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Analyses of shocked quartz at the global K-P boundary indicate an origin from a single, high-angle, oblique impact at Chicxulub

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Abstract

The precise cause and timing of the Cretaceous–Paleocene (K–P) mass extinction 65 Ma ago remains a matter of debate. Many advocate that the extinction was caused by a meteorite impact at Chicxulub, Mexico, and a number of potential kill-mechanisms have been proposed for this. Although we now have good constraints on the size of this impact and chemistry of the target rocks, estimates of its environmental consequences are hindered by a lack of knowledge about the obliquity of this impact. An oblique impact is likely to have been far more catastrophic than a sub-vertical one, because greater volumes of volatiles would have been released into the atmosphere. The principal purpose of this study was to characterize shocked quartz within distal K–P ejecta, to investigate whether the quartz distribution carried a signature of the direction and angle of impact. Our analyses show that the total number, maximum and average size of shocked quartz grains all decrease gradually with paleodistance from Chicxulub. We do not find particularly high abundances in Pacific sites relative to Atlantic and European sites, as has been previously reported, and the size-distribution around Chicxulub is relatively symmetric. Ejecta samples at any one site display features that are indicative of a wide range of shock pressures, but the mean degree of shock increases with paleodistance. These shock- and size-distributions are both consistent with the K–P layer having been formed by a single impact at Chicxulub. One site in the South Atlantic contains quartz indicating an anomalously high average shock degree, that may be indicative of an oblique impact with an uprange direction to the southeast $\pm 45^{\circ}$. The apparent continuous coverage of proximal ejecta in this quadrant of the crater, however, suggests a relatively high impact angle of $>45^\circ$. We conclude that some of the more extreme predictions of the environmental consequences of a low-angle impact at Chicxulub are probably not applicable.

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J. Morgan et al. / Earth and Planetary Science Letters 251 (2006) 264–279 265

1. Introduction

Alvarez et al. [\[1\]](#page-14-0) identified an extraterrestrial signature within the K–P boundary clay layer and, following a lengthy debate, this layer is now universally agreed to have been formed by a meteorite impact. The K–P layer also contains highly shocked terrestrial minerals (quartz, feldspar and zircon) that originated from rocks at the impact site, as well as impact-derived Ni-rich spinel [\[2](#page-14-0)–8]. After the Chicxulub impact crater was discovered [\[9\]](#page-14-0) numerous studies of this crater and the K–P boundary led to widespread agreement that Chicxulub was the site of a K–P boundary impact, with a few notable exceptions [\[10,11\].](#page-14-0) The cause of the K–P mass extinction is still contentious, and the precise role of the Chicxulub impact and Deccan volcanism remains a matter of some debate [e.g. [\[12,13\]\]](#page-14-0). The impact would have released dust and climatically active gasses into the atmosphere, and caused a period of extreme temperature change and darkness, but calculations of the nature, extent and duration of these changes differ [e.g. [\[14](#page-14-0)– [17\]](#page-14-0)]. Estimates of climatic effects are dependent, in part, upon knowledge of the energy of impact, the chemistry of the target rocks, and the obliquity of impact. An oblique impact would have been much more damaging than a sub-vertical one because the shock wave would be more focused within the near-surface, volatile-rich, sedimentary rocks, and this would cause the release of higher volumes of climatically active carbon- and sulfur-rich gasses [\[17](#page-14-0)–19]. The impact size and target chemistry are now reasonably well constrained [e.g. [\[9,20\]\]](#page-14-0), but the obliquity of impact is unknown.

The best indicator of direction of impact for craters on other planetary bodies is found in the asymmetric distribution of their proximal ejecta ([Fig. 1a](#page-2-0)). Highly oblique impacts $($ < 10 $^{\circ}$) produce elliptical craters and a butterfly ejecta pattern with ejecta-free zones in the up and downrange directions. For impact angles of greater than ∼15° near-circular craters are produced and, for oblique impact angles up to ∼45°, ejecta is missing in the uprange direction [\[21\]](#page-14-0). Chicxulub is roughly circular [\[9,20\],](#page-14-0) suggesting an impact angle of $>15^\circ$. Although asymmetries in the apparent crater structure have been used to estimate impact angle and direction [\[18,22\],](#page-14-0) the feature applied (offset of the central uplift) has been shown not to be diagnostic of impact direction for craters on Venus[\[23\]](#page-15-0). At Chicxulub much of the proximal ejecta is buried and, as the impact site was submerged continental shelf, is likely to have been re-distributed by tsunamis or erosion immediately or soon after impact.

During the excavation stage of cratering the so-called ejecta curtain material is ejected at particle velocities of

up to ∼2 kms-1 to form the proximal ejecta deposits [\[24,25\],](#page-15-0) and for Chicxulub these deposits should extend up to ∼400 km from the impact site. This ejection mechanism cannot explain the presence of shocked quartz in the global K–P layer as velocities of >9 kms-1 are required to eject quartz to the other side of the planet. The only plausible explanation is that these ejecta were accelerated to high speeds within the expanding vapor plume [\[25\].](#page-15-0) The expanding plume would consist of a mixture of vapor, melt and clastic minerals metamorphosed at different shock pressures. However, the precise kinematics of ejection are unknown and different assumptions about the velocity/mass distribution lead to quite different final distributions of ejecta [\[26,27\].](#page-15-0)

Observational, experimental and numerical modeling of oblique impacts all indicate that vapor plumes initially expand in the downrange direction [\[19,28,29\]](#page-14-0). Hence it appears plausible that the shocked quartz could be ejected in an asymmetric pattern from the impact site. Previous studies of the K–P ejecta have suggested some intriguing asymmetries in the distal ejecta pattern [\[18,30,31\]](#page-14-0), with shocked quartz being more abundant in sites to the west of Chicxulub [\[30,31\].](#page-15-0) However, these observations are based upon data acquired by a number of groups, who have adopted different sampling and/or analytical procedures. Detailed mineralogical studies that report the absolute abundance and grain size of shocked quartz are rare [\[31](#page-15-0)–33], and some sites are better documented than others. Hence, it is difficult to make a quantitative assessment of asymmetries in the distribution of the distal ejecta using the existing data alone.

This paper reports the results of a new systematic study of shocked quartz from the global K–P boundary that covered a range of paleodistances and azimuths [\(Fig.](#page-2-0) [1b](#page-2-0); [Table 1](#page-3-0)). Observations of the grain size and degree of shock in quartz have been made using optical and scanning electron microscopy (SEM). The objective of this study was to evaluate whether the shocked quartz is asymmetrically distributed around the crater and, if so, whether that distribution offers a clue to the direction and angle of impact at Chicxulub. In contrast to previous studies, site-to-site comparisons of these new data can be made with confidence because the same protocol was used to analyze each sample.

2. Sampling and analytical procedures

To cover a range of paleodistances and azimuths, a number of sites in Europe, Brazil, New Zealand, North America, and the Pacific and Atlantic ocean regions were chosen for this study [\(Fig. 1b](#page-2-0); [Table 1\)](#page-3-0). Samples

Fig. 1. a) Photographs of lunar craters from Apollo 13. Arrows indicate the direction of travel of the impacting body. Highly oblique impacts produce elliptical craters and a butterfly ejecta pattern, impact angles up to ∼45° produce symmetrical craters with ejecta missing in the uprange direction [\[21\]](#page-14-0). Impacts of >45° show continuous ejecta deposits around the crater. b) K–P paleogeographic map showing maximum size of shocked quartz grains as collated by Claeys et al. [\[48\].](#page-15-0) Locations sampled for this study are identified with a gray rectangle and numbered $1-15$.

were collected from a number of outcrops at each location, to investigate whether there was significant local variation in the K–P layer. In each case we attempted to sample the entire shocked quartz-bearing layer. For North American sites, the Fireball Layer and Boundary Layer (see definition in [\[7\]\)](#page-14-0) were analyzed separately. In New Zealand we sampled the 1-cm thick K–P layer, and in European sections where K–P layer thickness varies between 2 and 4 cm, we took samples from the 2-cm thick sections. Samples from the Ontong Java Plateau (ODP site 803D), Demerara Rise (ODP site 1258A), Shatsky Rise (ODP site 1209), Walvis Ridge (ODP site 1262) and North Pacific (site GPC3) were taken from intervals where either spherules or shocked quartz had been previously identified.

Optical analyses were used to determine the abundance and size of shocked quartz in the bulk volume of K–P rocks in samples from all land-based locations. Quartz grains were extracted after sequential leaching of fossil, carbonate and clay mineral components (as described in Bostwick and Kyte [\[31\]\)](#page-15-0) and dry sieving of the remaining mineral separate. Analyses were carried out on grains extracted from 50 g bulk samples and then mounted into smear slides with ultra violet cement, following the methodology of Montanari and Koeberl [\[33\].](#page-15-0) Shocked quartz grains were identified and marked on the slides, and then counted and measured. The approximate total number of quartz (shocked and unshocked) was estimated using a subset of the prepared slides.

Table 1

Site locations, current distance from Chicxulub, distance from Chicxulub during the K–P, direction from Chicxulub during the K–P, depositional environment, and the impact related minerals identified in this study

See [Fig. 1](#page-2-0)b for site locations. ODP sites are identified by their leg and site number.

The identification of a K–P section in the Poty Quarry in Northeastern Brazil has been the subject of some debate [\(\[34](#page-15-0)–37]). We have analyzed ten 50-g samples from the Poty section from locations where shocked quartz had been reported. Five samples were extracted from the base of a 20- to 50-cm thick layer of limestone breccia, where Koutsoukos [\[35\]](#page-15-0) suggested that ∼20% of the 300 quartz grains counted were shocked. The remaining samples were from the overlying layers from which Koutsoukos [\[35\]](#page-15-0) also reported grains of shocked quartz.

SEM was used to determine the size-distribution of the smaller grains of shocked quartz from all land and marine sites, as well as to document the shock features. The mineral separates for SEM analysis were extracted from 1 g of sample and attached to 1-cm diameter carbon tapes mounted on aluminum stubs. Backscattered electron imagery (BEI) and X-ray microanalysis (EDS) of shocked minerals were carried out on a JEOL JSM840 SEM with an Oxford Instruments exL Pentafet energy X-ray spectrometer (EDS) at the Natural History Museum, London. All shocked mineral grains were photographed, counted and measured, comprising a dataset of nearly 5500 shocked grains in total.

3. Results

3.1. Optical analyses

In [Table 2](#page-4-0) we show the number of shocked quartz grains observed as well as an estimate of the total number of quartz grains, and the maximum and average size for each sample. During the course of the optical analysis we noted that the shallow marine sites from Europe and New Zealand contained large amounts of unshocked quartz with grain sizes >60 μm. The high number of unshocked quartz grains in these samples is indicative of their depositional environment in shallow marine seas (Table 1) (e.g. [\[38\]\)](#page-15-0). We sieved a number of samples from these sites to recover grains of ≤ 100 μm in size, to investigate whether the smaller grains of shocked quartz were also diluted by unshocked grains (bottom half of [Table 2](#page-4-0)). The dilution is similar in both size ranges.

Table 2

Table 2			
Optical analyses on mineral separates from land-based sites			

Samples are named by their location and date of analysis. The total number of quartz grains has been estimated, whereas the number of shocked quartz grains have been carefully counted. The mean size is calculated using all identified shocked quartz grains, except for North American samples for which 100 grains were used to calculate the mean.

 $(**)$ Estimated.

Shocked quartz is abundant in the North American fireball layer with a few thousand shocked quartz grains per 50 g of sample. The maximum observed grain size in the Fireball Layer is 500 μm, but the majority of shocked grains were in the size range from 100 to 200 μm [\(Table 2\)](#page-4-0). The Boundary Claystone yielded relatively few shocked grains. As has been documented in previous optical studies (e.g., [\[7\]\)](#page-14-0), there is no substantial lateral variation in grain size from the southernmost to the northernmost K–P sites in North America. European sites have between a few and 20 shocked quartz grains per 50 g of sample, and shocked quartz is relatively rare in New Zealand. The mineral separate samples of the size range ≤ 100 μm consistently contain more shocked quartz grains than the separates with a size range of $>60 \mu$ m [\(Table 2](#page-4-0)). On average there are ∼1.8 times more shocked quartz grains in the smaller size-range. The abundance of shocked quartz does not vary substantially between samples from the same location, with individual counts varying by less than a factor of 2. Our results are consistent with those of previous optical analyses (e.g. [\[7\]](#page-14-0)).

Our mineral separates (\sim 0.1 g) from the limestone breccia at Poty Quarry comprise quartz (50–60%), feldspar (40–50%) and trace minerals $($ 1%) such as pyroxene, zircon and sphene. Some quartz and feldspar were intensely fractured, but we did not find a single quartz grain with planar deformation features (PDF). Some quartz grains showed more than two parallel fracture planes, but they were widely spaced and discontinuous. It appears that Poty Quarry is not a K–P section.

3.2. SEM analyses

Results from our SEM analysis of 1 g of samples [\(Table 3](#page-6-0)) show that much larger numbers of shocked quartz are identified using the SEM than with the optical studies. For example, there is an average of 55 shocked grains within a 1-g SEM sample from the bottom centimeter at three sites in Italy, whereas, in the optical analysis, the average is 10.4 grains per 50 g in the size range >60 μm and 19 grains per g for particle sizes $100 \mu m$ ([Table 2](#page-4-0)). A histogram of the size of shocked quartz grains, along with the mean at each site [\(Fig. 2\)](#page-7-0) shows that more than 70% of the shocked grains are in the size range from 20 to 80 μm.

Shocked quartz is most abundant in samples from the southernmost sites in North America (Clear Creek North, Berwind Canyon and Madrid East South) with about 800 shocked quartz grains per 0.5 g ([Table 3\)](#page-6-0). The K–P sites in Canada appear to contain slightly less shocked quartz (640 to 760 shocked quartz grains per 0.5 g of sample). At the next closest site, ODP site 1258A in the South Atlantic, the number of shocked quartz grains varies between 42 and 68 per g of sample from the 3-cm thick K–P boundary interval.

The largest numbers of shocked quartz grains in Europe were found at the base of the clay layer in the

The location column indicates sampling depth at ODP sites, bottom or top centimeter at European sites, the entire 1-cm thick K–P layer in New Zealand, and layer thickness at North American sites. The number of quartz grains is estimated, whereas the number of shocked quartz are carefully counted. The average size is for all shocked quartz in the sample, except for North American sites for which the average is based upon 100 measured grains. Similarly, the number of PDF in each quartz grain is for all shocked quartz grains in a sample, except for North American sites for which the values are determined from 100 quartz grains.

 $(*)$ Average of 100 grains.

(**) Fireball layer only.

bottom centimeter, whereas in samples from the Pacific site GPC3 the number of shocked quartz grains varies unsystematically with depth (Table 3). Samples from the ODP site 1258A, in the Atlantic, contain between 11 and 18 shocked quartz grains per g. ODP sites 803 and 1209 in the Pacific, and the site in New Zealand, contain between 2 and 11 shocked quartz grains per g of sample.

3.3. Dilution and reproducibility

We were concerned that the large numbers of unshocked quartz at some sites may have affected our counting, as their presence could make it more difficult to identify the shocked quartz. One sample from the bottom centimeter at Petriccio contained unusually low

Fig. 2. Grain size frequency histograms for shocked quartz grains recovered in K–P samples (see [Table 1](#page-3-0) and [Fig. 1b](#page-2-0) for location), and mean grain size at each site.

volumes of unshocked quartz (50–100 grains), while the other sample contained 500–1000 grains [\(Table 3](#page-6-0)). The number of shocked quartz grains identified in these two samples was 74 and 52 respectively. This suggests that dilution by large numbers of unshocked quartz could reduce the number of shocked quartz identified by $~\sim$ 40%.

The reproducibility of the methods used was tested using aliquots of the same sample. Four aliquots of a sample from Caravaca (Cara-1-2-03-2006a, -2006b, -2006c and -2006d) were separated for optical and SEM observation [\(Table 4a\)](#page-8-0). The results from optical

analysis varied from 4 to 8 grains per 50 g of sample, giving a mean of 5.5 grains and standard deviation of 2.3 grains ([Table 4a](#page-8-0)). For SEM analysis a mean of 30 grains (standard deviation of 10 grains) was obtained [\(Table 4a](#page-8-0)). In each case, the original counts of 5 and 35 grains at Caravaca 1 ([Tables 2 and 3\)](#page-4-0) are within 1 standard deviation of the means derived from the aliquots.

The variation in counts from the same sample is similar to the variation between samples from the same location, with most individual counts varying by less than a factor of 2.

Table 4a Results from four aliquots of the same sample from Caravaca Spain

Sample	Weight	Number of shocked quartz
	(g)	
Optical analysis		
Cara-1-2-03-2006a	50	4
Cara-1-2-03-2006b	50	7
Cara-1-2-03-2006c	50	8
Cara-1-2-03-2006d	50	3
Average		5.5
Standard deviation		2.3
Scanning electron microscopy analysis		
Cara-1-2-03-2006a		37
Cara-1-2-03-2006b		24
Cara-1-2-03-2006c		38
Cara-1-2-03-2006d		19
Average		29.5
Standard deviation		9.5

3.4. Shock features

Our optical and SEM analyses show that shocked quartz (Fig. 3a), are found in all sites sampled except for Poty Quarry [\(Tables 2 and 3\)](#page-4-0), in accordance with previous reports on K–P sites from North America [\[5](#page-14-0)– [7\]](#page-14-0), the Pacific plate [\[31,32\]](#page-15-0) and Europe [\[8\]](#page-14-0). We have

identified quartz characterized by planar fractures, PDF and shock mosaicism, which have been shown to be indicative of increasing degrees of shock pressure [\[39\].](#page-15-0) Most planes of PDF are sharply defined and continuous across the grain (Fig. 3a), but some are only observed around the margins of grains. Some quartz grains with more than one set of PDF are partially isotropic (Fig. 3b). SEM analysis revealed that several shocked quartz grains show evidence of internal recrystallization (Fig. 3c, d). These quartz grains appear to preserve their original shape, but are recrystallized into aggregates of 1–5 μm wide anhedral quartz grains. The aggregates are mostly seen on broken surfaces of the grains and are porous, similar in many aspects to the granular (strawberry) textures reported for shocked zircon in K– P samples from North America (Fig. 3c, d) (e.g., [\[40\]](#page-15-0)).

In all samples we observed quartz with a range of shock features. With the aim of investigating the degree of shock at each site, we have identified all quartz grains with PDF and documented the number of PDF in every grain ([Table 3](#page-6-0)). It is evident that the proportion of grains with 1, 2 or 3 PDF varies between samples.

In contrast to shocked quartz, shocked zircon is very rare in samples from Europe. We recovered six and eight shocked zircon in the size-range 40–80 μm in samples

Fig. 3. a), c), and d) are backscattered electron images; b) was taken in cross polarized light. a) Shocked quartz with two sets of planar deformation features (PDF) from ODP Site 1258A. b) Shocked quartz with partial isotropization (dark area in the center) from Berwind Canyon, North America. c) Partly recrystallized shocked quartz, showing two sets of PDF (arrows) and porous textures from Petriccio, Italy. d) Recrystallized shocked quartz with internal porous textures from Caravaca, Spain. Porous textures are indicative of recrystallization.

from Petriccio and Caravaca respectively. SEM analysis indicates that the shocked zircon grains are characterized by one or multiple sets of planar features, which occur as intersecting planes with spacings of between 4 and 8 μm. Three grains displayed strawberry texture, and this texture has previously been interpreted as a product of recrystallization under high-T conditions after shockinduced amorphization [\[41,42\].](#page-15-0) As for zircon from North American sites (e.g., [\[40,43\]\)](#page-15-0), the granular texture in the porous zircon from the European sites occurs beneath the original planar surface of subhedral grains (Fig. 4c). The largest zircon grains were polished to half their thickness to expose internal structures (Fig. 4d–f). The zircon grains with multiple sets of planar features have the oscillatory zoning typical of crystallization in a felsic magma (Fig. 4e, f), but many grains have also preserved core structures. The cores and compositional zoning are best identified through cathodoluminescent imaging, but the overall poor (dark) luminescence of these grains makes unraveling of the more subtle internal structures difficult and sometimes impossible (Fig. 4f). This is particularly so for recrystallized zircon, which is characterized by extremely fine-grained crystals.

4. Discussion

4.1. Apparent paleodistance and azimuth

As discussed in the introduction, the only plausible mechanism for distributing shocked quartz around the

Fig. 4. a–e) Backscattered electron images. a) Shocked zircon with multiple sets of planar deformation features (PDF) from Caravaca, Spain. Arrows point to the most continuous and straight sets. b) Recrystallized shocked zircon, showing porous texture. Partly recrystallized shocked zircon, with preserved euhedral shape and one set of planar fractures (arrow), from Petriccio, Italy. d–e) Polished sections showing internal structure of porous d) and fractured e) textures of shocked zircon from Caravaca, Spain. f) Cathodoluminescent image of shocked zircon grain shown in e), with multiple zoning and dark luminescent core.

globe is by their acceleration to high velocity within the expanding vapor plume. If the impact at Chicxulub had been sub-vertical, we would expect a symmetrical ejection pattern in all directions. One factor that affects the final ejecta distribution is the Earth's rotation, as sites to the west of Chicxulub move closer as the ejecta is traveling towards them. To partially remove this effect, we have calculated the flight time of ejecta with a take-off angle of 45° to each site, and used this to determine the distance and azimuth of each site from Chicxulub when the ejecta arrived. We refer to this as the apparent distance and azimuth at each site (Table 4b). For example, the Canadian sites moved ∼300 km westwards while these ejecta were in flight, and, at the time of arrival, the site was ∼100 km closer to Chicxulub at an azimuth of 341°.

4.2. Abundance and size-distribution of shocked quartz

It is apparent from [Tables 2 and 3](#page-4-0) that more shocked quartz grains per gram were identified through SEM than optical analyses on all samples for which both methods were used. The grain size plots in [Fig. 2](#page-7-0) show that more than 70% of the shocked quartz grains are smaller than 80 μm, thus, a large fraction of the grains are not detectable by optical analysis. The optical and SEM results are only weakly correlated (correlation coefficient $R = 0.55$) and reflect the observed variation on a local scale within the K–P layer. The 1 g aliquots of the same sample showed a factor of 2 variation, with a minimum of 19 and maximum of 38 shocked quartz grains [\(Table 4a](#page-8-0)), and similar variations are found between different samples from the same location [\(Tables 2 and](#page-4-0) [3](#page-4-0)). These local differences may be directly related to initial variability within the plume or subsequent depositional processes, for example bioturbation and re-working would produce some vertical and lateral dispersion. Given the observed dispersion, and in order to obtain a robust estimate of the local size-distribution, we have grouped the data and determined the mean and variance for all samples that were collected reasonably close together. The new group values are indicated in the first column of Table 4b. For site-to-site comparisons we are forced to use our SEM data, because optical analyses could only be performed at land-based sites.

Sites in North America have the largest number of shocked quartz grains with >1400 grains per g of sample (Table 4b). Sites between 4000 and 8250 km have a few 10 s of grains per g, and sites located at distances of >9000 km have a range of between 5 and 17 grains per g. However, the shocked quartz is located in relatively thin layers in North America, Europe and New Zealand $(\leq 2$ cm) and over an extended interval in the Pacific sites. To measure the total abundance at each location, we have estimated the number of shocked quartz grains per unit surface area (nos/cm² in Table 4b, see caption for details). Shocked quartz at Pacific sites is more dispersed than elsewhere. Where detailed sampling has been performed across several 10 s of centimeters above and

The apparent distance and azimuth incorporate paleo-reconstruction, and rotation of the Earth during transit of material ejected at 45° and traveling on a ballistic path. Nos/g is the mean number of shocked quartz grains per g, and is determined by averaging all samples at the specified location. Nos/ cm² is the number of shocked quartz grains per unit area of the Earth's surface, and is determined using the mean number of shocked quartz grains in each sampled layer ([Table 3](#page-6-0)), summing each layer across the entire K–P interval from which shocked quartz has been extracted (Int. in table above), and using a density of 2 g cm². For example, in Italy the layer thickness is 2 cm, and the mean number of grains per g are 55 and 12 in the bottom and top cm respectively. Hence, in a volume that is 1 cm² by 2 cm deep (mass 4 g), the nos/cm² is $67 \times 2 = 134$. SEM analyses were not performed on the boundary claystone from USA and Canada, hence Int. at these sites is for the fireball layer only. If the boundary claystone were included we estimate that the nos/cm² would increase by 50–100 grains at each of these two locations. The mean number of PDF is determined from the last three columns in [Table 3](#page-6-0).

below the K–P iridium peak, for example site 576B [\(Table 2](#page-4-0) in Bostwick and Kyte [\[31\]](#page-15-0)), the shocked quartz is focused within a relatively thin interval of $>$ 3 and $<$ 7 cm. This contrasts with the iridium anomaly at K–P sites (in the Pacific and elsewhere), which continues to be elevated outside the shocked-quartz bearing intervals for up to a few 10 s per cm, and is thought to be mobilized through diffusion [\[44\]](#page-15-0). Thus, in determining the nos/cm² at Pacific sites, we have assumed that the majority of the shocked quartz lies within the interval that is spherulerich and from where shocked quartz has been previously identified. We recognize that there could be small numbers of K–P shocked quartz outside these intervals, and thus that the total abundance at Pacific sites is less well constrained than at other sites.

The maximum size, average size, average number of shocked quartz grains per gram, and total number of shocked quartz grains per square centimeter all decrease with apparent distance (Fig. 5), and this is consistent with previous K–P studies [\[14,26\]](#page-14-0), as well as with observations of proximal ejecta around craters on other planetary bodies [\[41\],](#page-15-0) and terrestrial volcanic craters [\[42\].](#page-15-0) The maximum size and abundance per g are best fit by power laws, abundance per square centimeter by an exponential law, and mean size with a linear relationship (Fig. 5). In Fig. 5 the solid dot is the apparent distance for ejecta with a take-off angle of 45°, and the arrow indicates the change in apparent distance before the arrival of ejecta with a take-off angle of 70°.

Even if our estimates of nos/cm² for the Pacific are slightly underestimated, they still appear to contradict previous reports that shocked quartz is relatively abundant in the Pacific [\[31,32\],](#page-15-0) that led to publications on the preferential westward distribution of the Chicxulub ejecta [\[18,30,31\].](#page-14-0) Schultz and D'Hondt [\[18\]](#page-14-0) suggested that the North Pacific was downrange of the impact, whereas Alvarez et al. [\[30\]](#page-15-0) argued that the ejection angle and Earth's rotation produced the high abundance of shocked quartz in the Pacific. The previously reported differences between the Pacific region and Europe occur, at least in part, because shocked quartz is more readily identified using SEM. However, our counts of the number of quartz grains per g for the Pacific sites using SEM are different to those reported by Bostwick and Kyte [\[31\].](#page-15-0) They identified concentrations of 95 and 140 shocked quartz grains per g in samples from ODP site 803D and 577B, whereas we identified between 5 and 10 grains per g in sites at

Fig. 5. Results from SEM analyses. Can. is Canada, 1258A is ODP Leg 207 site 1258A, GPC3 is Leg LL44 site GPC3, Den. is Denmark, 1209 is ODP Leg 198 sites 1209, 803D is ODP Leg 130 site 803D, 1262 is ODP Leg 208 site 1262, NZ is New Zealand. The horizontal axis is apparent distance, which incorporates paleo-reconstruction and rotation of the Earth during transit of the ejecta. The solid dots represent the distance from Chicxulub to the site for quartz ejected with a take-off angle of 45° and traveling on a ballistic trajectory. The arrow indicates the change in distance prior to the arrival of quartz ejected at 70°. The dashed line is the best-fit curve through all the data points; the corresponding equation and correlation coefficient (R) are given above the line. a) The maximum and mean size of shocked quartz in 1 g samples of K–P boundary clay plotted against the apparent distance from Chicxulub. b) The average number of shocked quartz found in 1 g samples of K–P boundary clay at each site, plotted against the apparent distance from Chicxulub. Vertical bars are the variance [\(Table 4b](#page-10-0)). c) The total number of shocked quartz grains per unit surface area plotted against the apparent distance from Chicxulub. There is a systematic decrease in number of grains with increasing distance. The top of the vertical bar represents the count assuming a 40% underestimate of shocked quartz, for sites that might be affected by dilution with unshocked quartz (see text for more details).

equivalent paleodistances and azimuths from Chicxulub. To verify our procedure we acquired a sample directly from one of the authors (Kyte), who provided us with 0.45 g of site GPC3 from the interval 2056–57 cm. Bostwick and Kyte [\[31\]](#page-15-0) recovered 351 shocked quartz grains per g from this sample. We recovered 157 quartz grains from this sample, from which only 11 showed definitive evidence of shock metamorphism. This is equivalent to 24 grains per g and almost equal to our original result of 25 grains per g [\(Table 3](#page-6-0)). It appears that some aspect or aspects of sample preservation, sample handling, analytical techniques and mineral extraction procedures can lead to different absolute numbers of shocked quartz recovered and identified. We observe that Bostwick and Kyte's and our results are internally consistent: they both show a decrease in number of shocked grains towards the western Pacific. This forces the conclusion that individual results are protocol dependent, and that comparisons of results from unrelated studies are problematic.

4.3. Reproducibility and dilution

Our counts have been shown to be reproducible (within a factor of 2) with our aliquot tests on the Caravaca sample ([Table 4a\)](#page-8-0), and through our second count of a sample from site GPC3. Hence we are confident that our results are robust enough to be used for site-to-site comparisons. In the Results section, we suggested that we could be underestimating our counts for samples diluted with large numbers of unshocked quartz by ∼40%. We represent the possible effect of this dilution with a vertical bar in [Fig. 5c](#page-11-0): the top of the bar represents the total counted quartz if we assume an original underestimation by 40%.

4.4. Distribution of shock level

Petrographic descriptions from impact craters show that shock metamorphism is highly heterogeneous on a small scale, and small samples of target rock commonly contain mineral phases such as quartz and zircon that display a wide range of shock metamorphic features. Thus, the range of shock features in ejecta at any one K– P site is likely to relate directly to the heterogeneous nature of the shocking process and also reflect mixing in the vapor plume prior to dispersal around the globe. However, the first ejecta incorporated within the expanding plume is, on average, the most highly shocked [\[24,25\]](#page-15-0), hence we might expect this to lead to a systematic change in degree of shock around the globe.

An indicator of the relative average degree of shock at each site has been estimated by determining the mean number of PDF in the shocked quartz grains [\(Table 4b\)](#page-10-0). As an increasing number of PDF in quartz is indicative of increasing shock pressures [\[39\],](#page-15-0) the mean number of PDF is an approximate measure of the relative shock degree at each site. A plot of this estimated shock degree (mean number of PDF in shocked quartz) against apparent distance (Fig. 6a) shows that shock degree increases away from the impact site and is best described by a linear equation. A

Fig. 6. a) The average number of PDF in shocked quartz grains from SEM analysis of 1 g samples of K–P boundary clay, plotted against the apparent distance from Chicxulub. See [Fig. 5](#page-11-0) for site abbreviations. Vertical bars are the variance ([Table 4b](#page-10-0)). The dashed line is the best-fit linear equation through all data points except site 1258A; the corresponding equation and correlation coefficient are given on the graph. The mean number of PDF is a measure of the relative degree of shock at each site. The mean shock level increases with distance from Chicxulub. From the shocked quartz grains at ODP site 1258A an anomalously high mean shock level was determined. b) Calculated probability density functions for the degree of shock at each impact site using the means and variances in [Table 4](#page-10-0)b. N. Am includes all the sites from Canada and the USA; Europe includes all sites in Italy, Spain, and Denmark. Shocked quartz grains at ODP site 1258A plot to the right of all other sites, confirming their anomalously high mean shock level.

plot of the probability density functions (as determined from the means and variances in [Table 4b](#page-10-0)) for grouped sites ([Fig. 6](#page-12-0)b) shows that ODP site 1258A (shaded gray) plots to the right of all other sites. This indicates that the relative degree of shock is anomalously high at this site, and a two-tailed T test confirms this with a significance level of over 99%.

Outputs from numerical models indicate that, in the early stages of impact, the target material is subjected to the highest shock compression pressures and is ejected from the crater at the highest velocities[\[25\].](#page-15-0) Although the initial ejection velocity is \leq 2 kms-1, this highly shocked material enters the first-forming outer part of the plume, and is then rapidly accelerated to high velocities during plume expansion [\[24\].](#page-15-0) Thus, the highest shocked material leaves the crater with the highest velocity and ends up furthest from the impact site. This is consistent with the observed gradual increase in shock degree with apparent distance [\(Fig. 6](#page-12-0)a). For oblique impacts the numerical model predicts that the shock varies with direction and that, for any particular distance from the impact site, the uprange ejecta is the most highly shocked [\[25\]](#page-15-0). The high mean shock degree for the sample from ODP site 1258A may, thus, be indicative of the uprange direction. Also, we note that site 1262, which is further away but in the same direction [\(Fig. 1](#page-2-0)b), has quartz grains that are slightly more highly shocked than expected. However, the error bars on this direction are currently quite large $(\pm 45^{\circ})$. More observational data, as well as more numerical model runs for a range of different impact angles, are required before we can be more conclusive.

4.5. Single or multiple impacts?

Several authors have suggested that there may have been more than one impact at the K–P boundary (e.g. [\[12\]](#page-14-0)), and have referred to the 24-km Boltysh crater in the Ukraine which has been dated as 65.17 ± 0.64 Ma [\[46\]](#page-15-0). A sub-vertical impact would produce symmetric ejection from the impact site, with a gradual decrease in number and size of shocked quartz grains with paleodistance that is independent of azimuth [\[25\]](#page-15-0). Accepting that we have only sampled a subset of regions around the globe, the observed size and shock-distribution of quartz at the global K–P layer are consistent with this layer having been formed from a single, sub-vertical, large impact at Chicxulub. It remains possible that impacts may have occurred around K–P times elsewhere, but unless they had been of similar magnitude to the Chicxulub event, they could not make a significant contribution to the shocked-quartz population at the global K–P boundary.

Stratigraphic interpretations of Cretaceous and Paleocene sections close to Chicxulub have recently been used to argue that the Chicxulub impact took place 300 ky prior to the K–P [\[10,11\].](#page-14-0) The gradual decrease in size and total number of particles ([Fig. 5\)](#page-11-0), and the gradual increase in shock [\(Fig. 6\)](#page-12-0) with paleodistance from Chicxulub, indicate that the impact that produced the K–P distal ejecta layer occurred in the same region of the globe as Chicxulub. The presence of zircon of Yucatán basement age at the distal K–P layer in Canada [\[45\]](#page-15-0), indicate that the K–P impact occurred into rock with the same basement age as Chicxulub, and the large volume of shocked quartz in the K–P layer tell us that the impact was into continental crust. These observations are all consistent with the more widely accepted interpretation that Chicxulub is of K–P age [e.g. [\[9,38\]](#page-14-0), and contrasting biostratigraphical interpretations of the micropaleontological data [\[47\]](#page-15-0).

4.6. Direction and angle of impact

As discussed in the Introduction, an oblique impact at Chicxulub could have produced an asymmetrical ejection of shocked quartz from the impact site. The relatively symmetrical size-and shock-distribution of shocked quartz around the globe, appears to suggest that the impact was not particularly oblique. However, the high shock levels in quartz at ODP site 1258A could be indicative of an uprange direction to the southeast $\pm 45^{\circ}$ [\[25\]](#page-15-0). For oblique impact angles of 15[°] to 45[°] we would expect the proximal ejecta to be missing in the uprange direction [\[21\]](#page-14-0), but thick sequences of proximal ejecta are observed to the southeast (Belize), east (Haiti) and south of Chicxulub (southern Mexico) [e.g [\[38\]](#page-15-0)]. These combined observations suggest that the impact, if oblique, was likely to be at an angle of $>45^{\circ}$.

5. Conclusions

A large volume of material was examined from the proposed K–P boundary at Poty Quarry in Brazil. We did not find shocked quartz within these samples, and conclude that this section is unlikely to be K–P in age. This is the first systematic study of shocked quartz at many K–P sites. The size and number of shocked quartz grains decrease with apparent distance from the Chicxulub crater, and this decrease is fairly symmetric. This contradicts previous studies that proposed that shocked quartz was more abundant than expected in the Pacific, and relatively rare in Europe. We conclude that individual analyses are protocol dependent, and that a comparison between unrelated studies may have led to

the incorrect interpretation that the shocked quartz distribution was asymmetric. The observed relative degree of shock increases with apparent distance, and this is consistent with predictions from numerical models. Both the size-distribution and mean shock degree are precisely as expected, if the K–P layer was formed by a single, large, sub-vertical impact at Chicxulub. This contradicts recent publications that dispute the link between Chicxulub and the K–P boundary. The anomalous mean shock level at a site to the southeast of Chicxulub indicates that the impact may have been oblique, and that this site was uprange. However, the error bars on this direction are currently quite large. The presence of thick sequences of proximal ejecta at sites to the south, southeast and east of Chicxulub indicate that, if the impact was oblique and from the southeast, the impact angle was $>45^\circ$. Previous studies have predicted that Chicxulub produced a dramatic effect on the Earth's climate because it was formed by a low-angle oblique impact into volatile-rich sediments. Our results suggest that these more extreme predictions are unlikely.

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