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Geochemical and O-isotope constraints on the evolution of lithospheric mantle in the Ross Sea rift area (Antarctica)

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Abstract Peridotite xenoliths found in Cenozoic alkali basalts of northern Victoria Land, Antarctica, vary from fertile spinel-lherzolite to harzburgite. They often contain glass-bearing pockets formed after primary pyroxenes and spinel. Few samples are composite and consist of depleted spinel lherzolite crosscut by amphibole veins and/or lherzolite in contact with poikilitic wehrlite. Peridotite xenoliths are characterized by negative Al₂O₃-Mg# and TiO₂-Mg# covariations of clino- and orthopyroxenes, low to intermediate HREE concentrations in clinopyroxene, negative Cr-Al trend in spinel, suggesting variable degrees of partial melting. Metasomatic overprint is evidenced by trace element enrichment in clinopyroxene and sporadic increase of Ti-Fe_{tot}. Preferential Nb, Zr, Sr enrichments in clinopyroxene associated with high Ti-Fetot contents constrain the metasomatic agent to be an alkaline basic melt. In composite xenoliths, clinopyroxene REE contents increase next to the veins suggesting metasomatic diffusion of incompatible element. Oxygen isotope data indicate disequilibrium conditions among clinopyroxene, olivine and orthopyroxene. The highest δ^{18} O values are observed in minerals of the amphibole-bearing xenolith. The $\delta^{18}O_{cpx}$ correlations with clinopyroxene modal abundance and geochemical parameters (e.g. Mg# and Cr#) suggest a possible influence of partial melting on oxygen isotope composition. Thermobarometric estimates define a geotherm of 80°C/GPa for the refractory lithosphere of NVL, in a pressure range between 1 and 2.5 GPa.

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L. Dallai IGG – CNR, Via Moruzzi 1, 56124 Pisa, Italy E-mail: dallai@igg.cnr.it Clinopyroxene microlites of melt pockets provide P–T data close to the anhydrous peridotite solidus and confirm that they originated from heating and decompression during transport in the host magma. All these geothermometric data constrain the mantle potential temperature to values of 1250–1350°C, consistent with the occurrence of mantle decompressional melting in a transtensive tectonic regime for the Ross Sea region.

Keywords Metasomatism · Lithospheric mantle · Mantle geotherm · Ross Sea Rift · Antarctica

Introduction

Spinel peridotite xenoliths found in alkali basalts provide evidence for the composition of lithospheric mantle and constrain the nature, evolution and thermal state of the uppermost mantle in regions of continental rift.

Ultramafic nodules are abundant in basic lavas of the McMurdo Volcanic Group (MMVG), in the Ross Sea region of Antarctica, an area of ongoing rifting where magmatism has been active since Cenozoic times. A general consensus for the origin of this magmatism has not yet been achieved. The main hypotheses proposed can be summarized as follows: (1) active rifting due to the presence of a thermally buoyant plume; (2) passive rifting related to transtensional tectonics and consequent melting of the low-melting domains in the lithospheric mantle. The latter hypothesis included the possibility that partial melting preferentially sampled a fossil HIMU plume emplaced during Gondwana break-up (LeMasurier and Rex 1989; Thessenson and Wörner 1991; Kyle et al. 1992; Rocholl et al. 1995; LeMasurier and Landis 1996; Hart et al. 1997; Wörner 1999; Storti et al. 2001; Rocchi et al. 2002).

Previous studies have shown that the mantle in this area is heterogeneous and that mantle xenoliths preserved evidence for partial melting and metasomatic processes (Berg et al. 1989; Hornig and Wörner 1991; Beccaluva et al. 1991; Hornig et al. 1992; Zipfel and

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and Wörner (1992), the mantle xenoliths record three main stages: (1) adiabatic rise and dynamic hightemperature recrystallization of mantle phases at lower pressure; (2) cooling and re-crystallization under low pressure; (3) local heating of lithospheric mantle linked to the magmatism of Ross Sea Rift. During this late stage and before xenoliths were brought to surface, metasomatic processes affected lithospheric mantle domains of this region likely due to the infiltration of basaltic melts at depth (Coltorti et al. 2004).

In this study we investigated two suites of spinel peridotites for major element, trace element and O-isotope composition in order to constrain their petrologic evolution and to define the thermal state of the lithosphere. Combining the REE distribution with oxygen isotope data, we constrained the processes of partial melting and metasomatism affecting the lithospheric mantle during the Cenozoic. 1991), a region of lithospheric thinning marked by a topographical trough from the Antarctic Peninsula to the Ross Embayment-northern Victoria Land (Fig. 1) (Behrendt et al. 1991, 1992; LeMasurier and Thomson 1990). The western rift flank in northern Victoria Land (NVL) is formed by the Transantarctic Mountains and represents the uplifted roots of the early Palaeozoic Ross Orogen. The uplift of the Transantarctic Mountains was substantially amagmatic and it was accompanied by extensional tectonics in the Ross Sea region during late Cretaceous time (e.g. Chand et al. 2001; Fitzgerald 1994; Fitzgerald and Stump 1997; Stagg and Willcox 1992).

Since the Eocene, diffuse igneous activity has characterized the WARS. In northern Victoria Land, plutons, like swarms and volcanoes were emplaced in an area of about 400×80 km (Fitzgerald and Stump 1997; Rocchi et al. 2002; Tonarini et al. 1997). The volcanic products were grouped in the McMurdo Volcanic Group (Kyle 1990), while intrusive-subvolcanic varieties are collectively referred to as Meander Intrusive Group (Muller et al. 1991; Tonarini et al. 1997). Mantle xenoliths were brought to surface by the primitive alkaline magmas (melanephelinite, basanite and alkali olivine basalt). These magmas erupted from monogenetic cinder

Geological setting

The MMVG igneous activity is linked to the West Antarctic Rift System (WARS, Tessensohn and Wörner

Fig. 1 Sketch map of north Victoria Land and sampling localities



cones near the margin of larger volcanoes or scattered on the crystalline basement, where the major active faults likely provided direct pathways for ascending to surface (Salvini et al. 1997; Orlando et al. 1997).

Analytical methods

The major elements of the whole rock were determined on powder pellets by XRF (Philips PW 1480) with the Franzini et al. (1975) correction, while trace elements were determined for selected samples by ICP-MS (Fisons PO2 Plus®) at "Dipartimento di Scienze della Terra", Pisa University. Samples were dissolved in screw-top PFA vessels on a hotplate at $\sim 120^{\circ}$ C with HF-HNO3 mixture. Analyses were performed by external calibration using basaltic geochemical reference samples as composition- and matrix-matching calibration solutions. The correction procedure includes (i) blank subtraction; (ii) instrumental drift correction using Rh-Re-Bi internal standardization and repeated (every five samples) analysis of a drift monitor; (iii) oxide-hydroxide interference correction. Precision, evaluated by replicate dissolutions and analyses of the in-house standard HE-1 (Mt. Etna hawaiite), is generally between 2 and 5% RSD, except for Gd (6%), Tm (7%), Pb and Sc (8%). Detection limits are in the range 0.002-0.02 ng ml⁻¹ in the solution (corresponding to 0.002-0.02 ppm for a 1,000fold sample dilution) for all the elements, except for Ba, Pb and Sr $(0.1 \div 0.2)$.

The major elements of mineral and glass were analysed by CAMECA electron microprobe (Instituto di Geoscienze e Georisorse, Rome Section, Italy) operated at 15 kV accelerating voltage and 30 nA beam current. Counting time was 100 s. To reduce alkali loss during glass analysis, we decreased the counting time to 20 s and the beam was defocused to 15 μ m. ZAF correction was applied to all data.

The concentration of trace elements in minerals of GP xenoliths were determined by laser ablation microprobe-inductively coupled plasma- mass spectrometry analysis (LAM-ICP-MS) at CNR- Istituto di Geoscienze e Georisorse in Pavia, Italy (operating conditions, analytical procedures, accuracy and precision have been described by Bottazzi et al. 1994).

Trace elements in minerals and intergranular glasses were determined by secondary ion mass spectrometry (SIMS) on polished sections at the LAM-ICP-MS laboratory of Pavia. Analytical procedures have been described by Wulff-Pedersen et al. (1999).

Oxygen isotope data were measured by conventional laser fluorination (Sharp 1995) at the CNR-Istituto di Geologia Ambientale e Geoingegneria in Rome, Italy. A 15 W CO₂ laser was employed to irradiate the samples and pure fluorine desorbed at 290–310°C from hexafluoropotassium–nickelate salt (Asprey 1976) was used to react 1–1.5 mg of cpx, opx, spl and ol fragments. The O₂ produced was purified of excess fluorine by a KBr trap at 160°C and of trace NF_x by-products of laser fluorination using a 13 Å zeolite molecular sieve and liquid nitrogen-ethanol mixture (Clayton and Mayeda 1983). The gas was then introduced into a Finnigan delta plus mass spectrometer and analyzed for oxygen isotope composition and mineral yields were controlled to be quantitative by an empirical regression curve based on samples with known δ^{18} O versus P_{O2} values in the MS bellow. A total of 23 NBS28 (δ^{18} O = 9.60%) standard samples were measured during the time of this study, with an average δ^{18} O value of $9.54 \pm 0.17\%$ (2 σ). During each set of analyses, four to seven aliquots of laboratory quartz standards were measured with an average reproducibility of $\pm 0.10\%$ (1 σ) and no data correction was adopted. All the results are reported in the standard per mil notation and the δ^{18} O values are relative to SMOW. At least two fragments were analyzed for each mineral and variation within the same sample is less than the precision of standards.

Samples

Spinel peridotite xenoliths were collected at Greene Point and Baker Rocks (Fig. 1). They show variable dimensions ranging from 3 to 10 cm. Modal compositions of larger nodules (>5 cm) were estimated by counting more than 1,500 points in each thin section (Table 1). Xenoliths show a continuous mineralogical variation from moderately depleted lherzolite (11 to 12% diopside) to harzburgite with 0–4% of diopside (Table 1).

The most common texture is coarse-grained protogranular, but porphyroclastic and equigranular varieties also occur. Olivines and orthopyroxenes of protogranular samples are equant and show curvilinear grain boundaries. The maximum size reached is about 1 cm; typically, olivine crystal size shows a scale invariant distribution due to brittle fracturing (Armienti and Tarquini 2002). Clinopyroxene occurs both as a primary phase (CpxI) and microlites in pyrometamorphic glass patches (CpxII). Cpx I size is related to the depletion degree of the whole rock, the larger grains (up to 2 mm) occurring in the more fertile samples (modal clinopyroxene displays exsolution lamellae of clinopyroxene.

Spinel is commonly found next to diopside, sometimes showing symplectitic intergrowths with enstatite crystals. Reaction textures of spinel consist of colourless glass, neoblastic clinopyroxene and olivine. In porphyroclastic xenoliths, "holly leaf" spinel occurs in re-crystallized areas and defines a foliation.

A few samples have pyrometamorphic textures with spongy pyroxenes and grain rims replaced by a reaction assemblage of clinoproxene + olivine + glass. There is no evidence of direct connection between glass blebs and host lava due to the lack of veinlets of host magma in the xenoliths.

Table 1 Modes for xenoliths from Greene Point and Baker Rocks

	Sample	Rock	Ol	Opx	Срх	Sp	Amph
	BR-213	Lh	69.6	22.5	7.3	0.6	
Composite	BR-214	Weh	73.9	4.6	20.9	0.6	
1	BR-214	Lh	68.4	24.3	6.7	0.6	
Composite	BR-218	Lh	71.9	19.2	8.9	0.2	
1	BR-218	Pargasite vein					100.0
Composite	BR-219	Weh	61.0	4.4	34.3	0.3	
1	BR-219	Hz	80.7	15.5	3.6	0.2	
	GP1	Hz-Lh	70.4	23.3	5.1	1.2	
	GP5	Hz	74.5	18.7	4.9	1.9	
	GP6	Hz	76.5	18.2	4.2	1.2	
	GP10	Hz	75.4	17.3	5.4	1.9	
	GP14	Lh	66.2	20.5	11.9	1.4	
	GP15	Hz	74.8	23.3	0.9	1.0	
	GP17	Hz	71.9	23.7	4.0	0.4	
	GP20	Lh	77.3	16.3	6.1	0.3	
	GP21	Hz	85.0	14.4	0.0	0.6	
	GP22	Hz	77.7	17.0	3.8	1.5	
	GP25	Hz	82.2	16.0	1.0	0.8	
	GP28	Lz	73.1	22.5	4.0	0.4	
	GP29	Hz	72.8	22.7	3.8	0.7	
	GP31	Hz	80.1	17.5	1.6	0.8	
	GP34	Lh	71.6	21.3	6.5	0.6	
	GP36	Lh	70.5	22.3	6.2	1.0	
	GP49	Hz	73.1	21.8	4.4	0.7	

Modal compositions are determined by point counting in thin section (1500 points) Lh lherzolite, Weh wehrlite, Hz harzburgite

Samples BR218 and BR219 are composite: in BR218 an amphibole-bearing vein (2 mm wide) cuts the lherzolite and in BR219 a cumulate wehrlite adheres to a coarse harzburgite. Amphibole in the vein is idiomorphic, while in the wehrlite poikilitic clinopyroxenes include rounded olivines. Amphibole vein and cumulate wehrlite have blebs of pale yellow glass with microclinopyroxene + olivine \pm plagioclase. crystals of Further details of individual samples are given in Perinelli et al. (1998).

Chemical composition

Whole rocks

Measured compositions fall in the field of worldwide spinel peridotites (Table 2) and define different degrees of melt extraction (Frey and Green 1974; Nickel and Green 1984; Frey et al. 1985). Covariation diagrams show that Al₂O₃ and CaO contents are inversely correlated with MgO (Fig. 2). Relatively higher CaO contents (1.3 wt%; sample GP21) are probably due to the occurrence of secondary calcite. SiO₂, TiO₂ and Na₂O define poor correlations with MgO. The Mg# is rather uniform (90-91.4) and only peridotite domains of composite xenoliths have lower Mg# (86.8), possibly indicating FeO enrichment (Table 2).

Trace element patterns normalised to primitive mantle (McDonough and Sun 1995) show moderate to high incompatible element enrichments (Fig 3a). Xenoliths with incipient pyrometamorphic features show Ba,

Nb and Sr negative anomalies while the other samples display high values of Nb, Ta, Zr and a negative Ti anomaly. All samples have concave-upward REE_N with (Ce/Yb)_N ratios between 1.8 and 17.3 (Fig. 3b). The few samples with Nb/Ta negative anomaly inherit these features from clinopyroxene composition.

Olivine

Olivines have a small compositional range (Fo₈₉–Fo₉₂, Table 3). In composite xenoliths, Fo-contents decrease towards the veins (from Fo₉₁ to Fo₈₆) due to Mg-Fe exchange reactions between the vein and whole rock. In porphyroclastic samples, olivine neoblasts have lower NiO contents than porphyroclasts (0.10–0.18 and 0.40– 0.75 wt%, respectively). Higher CaO contents and low NiO characterize olivine of pyrometamorphic glass patches (Table 3).

Orthopyroxene

This phase generally shows constant compositions $(En_{88,2-91,2}, Wo_{1,9-1,8}, Fs_{9,9-7}, Fig. 4)$ with Mg# values between 89.5 and 93.2, comparable with those of coexisting olivine (Table 3). Al₂O₃ contents decrease from lherzolite to harzburgite, whereas Cr₂O₃ increases (Table 3). In composite xenoliths, orthopyroxene, next to amphibole or wehrlite vein (BR218 and BR219), is enriched in FeO (Mg#=87.8 and 85.6, respectively) and TiO₂. Chondrite normalised REE patterns of depleted

Table 2 Whole rock compositions of xenoliths from NVL

XRF	BR 213	BR 218	BR 219	GP 5	GP 6	GP 10	GP 14	GP 21	GP 22
SiO ₂	44.14	43.99	40.25	42.19	42.60	42.09	44.59	41.84	43.46
TiO ₂	0.08	0.04	0.20	0.06	0.05	0.04	0.10	0.21	0.14
Al_2O_3	1.85	2.10	1.73	1.02	0.90	0.92	2.95	1.38	1.15
FeO*	7.70	9.76	11.66	8.59	7.75	7.86	7.15	7.88	7.70
MnO	0.12	0.16	0.18	0.14	0.14	0.13	0.13	0.14	0.14
MgO	43.50	41.26	42.95	44.42	45.10	46.48	41.22	44.62	43.96
CaO	0.98	1.87	1.33	0.82	0.83	0.78	2.51	1.30	0.89
Na ₂ O	0.17	0.11	0.18	0.19	0.10	0.21	0.33	0.38	0.26
K ₂ O	0.04	0.03	0.07	0.09	0.07	0.07	0.10	0.14	0.13
P_2O_5	0.02	0.02	0.03	0.05	0.04	0.05	0.03	0.08	0.10
H_2O	0.40		0.07	0.12	0.16	0.27		0.11	0.28
CO ₂	0.47		0.10	0.15	0.10	0.19		0.09	0.10
Total	99.47	99.34	98.75	97.84	97.84	99.09	99.10	98.17	98.31
Mg# ICP-MS	90.97	88.28	86.78	90.21	91.21	91.33	91.13	90.99	91.05
V	45	76	56	34	30	26	62	37	31
Cr	2975	2395	3096	2611	2715	2666	2883	3184	2601
Co	109.2	104	115	118.8	112.2	118.9	111	108	115.9
Ni	1926	1813	2011	1983	2006	2130	1849	1998	2051
Rb	2.26	1.95	2.68	3.11	3.32	3.26	1.3	3.18	3.4
Sr	18.92	15.1	29.29	38.59	37.47	33.69	11.1	80.51	66.85
Y	1.2	2.66	2.1	1.5	0.8	1.01	1.34	1.6	1.5
Zr	6.08	6	8.14	5.58	5.14	6.05	5.7	6.92	4.23
Nb	1.77	1.23	2.65	1.4	2.33	2.4	0.17	3.4	2.52
Mo	0.87		1.18	1.23	1.41	1.26		1.15	1.34
Cs	0.04	0.33	0.12	0.08	0.06	0.06	0.36	0.13	0.09
Ba	13.18	7.1	20.14	23.93	15.99	21.14	0.68	48.11	39.36
La	1.33	0.79	2.22	3.24	1.68	2.96	0.79	3.05	3.44
Ce	2.64	1.93	4.73	6.49	3.19	3.69	1.84	4.52	3.85
Pr	0.253	0.21	0.458	0.554	0.28	0.382	0.25	0.325	0.331
Nd	1.09	0.88	2.16	2.21	1.1	1.89	0.97	1.91	1.65
Sm	0.27	0.24	0.47	0.46	0.23	0.38	0.17	0.37	0.38
Eu	0.08	0.09	0.179	0.138	0.069	0.115	0.06	0.183	0.145
Gd	0.21	0.35	0.46	0.32	0.15	0.31	0.18	0.33	0.37
Tb	0.03	0.06	0.06	0.05	0.02	0.04	0.03	0.06	0.08
Dy	0.21	0.45	0.36	0.25	0.13	0.23	0.21	0.31	0.38
Ho	0.04	0.1	0.07	0.05	0.03	0.04	0.04	0.08	0.07
Er	0.12	0.28	0.2	0.15	0.08	0.14	0.15	0.17	0.21
Tm	0.023	0.04	0.024	0.018	0.011	0.022	0.02	0.038	0.029
Yb	0.15	0.18	0.16	0.11	0.08	0.09	0.12	0.12	0.12
Lu	0.025	0.04	0.024	0.021	0.012	0.019	0.03	0.018	0.029
Hf	0.24	0.16	0.22	0.16	0.14	0.17	0.11	0.18	0.12
Та	0.12	0.26	0.195	0.279	0.147	0.238	0.05	0.211	0.25
T1	0.05	0.01	0.05	0.09	0.04	0.03	0.01	0.03	0.03
Th	0.179	0.13	0.408	0.411	0.183	0.53	0.05	0.934	0.705
U	0.055	0.02	0.092	0.134	0.079	0.199	0.02	0.214	0.182

FeO* total iron as FeO

samples are enriched in LREE (Fig. 4a, b) and show trace element zoning from core to rim (Fig. 4c). Patterns of more fertile xenoliths (e.g. GP36 or GP42) display a smooth decrease from HREE to LREE (Fig. 4a, b).

Clinopyroxene

CpxI which are mainly diopsidic (Wo_{44,4-48,2}, En_{45,6-52}, Fs_{3,6-6.2}; Table 3) have Mg# values (89.5–94.4) that correlate with the Fo contents of olivine and that are higher than Mg# of coexisting olivine and orthopyroxene. The Al₂O₃, TiO₂ and Na₂O contents decrease with increasing Mg# (Table 3). Pyrometamorphic crystals exhibit a decrease of Al₂O₃ and Na₂O from core to rim. In composite xenoliths clinopyroxene next to veins is enriched in TiO_2 and FeO.

CpxII microlites have higher Cr_2O_3 contents and a more homogeneous Mg# (90.7–93.6) than CpxI (Table 3).

CpxI of different xenoliths reveals marked differences in trace element composition and shows variable enrichments of incompatible elements (Fig. 5a, c, e). Samples GP6, GP32 and GP35 also show intra-mineral zoning (Fig. 5g).

The REE contents (Table 4) allow for the CpxI to be characterized as:

1. Light REE (LREE)-depleted ("not-metasomatized") clinopyroxenes with (Ce/Yb)_N varying between 0.002

Fig. 2 Concentration of major in NVL xenoliths, plotted with respect to MgO. Shaded fields are defined by data of dry spinel peridotite xenoliths from different areas (GERM data repository, http:// www.earthref.org/GERM). The curves represent the near fractional melting and batch melting trend of a starting source of primitive spinel peridotite using the Niu (1997) model (after Takazawa et al. 2000). Data for primitive mantle (PM) are from McDonough and Sun (1995)



△ Baker Rocks

Rb Th Nb La Pr Nd Hf Eu Gd Dy Y Tm Lu Ba U Ta Ce Sr Zr Sm Ti Tb Ho Er Yb

LREE on the primitive normalized diagrams. They also show small to significant Zr and Ti negative anomalies.

La Ce Pr Nd Sm Eu Gd Tb Dy Ho Er Tm Yb Lu

Fig. 3 Incompatible trace element (a) and REE (b) patterns for spinel peridotites. *Shaded fields* represent the patterns of the remaining GP xenoliths. Trace element plots are ordered according to their decreasing incompatibility during partial melting. Normalizing values after McDonough and Sun (1995). Continental lithospheric mantle (CLM) data are from McDonough (1990)

and 0.58 (n = chondrite normalized value) and Yb_N from 2.9 to 10.8 (Fig. 5d). These clinopyroxenes are depleted in Th,U Nb and Ta with respect to

100

10

Whole Rock/Primitive Mantle

0.1

10

0.1

Table 3 Representative electron microprobe analyses of olivine, orthopyroxene, clinopyroxene and spinel of a selection of peridotites from northern Victoria Land

	BR 21	3				BR 21	8 far fr	om An	nph-v	ein	BR	218*	near	Amph	-vei	n	BR 2	218 Amj	oh-vein
	Ol	Opx	CpxI	CpxII	Sp	Ol	Opx	CpxI	Sj	þ	Ol		Opx	Сря	κI	Sp	Amp	h	CpxII
SiO ₂ TiO ₂ Al ₂ O ₃ FeO MnO MgO CaO Na ₂ O K ₂ O	$\begin{array}{c} 40.47\\ 0.09\\ 0.00\\ 9.01\\ 0.12\\ 48.84\\ 0.05 \end{array}$	56.58 0.16 3.30 5.85 0.09 32.54 0.92 0.41	52.51 0.38 5.21 2.55 0.00 16.24 20.24 1.66	48.67 1.54 6.19 2.38 0.03 15.33 22.76 0.29	$\begin{array}{c} 0.27 \\ 0.02 \\ 40.28 \\ 12.36 \\ 0 \\ 18.06 \end{array}$	40.24 0.00 0.00 11.97 0.02 47.31 0.07	55.38 4.00 6.70 33.10 0.76	52.39 0.49 5.19 2.84 0.18 15.54 21.46 1.31	0. 0. 49 13 0. 17	12 23 9.71 3.62 07 7.98	40.2 0.00 0.21 12.4 0.28 46.8 0.05	5 9 4	54.27 0.26 5.24 7.87 31.86 0.63	51.2 0.82 5.98 3.42 0.09 15.2 21.6 1.08	22 2 3 2 9 31 57 3	0.21 52.65 13.06 0.13 18.22	42.77 3.66 14.87 4.62 16.33 11.60 2.98 0.81	2 2 8 8	47.68 3.01 7.85 2.93 0.16 14.43 21.61 0.81
Cr ₂ O ₃ NiO Total Mg#	0.38 98.96 90.62	0.54 100.39 90.84	1.30 100.09 91.91	2.23 99.42 91.99	27.93 0.33 99.25 72.26	0.35 99.96 87.57	99.94 89.80	0.88 100.22 90.70	18 8 99 70	8.16 9.89).18	0.32 100.4 86.9	44 9	0.59 100.7 87.83	0.81 2 100 88.8	l .40 37	15.98 100.2 71.32	0.82 5 98.49 86.34	9 4	1.86 100.34 89.78
	BR 21	9 far fro	m Wehr	rlite	BR 2	19 near	from V	Vehrlite	;	GI	25							GP10	
	Ol	Opx	CpxI	Sp	Ol	Opx	Сря	KI S	р	Ol	II	Ol		Opx	C	pxI	Sp	Ol	Opx
SiO ₂ TiO ₂ Al ₂ O ₃ FeO MnO MgO CaO Na ₂ O	40.04 0.01 0.24 11.65 0.01 46.92 0.09	55.68 0.12 3.75 7.80 0.21 31.41 0.94 0.47	51.74 0.37 5.58 3.15 0.15 16.80 20.19 1.12	0.07 0.16 46.99 13.93 0.43 17.98	39.68 0.03 0.04 13.14 0.14 45.61 0.09	55.47 0.22 2.82 9.24 0.13 30.80 0.95 0.48	50.7 0.90 6.58 3.39 0.18 15.9 20.4	74 0. 0 0. 8 4. 9 1. 8 0. 91 1. 41	.07 .33 5.49 5.96 .32 7.02	39 0.0 0.3 15 0.1 43 0.2 0 1	.16 07 31 .20 .5 .82 26 0	40.72 0.05 0.07 9.39 0.08 49.1 0.08	2	56.73 0.08 2.91 5.75 0.28 33.35 0.71 0.44	53 0. 3. 2. 0. 16 22 0.	5.12 30 60 35 09 5.88 2.26 82	0.1 0.03 39.67 13.24 0.36 17.44	40.54 0.00 0.13 8.60 0.05 49.75 0.07	56.53 0.15 3.38 5.44 0.19 32.84 0.89 0.51
Cr ₂ O ₃ NiO Total Mg#	0.49 99.45 87.78	0.58 100.96 87.77	1.12 1.12 100.22 90.48	19.44 0.41 2 99.41 69.71	0.45 99.18 86.09	0.54 100.6 85.60	1.12 5 100 89.3	2 20 0 .68 10 32 6	0.74 .32 00.25 5.53	99 83	.07 .71	0.49 100.9 90.3	03 2	0.58 100.83 91.18	0.1 10 92	99 99 00.41 2.76	29.15 0.54 100.53 70.14	0.48 99.62 91.16	0.71 100.64 91.50
	GP10					GP14							GP1	6					GP28
	CpxI c	ore Cp	xI rim	CpxII	Sp	Ol	Opx	Срх	I	CpxI	l Sp)	Ol n	Ol	p	Opx	CpxI	Sp	Ol
SiO ₂ TiO ₂ Al ₂ O ₃ FeO MnO MgO CaO Na ₂ O	52.77 0.52 5.06 2.55 16.48 20.68 1.47	53. 0.6 2.9 2.5 17. 21. 0.6	18 5 3 0 67 82 9	53.02 0.79 2.78 2.17 0.06 17.05 22.10 0.85	0.19 0.49 38.09 11.99 0.58 17.37	40.93 0.04 0.25 9.75 0.33 49.13 0.08	54.95 0.15 5.43 6.11 0.17 31.37 0.91 0.53	52.0 0.48 6.69 2.74 0.16 15.5 19.8 1.60	9 2 4	52.91 3.42 2.36 0.00 16.23 22.73 0.79	0.1 0.0 56 9.8 20	1 09 0.45 83 0.63	41.3 0 8.97 0.35 49.9 0.13	7 40 0.0 9.3 0.2 2 48 0.0	.29)6 37 29 .17)7	55.58 0.06 3.3 5.65 0.14 32.37 1.07 0.5	52.44 0.24 3.99 2.68 0.16 16.78 21.04 0.92	0.29 0.24 34.09 11.41 0.2 17.34	41.139.0350.2
Cr ₂ O ₃ NiO Total Mg#	1.32 100.85 92.01	1.6 101 92.	1.05 65	1.60 100.42 93.34	30.92 0.3 99.93 72.09	0.41 100.92 89.98	0.57 100.19 90.15	0.95 9 100. 90.9	07 9	0.88 99.32 92.45	12 99 79	.09 .21 .85	0.18 100. 90.8	0.7 92 99 4 90	75 .00 .16	0.85 99.52 91.08	1.52 99.77 91.78	35.75 0.33 99.65 73.04	0.38 100.74 90.84
	GP28			GP2	9				GP	34					G	P49			
	Opx	CpxI	Sp	Ol	Op	x C	pxI S	Sp	Ol	(Opx	Cj	pxI	Sp	0	01	Opx	CpxI	Sp
SiO ₂ TiO ₂ Al ₂ O ₃ FeO MnO	55.78 3.37 5.36	52.36 4.33 2.12	36.73 13.4	41.1 9.37	7 55. 2.9 6.2	51 52 0. 6 4. 9 2.	2 21 19 4 39 1	41.6 13.15	40.7 8.82 0.1	74 : 2 : 1	55.00 4.23 5.41	51 0. 5. 2.	79 18 66 41	49.03 10.51	4 9.	1.12 .51	55.17 3.35 5.89	52.29 0.2 3.57 2.34	37.33 13.14
MgO CaO Na ₂ O Cr ₂ O ₃ NiO	34.34 0.94 0.66	16.17 21.02 1.4 1.59	17.54 32.41	. 49.9 0.32	6 34. 0.6 0.5	55 10 9 22 0 5 1	6.72 1 2.16 89 23 2	17.88 27.66	49.3 0.5	33 (33.81).92).54	16 20 1. 1.	5.46).67 3 3	19.6 20.66	4 <u>9</u> 0.	9.73	34.18 0.87 0.64	16.88 22.65 0.71 1.07	17.04 32.23
Total Mg#	100.45 91.95	98.99 93.15	100.0	90.4	82 100 8 90.).55 99 73 92	9.79 1 2.58 7	100.29 70.80	99.: 90.8	50 9 89 9	99.91 91.76	99 92	0.77 2.41	99.80 76.88	10 90	00.74 0.31	100.10 91.19	99.71 92.79	99.74 69.81

CpxI primary clinopyroxene, CpxII pyroxene crystallised in the glass, p porphiroclast, n neoblast crystal





2. LREE-enriched clinopyroxenes ("metasomatized") with $1 < (Ce/Yb)_N < 19$ (Fig. 5b, f). They show a wider range of heavy REE (HREE) contents $(Yb_N = 1.2-11.8)$. The chondrite normalized REE patterns vary from convex-upward (GP10, BR213 and BR218*) to "spoon shaped", that is enriched in LREE relative to middle REE (GP6, BR218; Fig. 5b, f). The primitive normalized diagrams for these clinopyroxenes show small or not positive anomalies of U and slight to marked negative anomalies of Nb (but not for Ta), Zr and/or Hf, and Ti. In this group, sample GP5 differs from the other clinopyroxenes for the relatively flat distribution of Th to Lu, with positive U anomaly (Fig. 5). CpxII, large enough to be analysed, has been found in one sample (BR213) and shows a trace element pattern similar to that of CpxI (Fig. 5).

Spinel

Spinel displays wide compositional variability (Table 3) with $Cr\# [(Cr/Cr + Al + Fe^{3+})\times 100]$ gradually increasing from 12 to 60 from lherzolites to harzburgites. The negative correlation between Mg# and Cr# (Fig. 6) suggests that partial melting strongly controls the spinel composition. In composite xenoliths, spinel shows relatively low Mg# and Cr# which is indicative of iron enrichment, a trend commonly associated with metasomatic effects (Dawson 1984).

Amphibole

Major and trace element contents of euhedral pargasite from the amphibole vein in BR218 are given in Tables 3 and 4. With respect to amphibole occurring in peridotite xenoliths from McMurdo Volcanic Group (Hornig et al. 1991; Beccaluva et al. 1991; Gamble et al. 1988; Coltorti et al. 2004), BR218 pargasite has lower Ti and alkali contents and higher Al₂O₃ values.

Chondrite normalized trace element patterns display positive Nb and Sr anomalies and a slight negative Zr anomaly. The convex-upward LREE-enriched patterns resemble those of megacrysts formed as a liquidus phase from alkaline magmas (Irving and Frey 1984; Fig. 7).

Glass

Glass patches are practically ubiquitous: at Green Point they are Si-rich and alkaline with Phonolite– Nepheline Trachyte compositions (Table 5). Glass blebs in BR213 are subalkaline and Quartz-normative and have Latite composition. Glasses show unusually high Mg# independently from their silica contents (0.82 in alkaline glasses and 0.79 in sub-alkaline ones). In composite xenoliths, glass compositions range from Mugearite to Benmoreite and show lower Mg# (from 0.69 to 0.53).

Both alkaline and subalkaline glasses are enriched in LREE but differ for MREE and HREE contents; the REE_N patterns are fractionated in subalkaline glasses



and almost flat in alkaline ones (Fig. 8; Table. 6). Alkaline and subalkaline PM-normalized diagrams are both characterized by positive Rb and Zr anomalies.

Glass in amphibole vein of BR218 sample is enriched in Nb and exhibits a trace element pattern similar to the nephelinite SAX20, the Greene Point host lava, and to

	GP35 ^a rim	190 1.55 65.08 4.61 18.64 4.12	$\begin{array}{c} 22.32\\ 4.4\\ 10.67\\ 1.43\\ 6.31\\ 0.42\\ 0.42\\ 0.16\\ 0.16\end{array}$	$\begin{array}{c} 0.99\\ 0.61\\ 0.69\\ 0.69\\ 0.69\\ 0.52\\ 0.52\\ 0.52\\ 0.34\\ 0.1\end{array}$	GP22 rim	$\begin{array}{c} 51.53\\ 0.36\\ 0.117\\ 0.117\\ 0.29\\ 0.23\\ 0.23\\ 0.019\\ 0.003\\ 0.006\\ 0.003\\$
	GP35 ^a core	181 1.28 28.02 2.52 7.11 2.03	7.03 7.03 7.03 7.03 7.03 7.03 7.03 7.03 0.50 0.17 0.17 0.17 0.17 0.17	0.45 0.08 0.03 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.05 0.08 0.08 0.08 0.08 0.08 0.03 0.05	GP22 Core	51.56 0.28 0.089 0.145 0.22 0.22 0.038 0.008 0.037 0.016
	GP34 rim	15.91 12.16 8.23	$\begin{array}{c} 0.24 \\ 0.77 \\ 0.63 \\ 0.32 \\ 0.23 \end{array}$	$\begin{array}{c} 0.43\\ 1.38\\ 0.20\\ 1.02\\ 0.14\\ 0.51\end{array}$	GP14 ^b core	0.138 0.95 1.57 0.015 0.015 0.012
	GP34 core	12.49 11.22 6.51	$\begin{array}{c} 0.32\\ 0.97\\ 1.20\\ 0.70\\ 0.33\end{array}$	2.01 0.41 1.00 0.17 1.20 0.19 0.34 0.34 0.01	GP10 ^b core	$\begin{array}{c} 0.34 \\ 1.1 \\ 2.1 \\ 0.078 \\ 0.063 \\ 0.036 \end{array}$
	GP32 ^a rim	255 289 7.31 25.05 4.01	$\begin{array}{c} 0.31\\ 13.52\\ 38.04\\ 15.86\\ 15.86\\ 1.11\\ 1.11\\ 2.22\\ 0.32\end{array}$	$\begin{array}{c} 1.64 \\ 0.30 \\ 0.71 \\ 0.09 \\ 0.83 \\ 0.83 \\ 0.83 \\ 0.72 \\ 0.42 \\ 0.72 \\ 0.72 \\ 0.72 \end{array}$	GP6 rim	$\begin{array}{c} 60.79\\ 0.102\\ 0.102\\ 0.11\\ 0.56\\ 0.17\\ 0.56\\ 0.17\\ 0.27\\ 0.023\\ 0.1\\ 0.023\\ 0.1\\ 0.023\end{array}$
	GP32 ^a core	202 155 5.11 16.36 3.59	$\begin{array}{c} 0.16\\7.96\\1.8.40\\1.85\\7.07\\1.32\\0.53\\0.53\\0.53\\0.21\end{array}$	$\begin{array}{c} 0.93\\ 0.22\\ 0.58\\ 0.58\\ 0.54\\ 0.11\\ 0.11\\ 0.23\\ 0.61\\ 0.22\end{array}$	GP6 core	61.74 0.071 0.097 0.0497 0.041 0.041 0.041 0.038 0.007 0.007
	GP31 rim	211 12.00 8.74 3.75 0.13	$\begin{array}{c} 0.26\\ 0.97\\ 0.19\\ 0.25\\ 0.25\\ 0.25\end{array}$	$\begin{array}{c} 1.97 \\ 0.34 \\ 0.92 \\ 0.15 \\ 0.15 \\ 0.1 \\ 0.1 \\ 0.0 \\ 0.01 \\ 0.01 \end{array}$	GP5 rim	68.36 0.029 0.29 0.010 0.012 0.012 0.012 0.012 0.002 0.002
	GP31 Core	$213 \\ 0.15 \\ 0.15 \\ 7.29 \\ 3.90 \\ 0.10 $	$\begin{array}{c} 0.45\\ 0.25\\ 0.26\\ 0.29\\ 0.29\\ 0.29\\ 0.20\\ 0.20\\ \end{array}$	$\begin{array}{c} 1.75 \\ 0.35 \\ 0.16 \\ 0.15 \\ 1.03 \\ 0.14 \\ 0.31 \\ 0.01 \end{array}$	Opx GP5 core	$\begin{array}{c} 72.81\\ 0.63\\ 8.96\\ 0.47\\ 0.66\\ 0.025\\ 5.42\\ 5.42\\ 5.42\\ 5.42\\ 0.025\\ 5.42\\ 0.025\\ 0.45\\ 0.07\\ 0.031\end{array}$
	GP29	57.67 6.25 11.23	$\begin{array}{c} 0.05\\ 0.82\\ 1.64\\ 0.68\\ 0.34\\ 1.02\\ 0.19\end{array}$	$\begin{array}{c} 1.08 \\ 0.22 \\ 0.63 \\ 0.08 \\ 0.28 \\ 0.08 \\ 0.29 \end{array}$	BR219 ⁶ rim	253 69.71 9.42 19.84 0.56 5.98 5.98 1.74 1.51 1.51 0.66
	GP29	56.54 5.79 10.24	$\begin{array}{c} 0.09\\ 0.82\\ 1.78\\ 0.81\\ 0.30\\ 0.97\\ 0.17\end{array}$	$\begin{array}{c} 1.19\\ 0.24\\ 0.62\\ 0.08\\ 0.47\\ 0.07\\ 0.42\end{array}$	BR219 ^a core	258 61.93 7.98 9.37 0.72 0.72 1.24 1.24 1.24 0.44
	GP28 rim	41.59 9.57 19.93	$\begin{array}{c} 0.32\\ 1.46\\ 2.50\\ 0.90\\ 0.40\\ 1.10\end{array}$	$\begin{array}{c} 1.60\\ 0.38\\ 1.01\\ 0.14\\ 0.90\\ 0.90\\ 0.12\\ 0.36\\ 0.36\end{array}$	218 ^{a,b,c}	221 26 26 21 21 21 21 21 22
	GP28 core	42.28 9.24 20.07	$\begin{array}{c} 0.32 \\ 1.52 \\ 1.52 \\ 2.42 \\ 0.91 \\ 0.40 \\ 1.35 \end{array}$	$\begin{array}{c} 1.33\\ 0.36\\ 0.36\\ 0.80\\ 0.14\\ 0.80\\ 0.09\\ 0.50\\ 0.02\\$	^a BR2	$\begin{array}{cccc} & 138 \\ 138 \\ 84. \\ 84. \\ 0. \\ 24. \\ 8. \\ 0. \\ 0. \\ 1. \\ 1. \\ 1. \\ 1. \\ 1. \\ 1$
	GP25 ^a rim	$\begin{array}{c} 155\\ 0.11\\ 2.318\\ 1.35\\ 2.90\\ 0.25\end{array}$	$\begin{array}{c} 0.26\\ 0.34\\ 0.34\\ 0.41\\ 0.12\\ 0.03\\ 0.03\end{array}$	$\begin{array}{c} 0.32\\ 0.06\\ 0.16\\ 0.03\\ 0.02\\ 0.02\\ 0.03\\ 0.03\\ 0.03\\ 0.03\end{array}$	BR218 rim	228 5.74 5.74 1.03 1.03 0.79 0.73 0.77 0.77 0.77 0.77 0.77 0.77 0.77
	GP25 ^a Core	$159 \\ 21.79 \\ 1.17 \\ 2.79 \\ 0.26$	$\begin{array}{c} 0.21\\ 0.75\\ 0.31\\ 1.48\\ 0.41\\ 0.11\\ 0.05\end{array}$	$\begin{array}{c} 0.26\\ 0.06\\ 0.11\\ 0.12\\ 0.12\\ 0.03\\ 0.03\\ 0.05\\ 0.01\end{array}$	BR218 ^a core	258 0.15 61.56 61.56 1.09 1.09 5.79 10.95 0.28 0.28 0.28 0.36 0.36
	GP22 ^a rim	200 164 5.92 5.45	$\begin{array}{c} 0.08\\ 6.87\\ 1.43\\ 5.21\\ 0.37\\ 0.37\\ 0.09\\$	$\begin{array}{c} 0.39\\ 0.09\\ 0.02\\ 0.02\\ 0.01\\ 0.04\\ 0.17\\ 0.43\\ 0.52\\ 0.52\\ 0.12\end{array}$	3R213 ^{a,b} tore	22 22.75 22.75 20.58 17.36 5.02 1.63
	GP22 ^a core	$189 \\ 0.09 \\ 171 \\ 1.91 \\ 5.48 \\ 5.31$	$\begin{array}{c} 0.76\\ 6.76\\ 1.347\\ 1.44\\ 0.35\\ 0.35\\ 0.07\end{array}$	$\begin{array}{c} 0.42\\ 0.07\\ 0.17\\ 0.02\\ 0.21\\ 0.03\\ 0.15\\ 0.58\\ 0.15\\ 0.15\end{array}$	GP49 1 rim c	33.09 1 6.94 1 12.20 1 1.46 1.46 1.46 0.75 0.75
	GP14 ^b core	17.16 15.36 12.04 0.11	$\begin{array}{c} 0.14 \\ 0.22 \\ 0.95 \\ 1.99 \\ 1.20 \\ 2.10 \end{array}$	2.48 1.53 1.59	GP49 core	33.82 7.99 13.61 0.24 0.32 1.50 0.73 0.73
	P10 ^{a,b} sre	79 .0.56 81.85 2.68	0.08 5.75 4.84 4.84 2.15 3.79 2.87 2.87	2.26 0.82 1.14 0.94 0.24	GP47 rim	231 15.72 13.88 13.88 13.88 13.88 0.21 0.21 0.21 0.28 0.28 0.28 0.28 0.28 0.28 0.28 0.28
	P6 ^a G n cc	3 0.54 5.63 17 1.73 1 5.27 8 1.11	2.43 1.37 0.55 0.16 0.16 0.37 0.06	$\begin{array}{c} 0.37\\ 0.06\\ 0.02\\ 0.03\\ 0.03\\ 0.18\\ 0.05\\ 0.06\\ 0.04\\ 0.06\end{array}$	GP47 core	241 17.47 11.481 12.79 0.22 0.08 0.01 0.11 0.01 1.35 0.85 0.85 0.05 0.05 0.05 0.061
	5 ^a G	18 61 2 20 13 62	02 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0	14 03 04 06 08 08 04 08	GP40 rim	245 9.70 0.09 0.09 0.09 0.09 0.20 0.21 0.17
	GP(core	9 178 9 19. 9 1. 7 0.	00000000000000000000000000000000000000	00000000000	GP40 Sore	209 0.22 9.8 0.51 0.02 0.02 0.14 0.18
	GP5 ^a rim	245 0.0 17.1 5.4 8.5 8.5 0.7	0.7 0.7 0.7 0.7 0.7 0.7 0.7	$\begin{array}{c} 0.9\\ 0.2\\ 0.1\\ 0.1\\ 0.1\\ 0.1\\ 0.1\\ 0.1\\ 0.1\\ 0.1$	3P36 e im	1.56 0.03 5.65 5.65 5.65 5.65 0.18 0.71 0.71 0.46 0.46
Cpx	GP5 ^a Core	$\begin{array}{c} 234 \\ 0.12 \\ 0.12 \\ 14.94 \\ 4.82 \\ 7.49 \\ 0.55 \\ 0.55 \\ 0.55 \end{array}$	0.29 0.05 0.05 0.12 0.12 0.12 0.12 0.12 0.12 0.12	$\begin{array}{c} 0.77\\ 0.18\\ 0.56\\ 0.09\\ 0.09\\ 0.020\\ 0.020\\ 0.09\\ 0.05\end{array}$	Cpx GP36 (Core r	12.77 10.29 6.06 0.77 0.77 0.65 0.29
		Sr Sr Sr	Tb Tb Tb Tb Tb Tb Tb Tb Tb Tb Tb Tb Tb T	U HH HU HU DY U HO DY U DY		Sm Sm Sm Sm Eu

Table 4 LAM ICP-MS analyses (ppm) of clynopyroxene, orthopyroxene and amphibole separates

Table 4 (Contd.)

	GP22 rim	$\begin{array}{c} 0.003\\ 0.013\\ 0.014\\ 0.005\\ 0.005\\ 0.005\\ 0.008\\ 0.$		sR218 ^b im	07 3	95 32	60	5	16.5 45.8	36.3	8.6 7,7	2.12 8.2	6.7	5 81 81	10.7	2.93				
	GP22 Core	$\begin{array}{c} 0.001\\ 0.007\\ 0.012\\ 0.002\\ 0.002\\ 0.002\\ 0.002\\ 0.002\\ 0.003\\ 0.003\\ \end{array}$	Чdu	218 ^b B re ri	4	∞		ۍ ۱	6.9	5.2	.6	8.2	6.6		j.	2.88				
	GP14 ^b core	0.051 0.099 0.13 0.21 0.04 0.04	An	19 BR Co	416 3	81 895 77 32	66 163 06 163	316	17 23 46	018 19 36	0, (31	016 31	10 10	, 043	36	25	5	04	
	iP10 ^b ore	.07 .11 .096 .034 .031		9 BR2 rim	209	0.1	ivi d		0.0	8 1 5 0.0	~ ~	5. .0	00	~ ~ ~	20 0.0	~ ~	0	0	38 0.	
	3P6 C im c	0.017 0 0.005 0 0.017 0 0.017 0 0.016 0 0.006 0 0.000 0 0 0.001 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0 0		BR21 core	166	1.0]	2.98		0.13	0.02	0.08	0.0	0.0	0.0	0.0	0.28		0.1	0.0	
	GP6 (sore r	0.04 0 0.002 0 0.024 0 0.03 0 0.03 0 0.002 0 0.002 0 0.001 0		BR218 rim	126	2.53 0.8	0.95		$0.16 \\ 0.35$	0.042 0.17	0.085		0.022	0.058	0.016	0.19	000	20.0 0.09	0.00	
	GP5 6	$\begin{array}{c} 0.026 \\ 0.003 \\ 0.013 \\ 0.008 \\ 0.024 \\ 0.002 \\ 0.002 \\ 0.005 \\ 0.007 \\ 0.007 \\ 0.001 \\ 0.000 \\ 0.0001 \\ 0.0000 \\ 0.000 \\ 0.000 \\ 0.000 $		R218 ore	16	2.01 0.84	0.2		$0.13 \\ 0.24$	0.015 0.055	50.0	0.01	$0.008 \\ 0.17$	0.039	0.024	0.26	0.016	0.049	0.012	
Opx	GP5 core	$\begin{array}{c} 0.07\\ 0.011\\ 0.065\\ 0.075\\ 0.075\\ 0.013\\ 0.036\\ 0.036\\ 0.036\\ 0.036\\ 0.036\\ 0.036\\ 0.033\\ 0.0$		P49 B cc	38 1	$0.11 \\ 0.63$	1.43		0.02	0.02			0.08	0.03	0.02	0.18	0.05			
	R219 ^a m	268 268 268 268 268 268 268 268 268 268		P42 G m	9.42 10).58	0.63	0.014).003).003).026	007	0.026	0.012	0.020	0.019	0.136	0.018			
	219 ^a B e ri	-0-0-0000000		42 G e ri	32 69	67 (012	U	087 (0059 (015 (010	000 (000 (000 (000 (000 (000 (000 (000	023 (016 (149 (029 (034 (001		
	o,c BR core	0.130		40 GF	06 76. 019	024 33 0.	0.0	4	0.3030.	00	004	00	003 0. 027 0.	012 0.	.0 000	107 0. 014 0.	011 0.	0. 0.		
	8R218 ^{a,}	.29 .69 .9 .14		0 GP rim	2 65.0 29 0.0	0.0	0.0	36 0.	0.0		0.0	194	5. 0. 0. 0.	15 0.0	0.0	0.0	0.0	4 0.0	01	
	218 ^a E	00 1 1 7 3		GP4 core	68.72 0.02	0.38	0.0(0.0				0.0	0.0	0.0	0.0	0.11	0.0	л. С	0.0(
	^a BR2 rim	$\begin{array}{c} 0.57\\ 0.14\\ 0.057\\ 0.025\\ 0.13\\ 0.04\\ 0.03\\ 0.0$		GP36	88.41	0.84	0.40		0.01				0.07	0.02	60.0	0.13	0.04	0.00		
	BR218 core	$\begin{array}{c} 0.42\\ 0.14\\ 1.27\\ 0.29\\ 0.93\\ 0.05\\ 0.35\\ 0.03\\ 0.08\\$		GP35 rim	64.39 0.36	5.37 0.221	1.58	3.95	0.43	0.081	0.03	0.064	0.029	0.008	0.002	0.028	0.045	0:059 0:059	0.025	
	213 ^{a,b} re	5 0 3 0		GP35 sore	51.32 0.51	6.83 0.26	1.87 0.78	4,48	0.48	$0.101 \\ 0.34$	0.051	0.074	$0.002 \\ 0.034$	0.005	0.004	0.061	0.033	0.065	0.021	
	49 BF	8 5.5 9 4.7 9 4.7 1 1		SP32 (im (6.98 ($0.35 \\ 0.305$	$0.71 \\ 0.082$	0.021	0.013	0.009 0.08	0.036	0	$0.008 \\ 0.034$	0.008	0.003	0.065	0.005	0.002 0.002	0.002	
	49 GF e rin	0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0		P32 C	3.26 5	0.258 0.319	0.67		0.013 0.058	0.010 0.052	0.025	0.061	0.034 0.034	0.011	0.013	0.051	0.005	0.003 0.003	0.001	
	47 GP cor	2.1.0 0.0 0		31 G cc	.9	2.0 01		36			0	60.	.066	0.49	.042	.28			0	
	7 GP	$\begin{array}{c} 1.88\\ 0.35\\ 0.27\\ 0.27\\ 0.26\\ 0.26\\ 0.26\\ 0.00\\$		GP rim	121	4 0	6	C	2	61	87	Ð	18 7 0	33 0	1 25 0	7 55 0	35	17		
) GP4 core	$\begin{array}{c} 2.38\\ 0.37\\ 0.61\\ 1.82\\ 0.64\\ 0.25\\ 0.25\\ 0.01\\ 0.01\\ 0.01\\ \end{array}$		GP31 core	105	0.8	0.2		0.0	0.0	0.0	4	0.0	0.0	0.0	0.0	0.0	0.0		
	GP4($\begin{array}{c} 0.81\\ 0.26\\ 1.69\\ 1.69\\ 0.46\\ 1.54\\ 1.54\\ 0.1\\ 0.1\\ 0.1\\ 0.1\\ 0.1\\ 0.0\end{array}$		GP29	99.04	$0.11 \\ 0.47$	0.74		0.00				0.03	0.01	00.0	0.0				
	GP40 core	$\begin{array}{c} 0.96\\ 0.18\\ 0.18\\ 1.49\\ 0.21\\ 1.31\\ 0.17\\ 0.01\\ 0.01\\ 0.01\end{array}$		GP28	87.04	0.54	1.05						0.04	0.07	10.0	0.18				0
	GP36 rim	$\begin{array}{c} 1.09\\ 0.2\\ 1.29\\ 0.32\\ 0.14\\ 0.14\\ 0.14\\ 0.22\\ 0.22\end{array}$		GP25 rim	47.82	0.04 0.068	0.198		$0.004 \\ 0.004$	0.005 0.006	2000	cuu.u	0.002	0.006	0.003	0.011	0.011		0.001	oo poo:
Срх	GP36 Core	$\begin{array}{c} 1.20\\ 0.24\\ 1.42\\ 1.42\\ 1.08\\ 0.16\\ 0.16\\ 0.13\\ 0.35\\ 0.35\\ 0.01\end{array}$	Opx	GP25 Core	51.21	$0.067 \\ 0.102$	$0.34 \\ 0.032$	0.08	0.009	0.012	0.008	0.064	0.015	0.004	0.0010	0.036	0.010	0.001	0.065	1000000
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^aMetasomatised sample ^bTrace elements data determined by SIMS ^cClinopyroxene near amphibole vein

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Fig. 6 Composition of spinel in terms of Mg# $[(Mg/Mg+Fe^{2+})\times100]$ and Cr# $[(Cr/Cr+Al+Fe^{3+})\times100]$. The plot illustrates a negative correlation between the two parameters. More depleted xenoliths have higher Cr# suggesting a primary composition control by partial melting



Fig. 7 Primitive mantle (PM; McDonough and Sun 1995) normalized incompatible trace elements distribution of amphibole. The composition of amphibole megacryst (279 MCX), Beccaluva et al. (1991) and disseminated amphibole (amph-D) and amphibole in vein (amph-V) from Coltorti et al. (2004) are also reported

glasses in a composite lherzolite described by Coltorti et al. (2004; Fig. 8).

Oxygen isotope composition

Olivine

Olivine in anhydrous spl-lherzolites exhibits constant Oisotope composition ($\delta^{18}O_{Ol} = 5.4 \pm 0.19\%$; Table 7), in the range of spinel-lherzolites on a worldwide scale (5.19±0.28;. Mattey et al. 1994). Only sample GP49 shows an anomalously high $\delta^{18}O_{Ol}$ (5.81‰). The $\delta^{18}O$ values are consistent with the homogenous Fo contents. Olivine grains from composite xenoliths show a variable O-isotope composition (5.0-5.8%); those from the amphibole-bearing sample (BR 218) have higher δ^{18} O values (5.7-5.8%) regardless of whether they were collected from domains adjacent to amphibole or from anhydrous domains. Olivines from samples BR213 and 219 show the lowest δ^{18} O values measured in this suite of xenoliths (5.0-5.1%).

Orthopyroxene

Enstatite in peridotite xenoliths has homogeneous δ^{18} O values (5.8±0.16‰; Table 7); the Δ^{18} O_{opx-ol} values (δ^{18} O_{enstatite}- δ^{18} O_{olivine}) are quite variable (Δ^{18} O_{ol-opx} 0.48±0.30‰) with respect to the 0.5‰ equilibrium fractionation at mantle conditions (Mattey et al. 1994). The highest δ^{18} O values (6.1‰) are found in amphibole-bearing sample.

Clinopyroxene

Clinopyroxenes have heterogeneous δ^{18} O values ranging from 5.0 to 5.7‰, the highest δ^{18} O values shown by cpx from amphibole bearing sample. Clinopyroxene-olivine fractionation ($\Delta^{18}O_{ol-cpx} - 0.07 \pm 0.19\%)$ indicates O-isotope disequilibrium (Fig. 9a). Among not-metasomatized samples, the $\delta^{18}O_{cpx}$ values inversely correlate with Mg# and Cr₂O₃ contents ($R^2 = 0.709$ and 0.595, respectively). Positive covariations are observed for the $\delta^{18}O_{cpx}$ versus TiO₂ (R²=0.727) and versus Y (R²=0.527) (Fig. 9e-f), whereas poor or no correlations are found for the metasomatic ones $(R^2=0.318)$ and 0.086, respectively). Poor covariation trends $(R^2 \le 0.277)$ are displayed by either group in the $\delta^{18}O_{cpx}$ versus SiO₂, (Fig 9d), and in the $\delta^{18}O_{cpx}$ versus Al₂O₃, versus La, and versus Ce.

Spinel

The measured oxygen isotope compositions of spinel are variable (3.2–4.9%; Table 7) and within the values reported for mantle spinel worldwide (e.g. Kyser et al. 1981; Chazot et al. 1997). Due to the paucity of samples (n=4), the $\delta^{18}O_{spl}$ correlations are poorly constrained; however, a remarkable positive covariation occurs between the $\delta^{18}O_{spl}$ values and spinel Cr# ($R^2=0.953$)

Discussion

Variations in modal composition, whole rock and mineral chemistry provide the evidence for heterogeneous upper mantle beneath northern Victoria Land which is due to the combined effects of partial melting and metasomatic enrichments.

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Fig. 8 Classification of glass included in Baker Rocks and Greene Point xenoliths after Le Bas et al. (1986). Discriminating curve of alkaline and subalkaline field is from Irvine and Baragar (1971). Chondrite normalized trace element patterns (a) and REE pattern (b) for glasses

Extraction of silicate melts, resulting from different degrees of partial melting, may account for: (1) decrease of clinopyroxene modal proportion; (2) whole rock and pyroxene depletion in "basaltic" components (e.g. Al, Na, REE), and (3) contemporaneous build up of refractory elements (e.g. MgO, Ni, Cr# of pyroxene and spinel; Table 2, 3). In the following discussion, we distinguished metasomatized versus not-metasomatized samples according to the occurrence of LREE-enrichment or LREE-depletion in clinopyroxene (Table 4).

Partial melting

Whole rock major element compositions define trends similar to those calculated for residuals of the primitive mantle melting (Niu 1997) in the spinel stability field (Fig. 2). Spinel Cr# of the xenoliths varies between 0.13 and 0.53 (Table 3) and constrains the extent of source melting to be 3-17% of perfect fractional melting in the spinel stability field (Hellebrand et al. 2001). For most samples, clinopyroxene HREE contents confirm that fractional melting occurred in the spinel stability field (e.g. Johnson et al. 1990; Norman 1998; Xu et al. 2000; Hellebrand et al. 2002). The degree of melting (*F*) varies between 3 and 18% (Fig. 10a). However, the fractionated HREE patterns of samples GP40, BR218 and GP6 likely indicate a garnet-bearing mantle residue. In order to produce the observed HREE abundances, an initial 5–20% of melting into the garnet stability field needs to be accounted for, then followed by

Table 6 SIMS analyses (ppm) of intergranular glasses

	GP 5	GP 14	BR 213	BR 213	BR 213	BR 218
v	36	8	244	234	220	218
Cr	4	17	62	56	65	148
Rb	154	210	103	116	98	15
Sr	274	63	397	455	472	861
Y	6.7	4.0	20	21	17	57
Zr	127	47	249	289	223	453
Nb	44	17	58	69	52	241
Cs	2.04	0.99	1.19	1.02	0.65	0.22
Ba	255	116	562	636	515	685
La	26	8.6	30.0	33.7	26.4	54.3
Ce	41	14.2	57.6	66.5	52.7	130.8
Nd	14	4.7	25.9	29.2	23.1	83.0
Sm	2.24	0.52	4.5	5.0	4.3	16.7
Eu	0.32	0.00	1.15	1.41	0.96	4.0
Gd	1.79	0.56	4.5	4.7	4.2	15.5
Dy	1.29	0.73	3.8	4.3	3.2	11.8
Er	0.73	0.45	2.00	1.81	1.63	5.3
Yb	0.69	0.49	2.18	2.09	1.72	5.1

Table 7 Oxygen isotopic composition of minerals from Greene Point and Baker Rocks spinel peridotites and standard deviation of the mean (SEM)

Sample	δ^{18} O Opx (‰ SMOW)	SEM	δ ¹⁸ O Ol (‰ SMOW)	SEM	δ ¹⁸ Ο Cpx (‰ SMOW)	SEM	δ ¹⁸ O Spl (‰ SMOW)	SEM
GP5	5.90	0.10	5.47	0.07	5.32	0.13		
GP6	5.72	0.03	5.12	0.13	4.96	0.19		
GP10	5.79	0.04	5.30	0.09	5.37	0.02	3.19	0.11
GP14	5.70	0.09	5.55	0.04	5.62	0.07		
GP22	5.77	0.08	5.39	0.10	4.99	0.12		
GP25	6.04	0.12	5.25	0.07	5.13	0.15		
GP28	5.61	0.03	5.09	0.03	5.24	0.13		
GP29	5.56	0.10	5.35	0.15	5.34	0.02		
GP31	6.00	0.06	5.36	0.13	5.15	0.04		
GP34	5.90	0.07	5.40	0.01	5.42	0.05		
GP36	5.89	0.01	5.42	0.02	5.26	0.06		
GP42	5.82	0.06	5.58	0.07	5.53	0.07	4.88	0.04
GP49	5.71	0.16	5.81	0.09	5.35	0.11		
BR213	6.00	0.18	4.80	0.08	5.10	0.13		
BR218	5.96	0.01	5.79	0.06	5.62	0.07	4.43	0.09
BR218 ^a	6.14		5.67	0.05	5.71			
BR219	5.87	0.16	4.96	0.01	5.12	0.20	3.79	0.21

BR218 Amphibole δ^{18} O = 5.16% and GP Lava Host (SAX20) δ^{18} O = 5.85%

^aMinerals near to the amphibole vein in composite xenolith

further 3-15% of melting in the spinel stability field (Fig. 10b).

The not-metasomatized samples are characterized by significant linear covariations of $\delta^{18}O_{cpx}$ versus cpx modal abundance ($R^2 = 0.761$) and versus $F(R^2 = 0.948;$ Fig. 11a, b), whereas poor correlations are defined by the $\delta^{18}O_{cpx}$ values of the metasomatized samples ($R^2 \le 0.433$). The $\delta^{18}O_{spl}$ values define a noticeable covariation trend ($R^2 = 0.953$) with spl–Cr# (and therefore F, $R^2 = 0.999$; Fig. 11c), regardless of the metasomatic overprint. These relationships are consistent with the effects caused by partial melting. If partial melts derived from spl-lherzolite are more siliceous than the residual assemblage (Hirschmann et al. 1998), the ¹⁸O-isotope fractionation in the more polymerized phase (e.g. Garlik 1966; Appora et al. 2003) is consistent with a decrease in δ^{18} O of cpx and depends on the extent of partial melting. In contrast, the $\delta^{18}O_{opx}$ and $\delta^{18}O_{ol}$ values show only slight variations likely resulting from the minor contribution of olivine and orthopyroxene to partial melting and their peritectic relationships. Moreover, the larger grain size of these phases in mantle xenoliths accounts for a difficult diffusion-assisted re-equilibration. Nevertheless, oxygen isotope disequilibrium fractionation has been documented for olivine and orthopyroxene in peridotites with no metasomatic (open-system) overprint, and this in spite of the large oxygen buffer of the mantle (e.g. Harmon et al. 1987; Deines and Haggerty 2000; Eiler 2001; Zhang et al. 2000).

If the $\delta^{18}O_{cpx}$ values are controlled by the liquid/solid partition coefficients at the melting temperature and vary as a function of T, we may assume a major cpx control on the isotopic composition of the melt. Using the approximation:

$$1000 \ln \alpha_{melt-cpx} = \delta^{18} O_{melt} - \delta^{18} O_{cpx} = \Delta^{18} O_{melt-cpx}$$

(e.g O'Neil 1986), simple mass balance considerations imply that melting processes is described by the equation:

$$\delta^{18} \mathbf{O}_{\mathrm{cpx}}^{\mathrm{initial}} = F \times \delta^{18} \mathbf{O}_{\mathrm{melt}} + (1 - F) \times \delta^{18} \mathbf{O}_{\mathrm{cpx}}$$

(*F* is the fraction of partial melting derived from the nomodal fractional melting equation of Hellebrand et al. 2002).

After little manipulation:

$$\delta^{18} O_{cpx} = \delta^{18} O_{cpx}^{initial} - F \Delta_{melt-cpx}.$$

In principle, this equation could explain the linear dependence of $\delta^{18}O_{cpx}$ from *F* (Fig. 11b). The intercept and the slope of the linear array also provide estimates for the peridotite ¹⁸O- starting composition and for the $\Delta^{18}O_{melt-cpx}$, respectively. The calculated $\delta^{18}O_{cpx}^{initial}$ approximates 5.68%, in the range of peridotite cpx on world-wide scale (Mattey et al. 1994). In contrast, the $\Delta^{18}O_{melt-cpx}$ derived from the slope (~2.95) is inconsistent with equilibrium fractionation between peridotite residuum and partial melts ($\Delta_{melt-cpx} = 0.2$; Zhao and Zheng 2003) and ultimately unrealistic for ¹⁸O-isotope fractionation at mantle conditions.

Rearrangement and differentiation of the above equation results into:

$$\frac{\partial (\delta^{18} \mathbf{O}_{cpx}^{initial} - \delta^{18} \mathbf{O}_{cpx})}{\partial F} = \Delta_{melt-cpx} = \frac{-\partial (\delta^{18} \mathbf{O}_{cpx})}{\partial F}$$

In this equation, F is imposed by the mass balance and the $\Delta_{\text{melt-cpx}}$ values strongly depend on the rate of change in $\delta^{18}O_{\text{cpx}}$ with F. This condition is achieved only if the melt is subtracted from the matrix after its Fig. 9 Oxygen isotopic composition for olivine, orthopyroxene and clinopyroxene for NVL spinel peridotites. The line in (a) and (b) corresponds to $\Delta^{18}O = 0$. *Open circle* represents not metasomatized sample; *filled symbols* represent metasomatized sample. c and d plots show the negative correlation of Mg# and SiO₂ content with the $\delta^{18}O_{cpx}$ values. On the contrary the $\delta^{18}O_{cpx}$ is positively correlated with TiO₂ and Y contents, (e) and (f)



formation, thereby implying that the dependence of $\delta^{18}O_{cpx}$ from *F* may be described only by a fractional melting model. Present data indicate that the correlations among the $\delta^{18}O_{cpx}$ and the parameters related to the extent of fractional melting hold in the not-metasomatized samples; the $\delta^{18}O_{spl}$ versus $Cr_{spl}\#$ suggest that these correlations may also survive in samples which experienced low degrees (cryptic) of mantle metasomatism.

Mantle metasomatism

First order evidence for metasomatic events are represented by the high FeO, TiO_2 and incompatible trace element contents of the whole rock and pyroxenes. REE_N clinopyroxene patterns can be interpreted in terms of a chromatographic fractionation during reactive porous flow (Navon and Stolper 1987; Bodinier et al. 1990; Bedini et al. 1997; Vernières et al. 1997). In the chromatographic model, the behaviour of elements during melt percolation through the peridotite depends on mineral-melt partition coefficient (D^{rock/melt}). The interaction of the percolating melt with the peridotite assemblage causes the melt to be progressively enriched in LREE, producing a transient metasomatic zoning with different enrichment patterns as a function of the distance from the magma source (Ionov et al. 2002). Metasomatism can be referred to as "wall-rock" metasomatism, linked to the transport of melt in fractures (veins and dykes), and "diffuse" metasomatism related to percolation of small melt fractions along grain boundaries in the peridotite matrix. Peridotites (and cpx) affected by "wall-rock" metasomatism are generally characterized by Fe-enrichment and convex-upward or



Fig. 10 REE chondrite normalised patterns of NVL clinopyroxene compared with the results of fractional melting model (Johnson et al. 1990) in the spinel stability field (a) and a combination of initial melting in the garnet stability field followed by additional melting in the spinel stability field (b). Initial source composition, sources modes, melt modes and distribution coefficients are after Hellebrand et al. (2002). Chondrite normalization values taken from McDonough and Sun (1995)

flat REE patterns, whereas Fe-enrichment are absent and the REE patterns are highly fractionated (from U-shaped to steadly enriched patterns) where "diffuse" metasomatism has been active (Bodinier et al. 1990; Bedini et al. 1997; Ionov et al. 2002, Xu and Bodinier 2004).

Composite sample BR218 is typical in this respect: the clinopyroxene next to the veins are FeO-enriched, with convex-upward REE patterns; far (\sim 5 cm) from the vein, cpx REE pattern becomes convex-downward. This can be interpreted as a shift from "wall-rock" to "diffusive" metasomatism linked to different melt/rock ratios, as required in the chromatographic fractionation (Kelemen et al. 1995; Bedini et al. 1997; Xu and Bodinier 2004).

Textural relations, FeO enrichments and the REE pattern of BR218*clinopyroxene, indicate that its composition approached equilibrium with the metaso-

matic melt. Using crystal/melt partition coefficients (Skulski et al. 1994; Halliday et al. 1995) and trace element composition of BR218* cpx, it is possible to estimate the trace element concentrations of the melt in



Fig. 11 (a) Variation of oxygen isotope composition of clinopyroxene with its modal abundance in NVL spinel peridotites. The δ^{18} O values of clinopyroxene (b) and spinel (c) decrease with the degree of partial melting suggesting that partial melting have a primary control on the oxygen isotope composition of clinopyroxene

equilibrium with this cpx. The calculated liquid shows trace element pattern similar to that of the GP host lava (SAX20, Tab. 5), although with lower Nb and Ta contents, and more negative Ti anomaly (Fig. 12). The differences between calculated liquid and host GP lava can be due to the crystallization of amphibole in the vein, that likely depleted the residual melt in Nb, Ta and Ti. This melt percolated away from the vein and interacted with the peridotite assemblage producing a cpx (sample BR218) with negative Nb, Zr and Hf anomalies, fractionated LREE patterns and decoupled Nb-Ta contents (Ionov et al. 2002; Xu and Bodinier 2004). Similarly, a process of diffusive metasomatism might be responsible for variable incompatible elements enrichments of GP samples. These samples probably represent peridotite portions at various distance from the melt conduit (vein and/or dyke). From the composition of calculate equilibrium melts (GP10 and BR218*), the metasomatic agents for BR and GP xenoliths is inferred



Fig. 12 PM normalized incompatible patterns of calculated liquid in equilibrium with enriched clinopyroxene of xenoliths. The calculations are made using cpx/liquid distribution coefficient proposed by Skulski et al. (1994) and Halliday et al. (1995). The figure also reports the range of variation of some incompatible elements (*grey field*) of basic McMurdo Lavas (Rocchi et al. 2002)

to be similar in composition to the nephelinite GP host lava, although the high ascent rates and the resulting short scale diffusion-controlled effects prevent the host lavas itself to be the very metasomatic agent.

The metasomatic events were able to shift the $\delta^{18}O_{cpx}$ values only where the (metasomatic) fluid/rock ratio was high, as likely in the case of sample BR218* (vein).

Assuming (1) the δ^{18} O of the metasomatic agent to be similar to the δ^{18} O value of nephelinite sample SAX20 ($\delta^{18}O_{WR} = 5.85\%$); (2) the $\delta^{18}O_{cpx}^{initial} = 5.60$ (Mattey et al. 1994); and (3) using a $\Delta^{18}O_{cpx-Met}$ values ($\Delta^{18}O_{cpx-Met}$ = fractionation between the peridotite cpx and the metasomatic melt) of 0.15\% at 1100°C (Zhao and Zheng, 2003), it is possible to calculate the variation of the $\delta^{18}O_{cpx}$ value as a function of the mass ratio between the rock (*P*) and the metasomatic melt ($\frac{P}{Met}$) by the simple mass balance calculation:

$$\frac{P}{Met} = \frac{(\delta^{18} \mathbf{O}_{cpx}^{\text{final}} - \delta^{18} \mathbf{O}_{cpx}^{\text{initial}})}{\delta^{18} \mathbf{O}_{Met}^{\text{initial}} - (\delta^{18} \mathbf{O}_{cpx}^{\text{final}} - \Delta_{cpx-Met})}$$

On these bases, a maximum $\frac{P}{Met}$ ratio of 0.38 would be required to shift the $\delta^{18}O_{cpx}$ values of sample BR218 to 5.71%. This value is plausible and consistent with textural relationships between vein and peridotite in composite sample BR218.

Diffusion-based calculation (Crank 1975) provides time estimates for zoning to be achieved in crystals with core-rim REE variations (Fig.5; Table 4). The equation $x = (Dt)^{-2}$ approximates the time required for an element to travel the distance x in a time "t". For diffusion coefficients $D_{REE} \approx 10^{-21}$ m²/s (Van Orman et al. 2001), larger crystals ($\emptyset = 2$ mm) would be completely homogenized in a time span between 10 and 40 millions of years. In a similar time interval, oxygen isotope disequilibrium may survive among coarse-grained minerals at mantle conditions (e.g. Ryerson et al. 1989; Ingrin et al. 2001

Origin of intergranular glass patches

Late evidence of melting is also provided by the occurrence of alkaline glasses in the spongy rims of clinopyroxenes and by interstitial blebs of alkaline glass in reaction with spinel \pm clinopyroxene. Both these textural features result from xenolith disequilibrium melting during the transport into the host lava, due to rapid P–T variations (Francis 1976; Dawson 2002). The high glass concentrations of K, P, Rb and Ba indicate that amphibole and/or phlogopite were probably involved into the melting reactions (e.g. Hornig and Wörner 1991; Beccaluva et al. 1991; Coltorti et al. 2004); the absence of these hydrous phases in Greene Point xenoliths is likely due to their complete exhaustion during melting.

Sub-alkaline glasses found in sample BR213 were produced from the reaction between SiO_2 -poor melt and the depleted peridotite, this reaction produces a distinctive Si–Al-rich hybrid melt (Zinngrebe and Foley 1995; Shaw and Edgar 1997). Similar glasses, although showing lower Al_2O_3 contents, have been experimentally produced by orthopyroxene/basanite interaction (Shaw 1999). Higher Al_2O_3 contents in glasses of NVL xenoliths, and high Cr_2O_3 contents of associated Cpx II, likely account for the contribution of spinel to the glass-forming reaction.

Though somehow related to trace element patterns of McMurdo Volcanic Group, subalkaline and alkaline glass compositions do not match those of mafic lavas of MMVG, confirming that they are the product of disequilibrium reactions.

Thermal regime

Thermometric (two-pyroxene geothermometer; Wells 1977; Brey and Köhler 1990) and barometric estimates (olivine–clinopyroxene Ca-exchange reaction; Köhler and Brey 1990) show that coarse crystals of xenoliths were equilibrated in the spinel peridotite field (Table 8; Fig. 13) at T between $950 < T < 1050^{\circ}$ C and P between 1.0 < P < 1.5 GPa. The P–T data of primary phases of the peridotite assemblage (squares) are close to the hydrous peridotite solidus (H₂O < 0.4wt%; Falloon et al. 1988; Falloon et al. 1999).

Estimates of equilibrium T and P obtained on the crystal cores were compared with data from Mt. Melbourne xenoliths (Beccaluva et al. 1991) and with the data defined by geotherms typical of stationary and dynamic rifts (Chapman 1986).

The higher temperature array of our data defines a geotherm of about 80°C/GPa at P > 1 GPa, compatible with a dynamic rift geotherm. This implies a larger heat supply than previously thought on the basis of a 40°C/GPa geotherm (Berg et al. 1989). This new geotherm

 Table 8 Temperature and pressure estimes for peridotites from

 Greene Point and Baker Rocks

	T (°C), Wells (1977)	T (°C), Köhler and Brey, (1990)	P (GPa), Köhler and Brey (1990)
BR213	1027	1041	1.1
BR218	981	981	0.9
BR219	1061	1047	1.2
GP 1	970	948	1.0
GP 5	959	940	
GP 6	1043	1035	1.2
GP 10	1004	1016	1.4
GP 14	998	1033	1.4
GP 15			
GP 17	936	1039	1.1
GP 20	1045	1019	1.4
GP 21			
GP 22	1062	1062	1.4
GP 25			
GP 28	944	944	1.4
GP 29	941	941	1.5
GP 34	1007	959	1.4
GP 36	1028	1031	1.0
GP 49	925	925	0.9

meets the hydrous solidus of peridotite at about 2 GPa, running close to the hydrous peridotite solidus. This suggests that low degree partial melts had the possibility to migrate through the peridotite matrix.

In near-primary melts erupted in NVL, the estimated conditions for olivine crystallization are in the range of 1250–1350°C at pressures between 2.0 and 2.5 GPa (Orlando et al. 1997). These conditions likely correspond to the mantle potential temperature (Tp; Green et al. 2001); they also match the common Tp values of the asthenosphere and its thermal boundary layer (Anderson 1995; McKenzie and Bickle 1988). Starting from these P–T conditions and assuming an adiabatic coefficient of 30° C/GPa, the estimated mantle geotherm sets the conditions for the genesis of McMurdo melts at $4 \le P \le 5$ GPa (corresponding to depths greater than 120 km) and 1350 $\le T \le 1400^{\circ}$ C. These mantle potential



Fig. 13 Temperature (°C) and pressure (GPa) estimates from North Victoria Land mantle xenoliths. Squares represent two pyroxene thermometer (Brey and Kölher 1990) combined to barometer Ol-Cpx (Kölher and Brey 1990). Error bars are reported in the figure. McMurdo geotherm from Berg et al. (1989). Static and dynamic rift geotherms are from Chapman (1986). PLG, SP and GT:PT ranges of stability of plagioclase, spinel and garnet. Curves 1 and 2 are the hydrous (0.4 wt% H₂O) and anhydrous peridotite solidi (Falloon et al. 1988; Falloon et al. 1997; Green 1988). Light grey field represents the PT range of mantle xenoliths from NVL reported in (Beccaluva et al. 1991). Grav rectangle represents the PT range for the onset of olivine crystallization in near primary melts of NVL (Orlando et al. 1997). Heavy grey line represents the NVL mantle geotherm visually fitted from the high temperature envelop of our PT determinations: P > 1 Gpa, the gradient is 80°C/GPa. See text for discussion

temperatures are at least 200–300°C lower than estimates expected for a thermally buoyant plume (Shubert et al. 2001). Thus decompressional melting of fertile or re-enriched lithospheric mantle (Orlando et al. 2000; Rocchi et al. 2002) in a regime of transtensional tectonics (Salvini et al. 1997) turns out to be a more satisfactory explanation for the genesis of magma in NVL.

The frame emerging from trace elements and oxygen isotopic data is in agreement with the conclusions derived by Zipfel and Wörner (1992), reconstructing the thermal history of NVL lithospheric mantle. These authors, on the basis of diffusion profile of Ca in the mantle olivines, stated that the thermal state of the lithospheric during Cenozoic evolved from normal temperature gradient to a dynamic rift geotherm as a consequence of the onset of magmatic activity.

Conclusions

The chemical and isotopic data on mantle xenoliths from northern Victoria Land constrain the compositional evolution of the mantle underneath the Ross Sea region of Antarctica. Extraction of silicate melt via different degrees of partial melting affected the spinelperidotite mantle. Partial melting was followed by cryptic and modal metasomatic events, resulting in chemical and mineralogical heterogeneities of the peridotite. These metasomatic effects are characterized by Fe–Ti addition and variable LREE enrichments of the clinopyroxene. The occurrence of veins of fertile wehrlite or hornblendite is related to the infiltration of magmas at mantle depths. In addition, disequilibrium melting/ reactions during the ascent to surface produced intergranular high Mg# and Si-rich melts.

The δ^{18} O values indicate non-equilibrium oxygen isotope fractionation at mantle depths and the δ^{18} O_{cpx} and δ^{18} O_{spl} correlations with *F* are consistent with a process of fractional melting. Metasomatic overprint affected the ¹⁸O/¹⁶O ratios of the minerals only where the metasomatic-melt/rock ratios were high and likely occurred no more than 40 Ma bp.

Thermobarometric estimates reveal a progressive heating of the mantle in a thermal regime of dynamic rifting. A geotherm of 80°C/GPa is proposed for the conductive refractory lithosphere of NVL close to the hydrous solidus of peridotite. The mantle potential temperature for the origin of near primary melts crosses the geotherm at 1350–1400°C and 4–5 GPa (120– 150 km). This thermal regime is coherent with the hypothesis of decompressional melting induced by transtensive tectonic regime and conflicts with the presence of an active plume in the area of the Ross Sea.

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