) stated that for peridotite minerals (garnet, olivine, clino-, and orthopyroxene) the values of  $Fe^{+3}/\Sigma$  Fe measured by the Mossbauer technique were significantly different from those calculated from the EMP analyses. Ballhaus et al. (1991) have demonstrated that stoichiometric  $Fe^{+3}$  in spinel agrees well with the ones measured by Mossbauer spectroscopy. Canil and O'Neill (1996) showed that the level of imprecision in Fe<sup>+3</sup> determined from EMP analyses is related to the total Fe content, and increases in the order: spinel < garnet < clinopyroxene. McCammon et al. (1998) and Sobolev et al. (1999b) used the Mossbauer milliprobe to determine  $Fe^{+3}/\Sigma$  Fe in eclogite garnet and clinopyroxene from George Creek diamonds and Udachnaya kimberlite. Woodland and Peltonen (1999), Woodland and Koch (2003), and McCammon and Kopylova (2004) used the Mossbauer milliprobe to determine  $Fe^{+3}/\Sigma$  Fe in peridotite garnet



**Table 4** The data of the garnet-interaction parameters on the octahedral site ( $G^e_{\{Andr\{II\}}$ ) and between dodecahedral and octahedral sites expressed in the Gibbs free energy of the exchange reactions (in J/mole,  $T^\circ K$ , P—kbar)

$G^e_{\mathrm{Andr}_{\mathrm{II}}}$			$G^e_{ m Andr}$ I - II			
W <sub>Al-Cr</sub>	W <sub>A1-Fe</sub>	W <sub>Cr-Fe</sub>	II	III	IV	V
2,508	3,700	1,267	25,080	40,200–7 <i>P</i>	122,181–49.7 <i>T</i>	106,903–43.45 <i>T</i>

The data were taken from: Aranovich (1991), Luth et al. (1990), Woodland and O'Neill (1993)

from Southern Africa, Finnish and Canadian kimberlites. Sobolev et al. (1999b) stated that the accuracy of the ferric iron estimate depends on the accuracy of the silica determination. A comparison of the Fe<sup>+3</sup>/ $\Sigma$  Fe values calculated by Schumacher's (1991) model with those determined by Mossbauer techniques for mantle garnets shows the correlation between the error in garnet Fe<sup>+3</sup>/ $\Sigma$  Fe determination and Si and Si+Ti contents (Fig. 2). The lowest error lies in the range of  $3.03 \pm 0.02$  f.u. (Si+Ti)<sub>Grt</sub> (Fig. 2).

Uncertainties of the calculated oxygen fugacity result from uncertainties in the activity model of the minerals involved and the calculated pressure and temperature. At temperatures more than 1,100°C, oxygen fugacity on 70% and more depends upon the mineral activities and only on 30% or less upon the P-T parameters. Otherwise, at low temperatures the  $f_{\{O2\}}$  on 70% depends upon P-T parameters and only on 30% or less upon the mineral activities. The andradite minal is the 45-55% part of the total mineral activity, the hedenbergite one is 30-35%, the ferrosilite one is 10-15% and the silica one is only 1–3% (for peridotite parageneses). The andradite activity coefficient is 10–15% part of the total mineral activity, the hedenbergite one is 5-7% and ferrosilite and silica ones are by 1-3%. The 4 kJ/mole error in the free energy of formation of andradite yields an uncertainty of about  $\pm 0.1 \log f_{O_2}$ . The reciprocal terms in the  $f_{O_2}$ calculation are typically in the order of 0.3 log units while the uncertainty in  $Fe^{+3}$  content of garnet leads to 0.5 log units in  $f_{O_2}$  uncertainty. The uncertainty in Fe<sup>+3</sup>

content of clinopyroxene leads to 0.5 log units in  $f_{O_2}$  uncertainty too. Assuming a normal distribution of these errors implies a total uncertainty of  $\pm 0.8$  log units

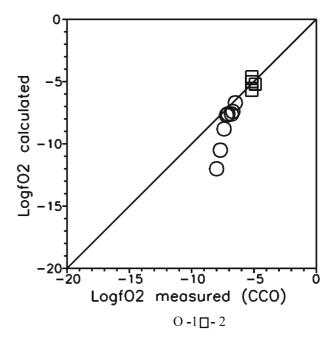


Fig. 3 The difference between the  $f_{O_2}$  values calculated by the equilibrium (I) and measured for synthesized nearly CCO buffer eclogites [1, Yaxley (1999)] and carbonatized peridotites [2, Ryabchikov et al. (1993)]

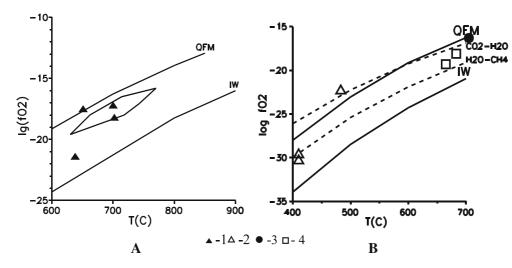
in  $f_{O_2}$  for our sensor. The accuracy of garnet-clinopyroxene-silica oxygen barometer was tested on the experimental data of garnet-clinopyroxene assemblages synthesized in the carbonate-bearing eclogite system at 30-35 kb and 700-900°C and on the garnet-twopyroxene-olivine assemblages synthesized in the peridotite system at 50 kbars and 1,400-1,550°C with the oxygen fugacity controlled by the CCO buffer (Ryabchikov et al. 1993; Yaxley 1999) (Fig. 3). Fe<sup>3+</sup> were calculated for the garnets with  $(Si + Ti)_{Grt}$ of  $3.0 \pm 0.02$  f.u. from the EMP analyses. As can be seen from Fig. 3, oxygen fugacities calculated from equilibrium (I) are in fair agreement with oxygen fugacities measured with the CCO sensor. The only data which lie outside the quoted uncertainties could be from those experiments in which garnet and clinopyroxene were not in exchange equilibrium. Accuracy of the barometer also was tested on the garnet-spinel lherzolites by comparing the oxygen fugacity calculated from the garnet-clinopyroxene-silica equilibrium with those from the olivineorthopyroxene-spinel equilibrium. These calculations

are correct when the minerals of the paragenesis are in mutual equilibrium. Yakutian kimberlitic garnet-orthopyroxene-clinopyroxene-olivine-spinel lherzolite was chosen for the oxygen-fugacity calculations. The difference between garnet-orthopyroxene, spinel-olivine, clinopyroxene-orthopyroxene, two-pyroxene and garnet-olivine thermometers did not exceed 50°C (Table 5). Accuracy of the barometer was also tested on the garnet lherzolites with Mossbauer garnet and clinopyroxene data by the comparing the oxygen fugacity calculated from the equilibrium (I) with those from the olivineorthopyroxene-garnet sensor of Gudmudson and Wood (1995) with an accuracy of  $\pm 0.6 \log f_{O_2}$  units. The parageneses were chosen for oxygen- fugacity calculations by the procedure described above. The difference between garnet-orthopyroxene, clinopyroxene-orthopyroxene, two-pyroxene and garnet-olivine thermometers did not exceed 155°C. Calculated  $f_{O_2}$  values for South African, Canadian, and Yakutian parageneses coincide with the calculations for the same analyses done with using the Gudmundson and Wood (1995)

Table 5 Garnet-spinel-orthopyroxene-clinopyroxene-olivine paragenesis separated for  $P-T-f_{102}$  calculations

Sample	NG85	H84	NW79	B91	Ai93	T98	$\Delta(f_{O_2})_1$	$\Delta log(f_{O_2})_2$
Yakutian xen Uv-624	nolith 42.68	847	856	824	874	847	-2.7	-2.1

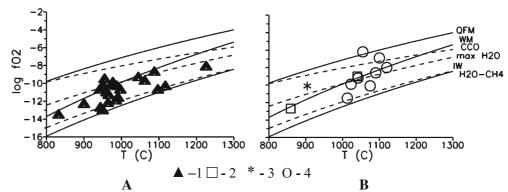
NG85, Grt–Opx barometer of Nickel and Green (1985) (in kbar); H84, Grt–Opx thermometer of Harley (1984); NW79, Grt–Ol thermometer of O'Neill and Wood (1979); Ai93, Grt–Cpx thermometer of Ai (1993); T98, Cpx–Opx thermometer of Taylor (1998) (in °C).  $\Delta log(f_{O_2})_1$ , calculated values by garnet–clinopyroxene–silica barometer;  $\Delta log(f_{O_2})_2$ , values obtained by spinel–orthopyroxene–olivine barometer of Ballhaus et al. (1991) with Taylor and Grenn's (1991) correction for high-chromium spinels at the same pressures and temperatures relative to the QFM buffer of Frost (1991). For  $f_{O_2}$  calculations *P*–*T* parameters obtained from garnet–orthopyroxene, thermobarometers of Harley (1984) and Nickel and Green (1985) were used. Analyses for the calculations were taken from Pokhilenko et al. (1991)



**Fig. 4** A comparison of the oxygen fugacities calculated using our sensor (1, Johnston and Essene 1982; Spear and Markussen 1997) with those, obtained by Lamb and Valley (1984) using the Mgt–Ilm sensor (*solid curve* between QFM and IW) for Andirondack complex (**a**), results of oxygen-fugacity calculations for Voltri Massif [2, Messiga and Scambelluri (1991)], Saulpian [3, Mottana

et al. (1968)] eclogites and Southern Karnataka (Indian) metamorphites [4, Hansen et al. (1984a); Janardan et al. (1982)] (b). For temperature estimation, the thermometer of Ai (1993) was used at P=7 kbar (*upper dashed line*—CCO buffer, *lower dashed line*—H<sub>2</sub>O–CH<sub>4</sub>—fence)

Fig. 5 The results of oxygenfugacity calculations for South African and Yakutian peridotite garnet-twopyroxene-olivine parageneses (see Appendix 1) with using Mossbauer data (1) (a) and for Yakutian (2, diamond bearing; 3, diamond-free); Namibian and Lesotho (4) peridotites with using EMP data (b) with using the garnet-clinopyroxene-silica oxygen barometer at 45 kbar



equilibrium (see Appendix 1). I conclude, therefore, that equilibrium (I), with free energy and the activity terms discussed above should, in general, yield the oxygen fugacity values to within  $\pm 0.8$  log units.

The oxygen barometer can also be applied to crustal eclogites using the model of Fe and Mg distribution on the M1 and M2 sites of clinopyroxene in accordance with the formulation of Dal Negro et al. (1982). The sensor was applied to crustal eclogitic and granulitic rocks from different metamorphic complexes. Andirondack complex in northern United States is one of the well-studied metamorphic complexes in the world (Johnston and Essene 1982; Spear and Markussen 1997). It was applied to olivine-bearing metagabbros and granulites from the Adirondack and the obtained  $f_{\Omega_2}$ values were compared with those calculated for the same rocks by the magnetite-ilmenite oxybarometer (Lamb and Valley 1984) (Fig. 4a). From another side, the results were compared to the compositions of fluid inclusions that were measured in the rocks. For the Voltrian Group eclogites, the calculated  $f_{O_2}$  values correspond to or are above the QFM buffer (Fig. 4b), for Southern Karnataka (India) granulites and charnokites, the calculated  $f_{O_2}$  values correspond to the QFM and WM buffers (Fig. 4b), in agreement with the water and water–carbon dioxide compositions of fluid inclusions in these rocks (Mottana et al. 1968; Messiga and Scambelluri 1991; Hansen et al. 1984b).

#### **Results and discussion**

Pressure–Temperature  $f_{O_2}$  parameters were calculated for inclusions within diamonds, diamond-bearing and diamond-free xenoliths from Australia, South Africa, and Yakutia using EMP and Mossbauer Fe<sup>+3</sup> analyses for garnet and clinopyroxene. For the calculations based on EMP analyses, only the garnet with (Si+-Ti)<sub>Grt</sub> of  $3.03 \pm 0.02$  f.u. was used (see Appendices 1, 2). For the eclogite garnet–clinopyroxene parageneses, the

Fig. 6 The results of oxygenfugacity calculations for Roberts Victor (a), Koidu and diamond-bearing Finnish (b), Mir (c), and Udachnaya (d) eclogites at 45 kbar (see Appendix 2). (1, diamond inclusion; 2, diamond-bearing eclogites; 3, xenoliths, and 4, diamond-bearing xenoliths of Udachnaya pipe, calculated with using Mossbauer data). solid lines: QFM quartsfayalite-magnetite, WM wustite-magnetite, IW ironwustite (Frost 1991), dashed lines: CCO buffer-correspond to upper limit of diamond stability in C-H-O system, max H<sub>2</sub>O—correspond to the reaction X, H<sub>2</sub>O-CH<sub>4</sub>—correspond to the lower boundary of water-rich  $(X_{\rm H_2O} > 0.5)$  fluids] (see Simakov 1998)

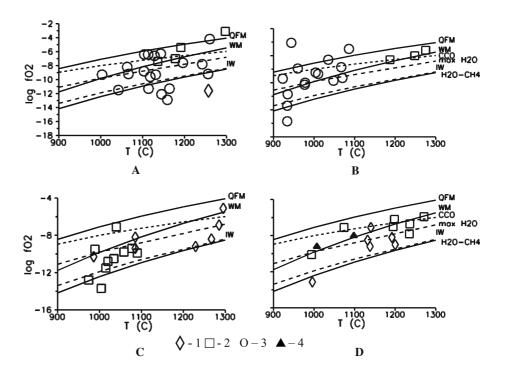
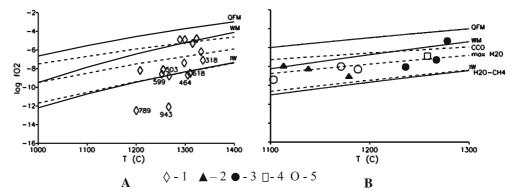


Fig. 7 The results of oxygenfugacity calculations for eclogite diamond inclusions from Premier (1) (a) George Creek (2, Mossbauer data), Argyle (3), Guinea (4) and Venezuela (5) (b) at 50 kbar (see Appendix 2). For Premier and Finsh diamonds, the data of nitrogen content of Deines's et al. (1989) in ppm is plotted



temperature was estimated using the garnet-clinopyroxene thermometer of Ellis and Green (1979). The peridotite garnet-two-pyroxene-olivine parageneses were chosen during the procedure described above.

Fluid compositions for eclogite and peridotite parageneses were estimated as a function of P, T and  $f_{O_2}$ with the assumption that they were free carbon-saturated. The fluid composition in equilibrium with diamond was calculated in accordance with the scheme described in Simakov (1998). In the C–H–O system, the upper limit of carbon stability in terms of oxygen fugacity corresponds to the reactions of the CCO buffer:

$$\begin{array}{ll} \mathrm{CO}_2 \rightarrow \mathrm{C} + \mathrm{O}_2, & (VIII) \\ \mathrm{CO} \rightarrow \mathrm{C} + 0.5\mathrm{O}_2. & (IX) \end{array}$$

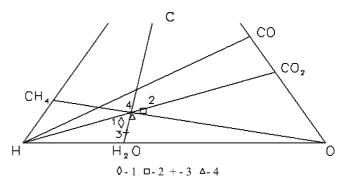
In the mantle, however, the CCO buffer is only of theoretical importance. In C–O–H system, diamond does not impose a lower  $f_{O_2}$  limit. One may be included to place a common lower limit near an  $f_{O_2}$  defined by the equilibrium:

$$CH_4 + O_2 \to C + 2H_2O, \tag{X}$$

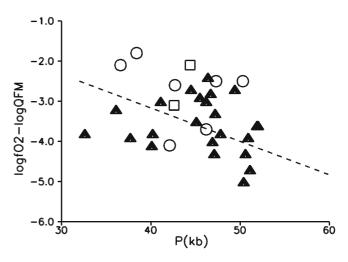
below, which the partial pressures of  $CH_4$  and  $H_2$  exceed the partial pressures of  $CO_2$  and CO and main part of the free carbon transfers to the methane.

Most of the peridotitic diamond-bearing and diamond-free xenoliths lie at oxygen fugacities below those of the CCO buffer (Fig. 5). Most of the eclogitic inclusions within diamonds also equilibrated at conditions below those of the CCO buffer, near WM (Figs. 6, 7). The results show that at these oxygen fugacities, the average calculated eclogite fluid compositions are waterrich [real apparent error of the Grt–Cpx–Qu(Cs) sensor  $(\pm 0.8 \log units in f_{O_2})$  could not have great effect on the average results] and close to the average composition of the gaseous inclusions in natural diamonds (see Fig. 8).

Calculated P-T-  $f_{O_2}$  results obtained for peridotite parageneses show that oxygen fugacity in the upper mantle mainly decreases with the depth (Fig. 9). It confirmed the conclusion of Taylor and Green (1989) that the upper mantle under the ancient cratons is zoned and consists of relatively oxidized lithosphere and reduced asthenosphere. Reduction and the fluid H<sub>2</sub>O/ CO<sub>2</sub> ratio increase with increasing depth (Fig. 5), which agrees with the experimental data in the carbonated peridotite-fluid system (Wyllie 1977) and with the thermochemical computation of oxidation state of the mantle (Saxena 1989).



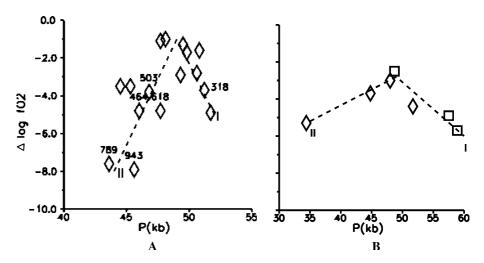
**Fig. 8** Average compositions of the calculated fluids for African parageneses: 1, eclogite inclusions from diamonds (average of 40 analyses); 2, eclogite xenoliths (average of 66 analyses); 3, peridotite xenoliths (average of 32 analyses); 4, fluid extracted from South African, Brazilian, and USA diamonds (average of 36 crystals) (Giardini and Melton 1975). Analyses for calculations were taken from Appendices 1 and 2



**Fig. 9** Log  $(f_{O_2})$  calc—log(QFM) dependencies for peridotites on the pressure (symbols—the same as on Fig. 5)

basis of CaTs barometer

(Simakov and Taylor 2000)

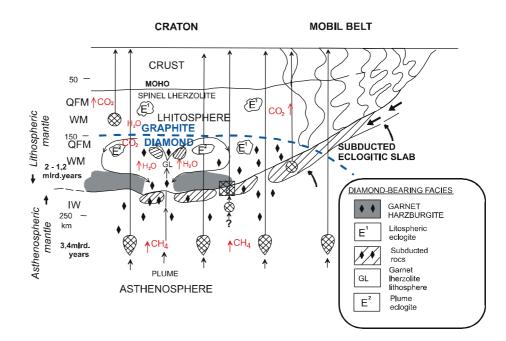


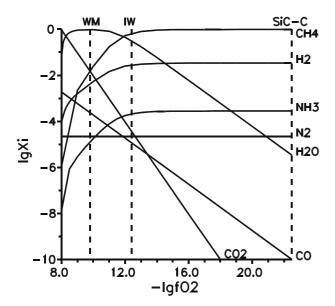
Calculated  $P-T-f_{O_2}$  results obtained for eclogite parageneses with using Grt-Cpx barometer of Simakov and Taylor (2000) show different trends for oxygen fugacity in the upper mantle at real apparent errors of the Grt–Cpx–Qu(Cs) sensors in  $\pm 0.8$  log units of  $f_{O_2}$ and in 5 kbars of pressure.  $P-T-f_{O_2}$  results obtained for eclogite inclusions within diamonds for Premier and Udachnaya pipes show a distinct kink in the oxygen-fugacity trajectory (Fig. 10 limbs I and II). At depths lower than 150 km, the oxygen fugacity decreases with depth (limb I). Between the level of 100-150 km, there is a tendency for oxygen fugacity to rise with increasing depth (limb II). At both the localities, maximum of oxygen fugacity corresponds to a depth of 140-160 km. From these data, we can infer that there are two main levels of diamond formation in the upper mantle: (1) crystallization of initial as then osphere diamonds under very low-  $f_{O_2}$  conditions at depths of 160-180 km; (2) diamond crystallization and growth in equilibrium with water- and carbon dioxide-rich fluids in the lithosphere at the depths of 130–160 km. In other words, we can conclude that the asthenospheric diamonds form at high pressures, low-oxygen fugacities, and CH<sub>4</sub>-rich fluids, whereas the lithospheric diamonds form at lower pressures, higher oxygen fugacities, and in  $H_2O-CO_2$ -rich fluids. The oxygen fugacity mainly decreases from the oxidized lithosphere to the reduced asthenosphere.

#### Mantle model

The observed oxygen-fugacity trajectory can be explained by the hot plume uplifting under the central parts of cratons (Fig. 11). The plume could be intruded from the deep levels of the asthenosphere to the lithosphere–asthenosphere boundary or the lower parts of the lithosphere, which contained subducted

Fig. 11 Hypothetical crosssection of an Archean craton and adjacent cratonized mobile belt in accordance with Haggerty (1986) and Mitchell (1991) models, showing the variation of the oxygenreduction conditions in the upper mantle





**Fig. 12** The result of calculation of O–H–N–C system up to SiC–C boundary [by the data of Woerman and Rosenhauer (1985)] at P = 45 kbar and  $T = 1,000^{\circ}$ C and  $P_{N_2} = 1$  bar. (Calculations of  $\gamma_i$  of the gases were performed by the methods of Belonoshko and Saxena (1992), who fit their expressions to experimental data for molar volume of these components up to 6,090° K and 80 GPa)

rocks with carbonates. The plume melts could contain ancient harzburgite diamonds formed in reduced conditions in the asthenosphere from metal-silicate melts. The plumes could contact with lithosphere subducted rocks (Taylor and Green 1989). Under the process of heating, the reaction  $Dol + Coes = liq + CO_2$  would be crossed, releasing  $CO_2$ . This is supported by the isotopic data, which shows that light crustal carbon participated in the formation of eclogite diamonds (Deines et al. 1987). From this model, it follows that the part of the eclogitic diamonds could be formed as a result of the interactions between CO<sub>2</sub> released during heating and the reduced upper mantle fluid components. As a result, the initial garnet peridotites could be melted forming carbon-silicate melts and water-carbon dioxide fluids. Diamonds could be formed from these melts (Pal'yanov et al. 1999), the traces of which were founded in natural diamonds (Schrauder and Navon 1993). The Mitchell (1991) reconstruction of the lithosphere and asthenosphere profile indicates that eclogites can form during the contact of the plume magmas with the garnet lherzolite part of lithosphere or with the remnants of ancient subducted rocks (Mitchell 1991) (see Fig. 11). Crystallization of the diamonds with lherzolite and eclogite parageneses in equilibrium with H<sub>2</sub>O-CO<sub>2</sub>-rich fluids could take place from carbon-silicate melts as a result of these processes.

At the levels of 130-160 km, there was a peak of the  $CO_2$  releasing and as a result, there is a maximum of the eclogite parageneses oxidation, which is two to three log units higher than the peridotite trend at the

same depths (see Figs. 9, 10). Deines et al. (1989) noted the tendency for the nitrogen content of eclogitic Premier diamonds to decrease with temperature of formation. It follows that the nitrogen content decreases with temperature and with increasing oxygen fugacity (Figs. 7, 10a—limbs I and II). Sobolev et al. (1966) proposed that nitrogen was included into the diamond structure at a deep-seated stage of degassing of the mantle by the reaction, which decreases the decomposition of NH<sub>3</sub> to H<sup>-</sup> and N<sup>+3</sup>.

$$\mathrm{NH}_3 \to \mathrm{N} + 3 + 3\mathrm{H}^-. \tag{XI}$$

The problem of nitrogen presence in the upper mantle was discussed in the previous paper (Simakov 1998). From the calculations and mineralogical data, it follows that nitrogen could be present in the upper mantle in the forms of  $N_2$  and  $NH_3$  gases. The calculations in C–H–O–N system shows that ammonia concentration must be inversely proportional to CO<sub>2</sub> (Fig. 12). This agrees with the tendency for nitrogen content of the Premier diamonds to decrease with increasing oxygen fugacity or temperature.

#### Conclusions

The calculations above show that  $f_{O_2}$  in the lithospheric upper mantle under the Archean cratons varies over a range of five to six log units, which agrees with the previously obtained results for spinel and garnet peridotite xenoliths (Ballhaus 1993; Simakov 1998; Woodland and Koch 2003). These results indicate that the upper mantle is zoned and the degree of its reduction increases with depth from lithosphere to asthenosphere.

From the calculations it follows that mantle eclogite xenoliths are more oxidized than the peridotite ones (Fig. 8). It agrees with the experimental data of Pokhilenko and Tomilenko (2001). It is known that the peridotite diamond inclusions are more ancient than eclogitic ones (Richardson et al. 1993), which is why eclogitic diamonds are enriched by light crustal carbon relative to the peridotitic ones (Galimov 1984; Deines et al. 1987).

Diamond preservation in the mantle depends upon the oxygen fugacity. The optimum conditions for diamond growth and preservation in C–H–O system correspond to those where a water-rich fluid would be stable (if the upper mantle is be fluid-saturated), which agrees with the previously obtained results for peridotite xenoliths (Ballhaus 1993; Simakov 1998).

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#### **Appendix 1**

**Table 6** Garnet–orthopyroxene–clinopyroxene–olivine parageneses chosen for  $P-T-f_{O_2}$  calculations

Sample	NG85	H84	NW79	Ai93	T98	$\Delta log(f_{ m O_2})_1$	$\Delta log(f_{O_2})_2$
South African x	renoliths						
FRB1350 <sup>a</sup>	33	880	857	725	746	-3.6	-3.0
89–719 <sup>a</sup>	47	1,130	1,225	1,120	1,121	-4.2	-4.6
VBD1140 <sup>a</sup>	40	937	910	964	897	-3.6	-3.1
BD1150 <sup>a</sup>	41	944	958	1,024	1,024	-2.8	-2.7
FRB909 <sup>a</sup>	52	1,274	1,348	1,380	1,343	-3.7	-3.6
FRB921 <sup>a</sup>	38	1,000	1,009	1,050	1,000	-3.8	-3.4
Canadian slave	xenoliths						
10–12a <sup>a</sup>	36	788	726	749	771	-2.9	-2.7
22–5 <sup>a</sup>	48	970	842	976	909	-1.9	-3.3
25–9 <sup>a</sup>	40	875	837	880	813	-3.3	-4.1
$14-107^{a}$	46	958	968	1,010	995	-2.5	-2.4
40–11 <sup>a</sup>	48	99	1,018	1,060	1,052	-3.1	-3.8
21–6 <sup>a</sup>	51	1,024	1,025	1,075	1,110	-4.8	-4.3
$21 - 4^{a}$	46	965	967	1.043	1,021	-3.3	-3.5
$21-3^{a}$	48	999	970	1,045	1,022	-4.6	-4.0
22–7 <sup>a</sup>	53	1,030	1,040	1,170	1,120	-3.4	-3.6
$41-1^{a}$	50	967	960	1,100	1,053	-3.2	-2.7
$23-5^{a}$	45	1,045	1,074	1,180	1,190	-2.6	-2.7
9–10 <sup>a</sup>	47	988	1,013	1,143	1,129	-2.5	-3.0
41–3 <sup>a</sup>	47	1,101	1,112	1,245	1,190	-2.5	-2.8
Lesotho xenolit		-,	-,	-,	-, •		
133	45	1,052	1,064	1,044	1,016	0.6	
A	50	1,155	1,202	1,315	1,217	-2.4	
B	43	1,009	1,050	1,138	990	-2.5	
BD1354 <sup>a</sup>	46	1,066	1,114	1,213	160	-3.0	-3.4
Namibian xeno		1,000	1,111	1,215	100	5.0	5.1
1	47	11.5	1,071	1,156	1,092	-2.4	
20	42	996	1,006	990	1,000	-3.9	
34	38	1,001	1,000	1,010	960	-1.7	
1	37	946	881	1,010	935	-1.9	
6	46	1,082	1,090	1,208	1,204	-3.6	
Yakutian Xeno		1,002	1,070	1,200	1,207	5.0	
Uv-624	43	847	856	874	847	-2.7	
BD-2125	45	1,037	1,120	995	1,054	-2.7	
UV417/89 <sup>a</sup>	50	974	955	963	850	-4.8	-4.4
$UV61/91^{a}$	51	1,031	1,045	1,117	1,168	-3.8	-3.4
0 101/91	51	1,051	1,045	1,11/	1,100	-5.0	-J. <del>4</del>

NG85, Grt–Opx barometer of Nickel and Green (1985) (in kbar); H84, Grt–Opx thermometer of Harley (1984); NW79, Grt–Ol thermometer of O'Neill and Wood (1979); Ai93, Grt–Cpx thermometer of Ai (1993); T98, Cpx–OPx thermometer of Taylor (1998) (in °C).  $\Delta log(f_{O_2})_1$ , calculated values by garnet–clinopyroxene–silica barometer relative to the QFM buffer of Frost (1991);  $\Delta log(f_{O_2})_2$ , calculated values by garnet–orthopyroxene–olivine barometer of Gudmundson and Wood (1995) relative to the QFM buffer. For  $f_{O_2}$  calculations P-T parameters obtained from garnet–orthopyroxene, thermobarometers of Harley (1984) and Nickel and Green (1985) were used. Analyses for the calculations were taken from: Carswell et al. (1979), Mitchell (1984), Pokhilenko et al. (1991), Sobolev et al. (1984) <sup>a</sup> specimens with garnet Fe<sup>+3</sup>/ $\Sigma$  Fe measured by the Mossbauer technique from Canil and O'Neil (1996) and McCammon and Kopylova (2004).

# **Appendix 2**

Table 7 Eclogite garnet-clinopyroxene parageneses chosen for  $f_{O_2}$  calculations

Sample	EG79	P(kbar)	(Si+Ti) <sub>Grt</sub>	%Fe <sup>+3</sup>	%Andr	$\Delta log(f_{\mathrm{O}_2})$
Yakutian incl	usions in diamond	S				
M-734	1,267	45.0	3.02	4.85	2.35	-5.4
M-742	1,285	45.0	3.03	5.97	3.33	-4.0
M-636	1,085	45.0	3.04	5.03	3.18	-4.7
M-748	1,229	45.0	3.03	1.56	0.88	-5.8

Table 7 Eclogite garnet-clinopyroxene parageneses chosen for  $f_{O_2}$  calculations

Sample	EG79	P(kbar)	$(Si + Ti)_{Grt}$	%Fe <sup>+3</sup>	%Andr	$\Delta log(f_{O_2})$
M-422	1,085	45.0	3.05	4.56	2.60	-3.6
M-423	1,305	45.0	3.03	7.26	4.13	-2.4
M-46D	986	45.0	3.02	2.52	1.67	-3.0
U-41/3	1,140	45.0	3.02	7.47	5.04	-2.6
U-66/3	1,192	45.0	3.01	7.02	3.95	-3.2
Ud-4/90-1	1,132	45.0	3.03	4.72	2.87	-2.9
Ud-4/90-12-4	1,200	45.0	3.04	3.44	2.02	-4.1
Ud-4/90-13-6	1,138	45.0	3.03	3.68	2.22	-3.7
Ud-10/90-23	996	45.0	3.04	1.03	0.72	-6.0
South African and Gu						
P-3	1,310	50.0	3.04	3.98	1.61	-4.9
P-5	1,252	50.0	3.02	4.18	2.64	-4.5
P-18	1,336	50.0	3.02	9.18	4.14	-3.7
P-35	1,323	50.0	3.02	11.28	5.11	-1.3
P-38	1,266	50.0	3.01	0.96	0.40	-8.1
P-44	1,267	50.0	3.01	6.53	3.49	-4.8
P-45	1,364	50.0	3.02	6.22	3.36	-3.0
P-46	1,299	50.0	3.01	11.99	5.46	-1.2
P-47	1,305	50.0	3.03	1.91	0.97	-5.0
P-65	1,255	50.0	3.04	4.17	2.56	-4.0
P-66	1,208	50.0	3.01	4.65	2.70	-3.7
P-103	1,298	50.0	3.01	6.54	3.97	-3.7
P-104	1,289	50.0	3.01	11.36	5.22	-1.1
P-109	1,200	50.0	3.05	0.65	0.40	-7.9
P-113	1,332	50.0	3.04	5.42	2.48	-2.7
P-115	1,315	50.0	3.02	11.62	5.33	-1.7
RV-69	1,277	50.0	3.02	0.61	0.40	-8.1
KK-96	1,258	50.0	3.01	6.95	3.33	-3.5
7 <sup>a</sup>	1,180	50.0		3.33	2.06	-4.2
20 <sup>a</sup>	1,113	50.0		3.14	1.96	-2.5
41 <sup>a</sup>	1,138	50.0		5.03	3.04	-3.0
Australian inclusions		-	2.05	o. / 1	2.04	
A29	1,267	50.0	3.05	3.41	2.04	-3.3
A33	1,236	50.0	3.03	3.17	1.83	-3.8
A31	1,278	50.0	3.03	6.52	2.71	-1.4
Venezuela inclusions i		50.0	2.02	2 (1	1 40	2.0
Gm-54	1,103	50.0	3.03	2.61	1.48	-3.9
037c-a v-19	1,188 1,171	50.0	3.03	3.12	1.62	-3.6
Yakutian diamond-be		50.0	3.02	5.00	2.83	-3.1
		45.0	3.02	5.26	2.51	-4.7
M-45 M-46X	1,058	45.0		5.26	2.51	-4.7 -4.0
M-46X	1,034	45.0	3.04	2.15	1.28	-4.0 -4.1
M-52 M-50	1,020 1,004	45.0 45.0	3.03 3.04	6.54 9.76	3.82 6.54	-4.1 -7.1
M-49	1,004	45.0	3.04	1.43	0.76	-4.8
A-811	989	45.0	3.04	3.55	2.06	-2.2
M-54-R1	1,040	45.0	3.04	0.54	0.24	-2.2 -5.6
BM-432	1,040	45.0	3.02	10.12	4.60	-0.6
U-464/86	994	45.0	3.02	5.51	4.08	-3.3
U-58/2	1,195	45.0	3.01	8.91	4.08	-2.5
U-55/2	1,074	45.0	3.01	10.30	4.23	
U-8/1	1,074	45.0	3.01	10.30	4.21 4.53	-0.8 -2.9
Ud-76/2	1,271	45.0	3.04	4.70	2.08	-2.9 -4.4
Ud-92/2	1,230	45.0	3.05	2.07	0.83	-4.4 -5.1
Ud-59/2	1,198	45.0	3.02	4.59	1.47	-1.5
Ud-28	1,198	45.0	3.02	4.59 3.67	2.56	-1.3 -2.8
M-85	1,077	45.0	3.02	1.71	1.03	-2.8 -4.5
M-2110	1,090	45.0	3.04	2.01	1.28	-4.8
M-63	974	45.0	3.02	0.77	0.38	-5.7
236 <sup>a</sup>	1,007	45.0	5.62	10.65	3.71	-2.3
$\frac{250}{281/2^{a}}$	1,098	45.0		19.94	7.40	-2.7
African diamond-bear		10.0		17.77		2.1
6	1,136	45.0	3.02	3.84	1.09	-1.9
17	1,700	45.0	3.02	6.10	2.12	-1.5
18	1.191	45.0	3.02	6.10	2.12	-0.6
19	1,297	45.0	3.02	4.55	1.68	-0.0
XM26	1,178	45.0	3.02	6.16	3.03	-2.4
1 11120	1,1/0	12.0	5.01	0.10	5.05	<i>2.</i> ¬

**Table 7** Eclogite garnet–clinopyroxene parageneses chosen for  $f_{O_2}$  calculations

Sample	EG79	P(kbar)	(Si+Ti) <sub>Grt</sub>	%Fe <sup>+3</sup>	%Andr	$\Delta log(f_{\mathrm{O}_2})$
JJG891	1,359	45.0	3.02	0.55	0.21	-7.3
KEC-86-DB-9	1,344	45.0	3.04	9.12	4.30	-1.7
Finnish diamond-bea	aring xenoliths					
L-78	1,275	45.0	3.02	15.14	4.21	-3.6
L-80	1,248	45.0	3.01	8.36	3.31	-3.0
SEI-4	1,187	45.0	3.01	2.96	0.82	-3.6
African xenoliths						
6913-5	1,113	45.0	3.01	4.79	2.16	-2.6
6913-10	1,128	45.0	3.02	8.83	3.79	-0.1
2	1,444	45.0	3.01	4.46	1.24	-1.1
4	1,196	45.0	3.03	4.36	1.21	-2.6
1	1,164	45.0	3.03	0.62	0.27	-7.4
4	1,114	45.0	3.03	3.75	2.02	-6.7
5	1,145	45.0	3.02	1.97	1.03	-7.5
8	1,066	45.0	3.02	4.98	2.96	-3.2
K-3	1,118	45.0	3.02	6.18	3.02	-4.2
K-5 K-4	1,131	45.0	3.02	5.85	3.41	-5.1
K-4 K-5	1,159	45.0	3.4	8.36	4.78	-3.1 -8.0
SRV1	1,139	45.0	3.03	2.16	0.69	-1.8
	1,143		3.02	6.08		-0.5
R-7		45.0			3.00	-0.5
R-8A	993	45.0	3.03	10.63	3.13	-2.0
R-22	1,063	45.0	3.02	13.19	9.39	-2.1
R-35	1,260	45.0	3.02	4.19	1.70	-1.0
R-51B	1,103	45.0	3.05	6.56	3.38	-4.1
R-71	1,042	45.0	3.01	3.52	1.31	-5.3
R-76A	1,242	45.0	3.02	4.64	1.96	-3.8
R-76B	1,256	45.0	3.03	4.34	1.77	-5.4
K-86-GB-70	934	45.0	3.03	0.45	0.29	-8.4
K-86-KB-3	1,035	45.0	3.02	11.28	4.17	0.1z
K-86-KB-74B	960	45.0	3.02	6.03	2.74	0.0
KG-86-76	978	45.0	3.01	6.03	3.72	-3.2
K-86-11	934	45.0	3.03	0.51	0.25	-5.4
K-86-72A	944	45.0	3.01	4.87	3.60	1.3
K-86-72B	1,004	45.0	3.02	6.97	4.85	-1.6
K-86-2	1,010	45.0	3.01	5.03	1.89	-2.4
K-86-58	923	45.0	3.03	11.63	2.56	-2.6
K-86-73A	1,070	45.0	3.03	6.20	2.19	-3.8
K-86-73B	977	45.0	3.01	10.81	3.83	-3.3
K-86-90	1,067	45.0	3.03	9.07	3.38	-2.5
K-86-107	1,086	45.0	3.03	7.47	1.67	-0.9
K-81-12	935	45.0	3.02	3.92	3.30	-4.3
K-81-21	1,048	45.0	3.01	3.55	2.14	-3.7
	1,0.0		2.01	2.22	2	2

 $(\% Fe^{+3} - 100*Fe^{+3})/(Fe^{+3} + Fe^{+2})$  in the garnets; %Andr, percent of andradite component;  $(Si + Ti)_{Grt}$ , silica and titan content in the garnet (in f.u.); EG79, thermometers of Ellis and Green (1979); *P*, pressure in kbar;  $\Delta \log(f_{102})$ , calculated values by garnet–clinopy-roxene–silica barometer relative to the QFM buffer of Frost (1991). Analyses for the calculations were taken from: Beard et al. (1996), Gurney et al. (1985), Hills and Haggerty (1989), Harte and Gurney (1975), Jaques et al. (1989), Jerde et al. (1993), Kaminsky et al. (2000), Kushiro and Aoki (1968), Lappin and Dawson (1975), MacGgregor and Manton (1986), O'Hara et al. (1975), Reid et al. (1976), Safronov et al. (1980), Smyth and Hatton (1977), Sobolev (1977), Sobolev et al. (1983, 1998), Spetsius et al. (1992), Stachel et al. (2000), Taylor et al. (1996)

<sup>a</sup> specimens with garnet and clinopyroxene; Fe<sup>+3</sup> / $\Sigma$  Fe Mossbauer taken from McCammon et al. (1998) and Sobolev et al. (1999b)

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