# **Calcic amphibole growth and compositions in calc-alkaline magmas: Evidence from the Motru Dike Swarm (Southern Carpathians, Romania)**

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### **ABSTRACT**

A Late Pan-African calc-alkaline dike swarm (basalt-andesite-dacite-rhyolite) has been investigated in a region of over 2000 km<sup>2</sup> in the Alpine Danubian window, South Carpathians (Romania). Amphibole phenocrysts and microphenocrysts have been investigated by wavelength-dispersive microprobe analysis and BSE imaging. The Ca-amphibole population, represented in all the lithologies, displays a large compositional range, interpreted as the result of two processes: (1) magmatic evolution (kaersutite  $\rightarrow$ Ti-pargasite → pargasite → Ti-magnesiohastingsite → magnesiohastingsite → edenite → tschermakite → magnesiohornblende) linked to magmatic differentiation from andesitic basalt to rhyolite; and (2) deuteritic alteration of the primary amphibole related to late-emplacement hydrothermal activity (yielding numerous varieties comprising those cited above). In all rock types, amphibole phenocrysts equilibrated at a nearly constant pressure of about  $0.6 \pm 0.1$  GPa, but their temperatures of crystallization ranged from 1000–900 °C for basaltic andesites to 700–600 °C for dacites. In rhyolites, edenite to magnesiohornblende crystals reflect a continuous range of *P*-*T* conditions from 700 °C/0.6 GPa to 600 °C/0.1 GPa, in agreement with their change of habit from euhedral to subhedral. Complex zonations in pargasite-magnesiohastingsite (including resorption) are interpreted in term of self-organization of oscillatory zoning without significant heating and/or magma mixing.

**Keywords:** Crystal structure, calcic amphibole, crystal growth, magmatic zonation, igneous petrology, calcalkaline magma, thermobarometry

## **INTRODUCTION**

In calc-alkaline magmas, such as those of the Late Pan-African Motru Dike Swarm (MDS) in Romania, Ca-amphibole occurs in all petrological facies from andesitic basalts to rhyolites. It represents the main ferromagnesian phase in the lithologic sequence and evolved in composition during magmatic differentiation. Because olivine is lacking in the MDS, even in the less differentiated basalts, amphibole phenocrysts took the place of this more common phase in playing an essential role since the onset of fractional crystallization.

Amphibole composition has been used classically as a differentiation index. However, the high  $H<sub>2</sub>O$  content of many calc-alkaline magmas results in pervasive late-emplacement deuteritic alteration to propylitization and partial modification and/or replacement of primary by secondary amphiboles. Distinction between amphibole parageneses may then become difficult, which so far has limited the potential interest of this phase for petrological models.

This study was aimed at providing a better understanding of the complex chemical compositions and thermobarometric conditions of primary and secondary amphiboles produced during magmatic growth. Using the Motru Dike Swarm as an

illustration, we have characterized the evolution of Ca-amphibole compositions in relation to magmatic and hydrothermal activity throughout melt emplacement. Back-scattered electron images and detailed microprobe investigations of a large population of amphiboles has allowed us to reconstruct the various stages of complex magmatic growth and resorption history of the phenocrysts and of their alteration products (up to rare talc pseudomorphs). This evolution is further constrained quantitatively through *P*-*T* estimates of the conditions of crystallization based on the Al-content empirical barometer (Schmidt 1992) and Ti-content empirical thermometer (modified expression of Otten 1984) for amphiboles.

### **GEOLOGICAL SETTING**

The dikes studied in this work belong to a major Late Pan-African calc-alkaline dike complex, the Motru Dike Swarm (MDS) that crops out within the Danubian window of the South Carpathian Mountains in Romania (Fig. 1). There, south Carpathian crystalline units were thrusted upon the Moesian cratonic platform during the Cenozoic Alpine orogeny in two major nappes: the upper and lower Danubian units (Berza et al. 1983, 1994) recognized on the basis of their Mesozoic cover (Stanoiu 1973; Kräutner et al. 1981). Intruding the Precambrian basement of the lower unit, the MDS, defined by Berza and Seghedi (1975) and \* E-mail: ofemenia@ulb.ac.be recently revisited by Féménias (2003), represents the last (pre-



**FIGURE 1.** (**a**) Map of the pre-Alpine units of the South Carpathian Mountains, from the IGR 1/50 000 map modified, after Berza et al. (1994) and Féménias et al. (2004). (**b**) Relative distribution of the dikes of the Motru Dike Swarm into the Lainici-Paius unit and location of the studied dikes (after Féménias et al. 2004).

Silurian) magmatic event in these Pan-African terranes.

This swarm contains numerous subvolvanic dikes defining a complete differentiated series from basaltic andesite to rhyolite  $(50-72 \text{ wt\% SiO}_2)$ . On the basis of their size, mean orientation, and magnetic fabric, two populations have been identified (Féménias et al. 2004): (1) narrow dikes (<1 m) with a N80° mean wall direction and a symmetrical AMS (anisotropy of magnetic susceptibility) fabric; and (2) thicker dikes (>1 m to several meters) with a N-S mean wall direction and an asymmetrical AMS fabric. The two populations describe a common regional stress field during a brittle emplacement continuum (Féménias et al. 2004).

#### **SAMPLING AND ANALYTICAL PROCEDURES**

Through a MDS regional survey in a mountainous area of 2000 km<sup>2</sup>, some 150 dikes have been sampled. A representative selection of thin sections of 55

samples has been the subject of petrological and mineralogical investigation. Representative analyses of the amphiboles are listed in Table 1. Major-element compositions of amphiboles have been determined by WDS electron microprobe (Cameca SX50) at the Centre d'Analyses par Microsonde en Sciences de la Terre (CAMST, University of Louvain) using a combination of natural and synthetic mineral standards. Operating conditions were as follows: accelerating voltage of 15 kV; beam current of 20 nA; counting time of 10 s for Fe, Mn, Ti, and Cr, 16 s for Si, Al, and K, and Mg, and 24 s for Na.

### **PETROLOGY OF THE MOTRU DIKE SWARM (MDS)**

The MDS is essentially composed of andesitic, trachyandesitic to dacitic dikes, with basaltic andesites, basaltic trachyandesites, and rhyolites being less common (classification according to Le Maitre 2002). Textures are typically aphyric to micro-porphyritic. Porphyritic basaltic andesites and andesites contain euhedral brown amphibole and/or plagioclase microphenocrysts, whereas porphyritic dacites and rhyolites have euhedral plagioclase, green hornblende, and partly resorbed quartz (Fig. 2). The dikes have chilled margins implying a high-level (subvolcanic) emplacement. These margins consist of few microphenocrysts (amphibole and clinopyroxene in basaltic andesites, plagioclase and biotite in andesites) in a secondary (=deuteric) groundmass. In highly hydrated basaltic to andesitic lithologies, plagioclase is locally deeply retrogressed into epidote ± calcite  $\pm$  white mica and commonly appears as ghosts. The groundmass consists of tiny quartz, plagioclase, ferromagnesian phases (amphibole and rare clinopyroxene), oxides, and devitrified glass. Mn-rich ilmenite is the main oxide of the MDS; it occurs either as euhedral to subhedral crystals forming glomeromorphic assemblages (mean grain size: 100-400 µm) or as single tiny (a few tens to a hundred micrometers) anhedral to skeletal grains. Deuteritic alteration and propylitization occurred during and/or just after dike emplacement and induced a low temperature - low pressure secondary paragenesis. Locally, pseudomorphs of leucoxene replace Ti-bearing silicate phases; chlorite, calcite, talc, epidote, and pyrrhotite-pyrite (tertiarily altered into goethite and hematite) developed locally as millimeter-sized euhedral grains. A pseudomorphic replacement of calcic amphibole by talc has been observed and interpreted as a strongly evolved propylitic process (Féménias 2003).

# **AMPHIBOLE COMPOSITION**

Structural formulae of amphiboles from the MDS were calculated on an anhydrous basis assuming 23 O atoms per half unit cell, with the general form  $A_{0-1}B_2C_5T_8O_{22}(OH)_{2}$  representing one formula unit. According to the IMA classification and  $Fe<sup>3+</sup>$  calculation proposed by Leake et al. (1997), all the amphiboles from the MSD are calcic ( ${}^B$ Ca >1 apfu) and belong to two sub-groups of this classification. Group 1 is defined by  $^{A}(Na + K) < 0.5$  apfu (Fig. 3a), and mostly contain high Si contents ( $TSi > 6.5$  apfu) characteristic of magnesiohornblende and actinolite, with minor tschermakite ( $TSi < 6.5$  apfu). Magnesiohornblendes correspond to core compositions of amphibole phenocrysts in rhyolites and dacites, and rim compositions of amphibole phenocrysts from andesites. Actinolite is present in numerous lithologies associated with the deuteritic paragenesis. Core compositions for amphiboles from basaltic andesites, andesites, and trachyandesites belong to Group 2, defined by  $^{A}$ (Na + K) > 0.5 apfu (Fig. 3a). This group mostly contains low Si contents  $(TSi < 6.5$  apfu)

**TABLE 1A.** Selected representative microprobe analyses (major elements in wt%, structural formula) for calcic amphiboles of the calc-alkaline Motru Dike Swarm

		1													
Rock	AB	AB	AB	AB	AB	AB	AB	AB	Α	Α	Α	Α	Α	Α	
Name			3	4	5	6		8		9	3	4	10	5	
SiO <sub>2</sub>	55.65	40.90	44.08	49.91	43.70	41.71	42.19	41.71	53.17	45.02	43.60	45.40	43.50	44.86	43.30
TiO <sub>2</sub>	0.00	4.44	1.96	0.20	2.37	2.75	3.63	3.68	0.03	1.98	2.19	1.68	2.06	2.81	2.33
$Al_2O_3$	0.93	11.10	10.63	6.47	9.41	10.07	11.19	10.88	2.44	11.12	11.18	9.82	11.49	9.60	10.97
FeO	10.16	14.24	12.15	12.94	15.71	19.43	10.45	13.60	15.26	10.77	11.28	11.23	11.47	12.94	12.53
MnO	0.34	0.27	0.19	0.21	0.31	0.30	0.18	0.19	0.34	0.14	0.16	0.19	0.10	0.18	0.13
MgO	17.30	11.69	14.18	14.76	11.45	9.11	14.54	12.04	13.82	14.39	15.03	15.44	13.95	13.39	13.67
CaO	12.84	11.28	11.16	11.86	10.75	10.86	11.34	11.53	12.10	11.59	11.43	11.36	11.69	10.79	11.72
Na <sub>2</sub> O	0.15	2.50	2.18	0.97	2.27	2.39	2.43	2.29	0.28	1.71	2.06	1.62	1.97	2.14	1.96
K,O	0.04	0.94	0.41	0.12	0.95	1.10	0.48	1.00	0.05	0.80	0.54	0.45	0.61	0.42	0.55
$Cr_2O_3$	0.00	0.03	0.08	0.00	0.00	0.00	0.54	0.13	0.00	0.00	0.03	0.11	0.09	0.00	0.00
Cl	nd	nd	nd	nd	nd	nd	nd	nd	nd	0.03	0.08	0.06	0.04	nd	0.07
F	nd	nd	nd	nd	nd	nd	nd	nd	nd	0.23	0.26	0.27	0.25	nd	0.35
Sum	97.41	97.39	97.02	97.43	96.92	97.72	96.97	97.03	97.48	97.78	97.82	97.62	97.23	97.13	97.58
	Notes: AB = Basaltic andesite; A = Andesite. 1 = actinolite; 2 = kaersutite; 3 = magnesiohastingsite; 4 = magnesiohornblende; 5 = titanian-edenite; 6 = ferro-pargasite; 7 =														
titanian-magnesiohastingsite; 8 = titanian-pargasite; 9 = edenite; 10 = pargasite; 11 = titanian-magnesiohornblende; 12 = tschermakite; 13 = ferro-hornblende.															

**TABLE 1B.** Selected representative microprobe analyses (major elements in wt%, structural formula) for calcic amphiboles of the calc-alkaline Motru Dike Swarm



Notes: A = Andesite; T = Trachyandesite; D = Dacite; R = Rhyolite.1 = actinolite; 2 = kaersutite; 3 = magnesiohastingsite; 4 = magnesiohornblende; 5 = titanian-edenite; 6 = ferro-pargasite; 7 = titanian-magnesiohastingsite; 8 = titanian-pargasite; 9 = edenite; 10 = pargasite; 11 = titanian-magnesiohornblende; 12 = tschermakite; 13 = ferro-hornblende.



**FIGURE 2.** Sketch illustrating modal evolution of the main phases during magma differentiation of the Motru Dike Swarm.

characteristic of magnesiohastingsite, pargasite, and kaersutite, with minor edenite ( $TSi > 6.5$  apfu). Such magnesiohastingsite, pargasite, and kaersutite correspond to core compositions of amphibole phenocrysts from andesitic basalts, trachyandesites and andesites, whereas edenites result from a moderate alteration of the Si-poor amphiboles and, accessorily, to core compositions of amphiboles from dacites. Core compositions of amphiboles from andesites are less titaniferous than those from basaltic andesites and trachyandesites (Fig. 3b), with the exception of some titanian-edenite. Basaltic andesites and trachyandesites also contain titanian-pargasites and kaersutites that represent the first step in the alteration of titanian-edenite compositions. All the amphiboles in rhyolites and dacites have low Ti contents

 $(Ti < 0.25$  apfu). Regardless of the amount of Na + K in the A site, the  $X_{\text{Me}}$  [=Mg/(Mg + Fe<sup>2+</sup>)] of the MDS amphiboles ranges from 0.5 to 1 (Figs. 4a and 4c), making them all magnesian. Rare ferro-hornblende, ferro-edenite, titanian-ferro-paragasite, and ferro-actinolite have been observed as accessory alteration products near rims or cracks. Primary magnesiohornblendes in the rhyolites are more magnesian than those resulting from secondary recrystallization of andesite (Fig. 4c). Actinolites derived from basaltic andesite phenocrysts are somewhat more magnesian than those from andesites (Fig. 4c). Within the magnesiohastingsite-pargasite population, the Fe<sup>3+</sup>/<sup>IV</sup>Al ratio (Fig. 4b) discriminates andesite magnesiohastingsites ( $Fe^{3+/IV}Al > 1$ ) from basaltic andesite and trachyandesite pargasites  $(Fe^{3t/N}Al < 1)$ . Most amphiboles from these two lithologies are thus characterized by a low  $Fe<sup>3+</sup>$  content, although a local increase in  $Fe<sup>3+</sup>$  has been observed, linked to zonation in andesite phenocrysts and alteration products.

# **AMPHIBOLE GEOTHERMOBAROMETRY**

Numerous studies have dealt with parageneses of calcic amphibole in mafic (meta-) igneous rocks (Allen et al. 1975; Allen and Boettcher 1978; Helz 1973; Spear 1981; Ernst and Liu 1998). These studies make it clear that, with increasing *P*-*T* conditions, calcic amphiboles exhibit increases in  $Mg/(Mg + Fe)$ and in K, Al, Na, and Ti contents, and commensurate decreases



**FIGURE 3.** Compositional variations of calcic-amphibole from the Motru Dike Swarm. Cation proportions are per formula unit. Amphibole classification after Leake et al. (1997).

in Si and total  $Fe + Mg + Mn + Ca$ .

Pressure conditions of crystallization, deduced from contact aureoles or experimentally controlled, have been observed to be correlated linearly with the Al-content in amphibole, and, many calibrations of the Al-in-amphibole barometer have been published (Hammarstrom and Zen 1986; Hollister et al. 1987; Johnson and Rutherford 1988; Rutter 1989; Blundy and Holland 1990; Schmidt 1992; Anderson and Smith 1995; Ernst and Liu 1998). The Anderson and Smith (1995) calibration is *T*-dependent; the pressure estimate is corrected for the temperature effect. It appears that this correction dramatically increases when the temperature increases from 800 up to 1100 °C. Anderson and Smith (1995) recognized that "the new expression probably should not be applied to plutons with temperatures much >800 °C because such conditions are far outside the range of experimental calibration…" p. 554). Indeed, when this calibration is applied to the high-*T* Ti-pargasite of the MDS, it leads to much lower (even negative!) *P* estimates than for the (low-*T* and low-Ti) magnesiohornblende. According to experimental data summarized by Ernst and Liu (1998), the range of *P*-*T* estimates for the Motru amphiboles does not justify the use of a coupled geothermobarometer. Indeed, Ernst and Liu (1998) have clearly shown in their Figure 8 that the  $Al_2O_3$  content of Ca-amphibole and  $T$  are very poorly correlated and that the  $TiO<sub>2</sub>$  content of Ca-amphibole and *P* are not correlated in the temperature range 700–1100 °C at  $P = 0.5 \pm 0.5$  GPa. All the barometer expressions usable for the Motru amphiboles give, within error, similar *P* estimates [see Fig. 1 of Anderson and Smith (1995) for a similar observation]. Comparisons of these experimental calibrations,



**FIGURE 4.** Compositional variations of calcic-amphibole from the Motru Dike Swarm. Symbols as in Figure 3.

however, show the important influence of fluid composition on equilibration of amphibole and that of the buffer phase assemblage. Considering the normal MDS assemblage amp  $+$  pl  $+$  Kfs + qz+ il and the usual lack of fresh plagioclase, the experimental calibration of Schmidt (1992) has been used, for it reduces the uncertainty on pressure estimates  $(\pm 0.06 \text{ GPa})$ .

In the reduced *P*-*T* domain (*P* =  $0.5 \pm 0.5$  GPa; *T* =  $900 \pm$ 400 °C), the only usable experimental data for *T* estimates are those from Helz (1973, 1976, 1979). According to these data and to the more recent works (Otten 1984; Ernst and Liu 1998), Ti solubility appears to be a suitable geothermometer for the MDS amphiboles. Indeed, as in the mineral assemblages reported here, the synthesized calcic amphiboles [Helz (1973) and references in Ernst and Liu (1998)] are associated with Ti-rich minerals (ilmenite, rutile, titanite) and may be considered as Ti-saturated. However, the graphical semi-quantitative calibration of Ernst and Liu (1998) yields large uncertainties whereas the independent quantitative calibration of Otten (1984) introduced a change of slope (in fact a mathematical artifact) at 970 °C in the linear

relation between Ti and *T* that makes this calibration inapplicable to the high-Ti amphiboles of the Motru Dike Swarm. However, in the original data set of Helz (1973), the experimental bulk compositions, the *P*-*T* domain of the experiments, and the related mineral association (Helz 1982), on the one hand, and the fact that the FMQ buffer is applicable to the MDS samples on the other hand, it is possible to derive a unique *T* calibration from the Ti content of these amphiboles. We propose a quantitative thermometer applicable to high- $T$  ( $>700$  °C) amphibole crystallizing in Ti-saturated calc-alkaline magma. Our *T* estimates are globally in agreement with those deduced from the Blundy and Holland (1990) thermometer. Nevertheless, we did not use their calibration because fresh plagioclase is rare; most plagioclases of the Motru rocks have suffered strong deuteritic alteration.

Solubility of Ti in calcic amphibole buffered by a Ti-rich phase displays an ideal solid-solution behavior:

$$
K_{\rm D} \approx \left[{\rm Ti}\right]_{\rm amphibole} \tag{1}
$$

Because it is independent of low-pressure variations (Otten 1984; Ernst and Liu 1998),  $K<sub>D</sub>$  may be expressed as a simple function of temperature:

$$
\ln K_{\rm p} = \frac{\rm b}{T} + \rm c \tag{2}
$$

Hence,

$$
\ln[Ti]_{\text{amphible}} = \frac{b}{T} + c \tag{3}
$$

where b and c are two constants. According to the FMQ data of Helz (1973),

$$
\ln[\text{Ti}]_{\text{amphible}} = \frac{2603}{T} - 1.70\tag{4}
$$

where  $[Ti]_{\text{amphibole}}$  is expressed in atoms per formula unit (apfu). The quality of the fit to the data ( $r^2 = 0.951$ ) yields an uncertainty (2 $\sigma$ ) that ranges from  $\pm 15$  to  $\pm 55$  °C for temperatures ranging from 600 to 1000 °C, respectively.

Most of the calcic amphiboles in the basaltic andesites from the MDS crystallized at relatively high pressures (Figs. 5a and 5b: ~70% of analyses yield pressures between 0.6 to 0.7 GPa) regardless of the crystal size, the location of the analysis (core or rim), or the nature of the amphibole (kaersutite, titanian-pargasite, magnesiohastingsite, or titanian- magnesiohastingsite). Titanian-magnesiohastingsites, and pargasites from andesites, as well as core compositions of edenites, ferro-edenites, and tschermakites from dacites and rhyolites, display similar pressure conditions of crystallization (Figs. 5a and 5b). In contrast, magnesiohornblendes from rhyolites and dacites show a broader range of values  $(0.7 \text{ to } < 0.2 \text{ GPa})$  in relation to the subhedral habit of these amphiboles, which may reasonably represent the progressive cooling of the dikes. Titanian-pargasites and titanianmagnesiohastingsites from trachyandesites display a slightly but significantly lower mean pressure  $\left(\sim 75\% \text{ of analyses yield pres-} \right)$ sures between 0.5 to 0.6 GPa), which distinguishes them from basaltic andesite amphiboles.

In contrast to crystallization pressure, temperatures (Figs. 5a and 5b) are more sensitive to the nature and habit of crystals. Temperatures vary according to amphibole mean compositions (as well as  $X_{\text{Mg}}$ ), and show a large range of values from 1100 to 600 °C (temperatures for actinolite were not investigated) irrespective of the nature of the host lava. Nevertheless, amphibole-core temperatures decrease continuously from basaltic andesites (1000–1100 °C for kaersutite and titanian-pargasite) to dacites and rhyolites (60–750 °C for magnesiohornblendes and edenites); rim compositions and alteration products are at the origin of the dispersion of the values for any a given chemical environment. In andesitic basalts, trachyandesites and andesites, the correlation between Ti-content of amphibole (temperature) and amphibole type is thus characteristic of the late-emplacement (at sub-surface conditions) alteration features of the deep amphiboles (Al does not vary significantly). This correlation results from easier diffusion of Ti than Al in the crystal structure during the deuteritization and/or propylitization events.

Populations of amphibole compositions (Figs. 5a and 5b) hence characterize a progressive decrease in temperature during a first stage of crystallization, with magma differentiation (from basaltic andesites to dacites) at a constant pressure of 0.6–0.7 GPa and, in a second stage, a progressive decrease in pressure, showing that crystallization proceeded during the ascent of the dacites and rhyolites. Amphiboles from trachyandesites also integrate this scheme of evolution, but at a lower pressure.

### **AMPHIBOLE GROWTH STRUCTURES**

Almost all phenocrysts of calcic amphibole in basaltic andesites to andesites display zoning observable microscopically in polarized light and on BSE images. Zoning is complex and characterized by two types of internal boundaries. The first type, which is dominant, is "geometric," concentric, and regular (Fig. 6, stages 1, 3, 5, 7, and 9); it shows up as transitions from one band to another without sharp discontinuities and looks like a crystalline growth structure (i.e., Vance 1965). The second type is not concentric, but is commonly irregular (Fig. 6, stages 2, 4, 6, and 8); it cuts across internal boundaries of type 1 and/or 2 without distinction, and shows up as a sharp discontinuity within the grain. Many boundaries of type 2 are curvilinear and commonly associated with resorption gulfs and/or slight blunting of the internal crystal edges. Some of them are associated with cracks and penetrate throughout the core of the crystal, modifying its morphology during growth (Fig. 6, stage 6). All the observed zoned amphiboles from the MDS, whatever their size, display several boundaries of each type.

A detailed electron microprobe investigation of these zonations was performed on several phenocrysts and microcrysts of amphibole. Fifteen to 90 points were analyzed on each crystal, depending on size (from 100 up to 2000 µm). All the results are roughly similar and only one representative magnesiohastingsitepargasite zoned crystal is shown here as an example (Fig. 6). Microprobe analyses of this single amphibole yield an overall magnesiohastingsite composition, with some pargasite and rare tschermakite point analyses. In a general way, Si and Mg-contents increase slightly from core to rim (from  $\sim 6.3$  to  $\sim 6.4$  and  $\sim 3.0$  to  $\sim$ 3.3 apfu, respectively) whereas the Ca-content decreases from 1.83 to 1.74 apfu Sodium, K, Ti, and Al-contents are roughly constant all along the profile, whereas the Mn-content shows a U-shaped profile  $(0.25-0.1-0.25$  apfu). In detail, a majority of



**FIGURE 5.** Statistical pressure and temperature distribution of amphibole crystallization for the principal lithologies of the Motru Dike Swarm. Symbols as in Figure 3.

the elements—with the exception of Na, Ti, and Al—display complex variations, which may be described at different scales. Indeed, at a small scale (ca. 20 µm), the crystal displays oscillatory characteristics (i.e.,  $\text{Fe}^{2+}/\text{Fe}_{\text{tot}}$  and  $X_{\text{Mg}}$  between 170 and 400 µm on Fig. 6), more or less linked to the visible internal boundaries of both types. This oscillatory zoning is roughly "normal" with an outward decrease of  $X_{\text{Mg}}$  and an increase of the Fe<sup>2+</sup>/Fe<sub>tot</sub> ratio. At a larger scale (up to 200  $\mu$ m), chemical evolution within the crystal translates into continuous changes without clear polarities; both types of internal boundaries are then generally linked with chemical steps and/or jumps. However, no evident systematic distinction can be made between these two types of boundaries, especially when the type 2 is not clearly associated with resorption gulfs (Fig. 6). Beyond this detailed description, zoning may be summed up according to the thermodynamic behavior of each element in calcic-amphiboles: Ti, Na, and Al, which are strongly *P*/*T*-dependent do not vary significantly, whereas elements that are dependent on the system chemistry show small-scale variations, showing up as slight oscillations. The very weak apparent temperature  $(T =$ 792  $\pm$  17 °C) and pressure ( $P = 0.63 \pm 0.02$  GPa) variations are not interpreted as having any geological significance.

# **DISCUSSION**

### **Calcic amphibole composition variations**

The calc-alkaline Motru Dike Swarm is characterized by prominent calcic-amphibole phenocrysts and/or matrix microcrysts. These calcic amphiboles represent the main ferromagnesian phase in the absence of olivine, but unlike olivine, they remain on the liquidus throughout the magmatic sequence ranging from basaltic andesites to rhyolites. These amphiboles display a wide range in composition from Si-poor, Ti-rich kaersutite to Si-rich actinolite. The absence of significant differences in composition between small  $\left($ <50  $\mu$ m) and large (up to 3000 µm) amphiboles precludes the possibility of many generations of crysts. Although this range of compositions does not seem to be correlated directly with the nature of host lava (many types of amphibole being present in each lithology), detailed studies of each population show that this variety results from superimposition of two contrasted processes illustrated by: (1) cores, which represent unperturbed and unaltered amphiboles, illustrating a magmatic evolution (kaersutite  $\rightarrow$  Ti-pargasite  $\rightarrow$  $pargasite \rightarrow Ti-magnesiohastingsite \rightarrow magnesiohastingsite \rightarrow$ edenite → tschermakite → magnesiohornblende) correlated to the degree of magmatic differentiation of the host lava (respectively from basaltic andesite to rhyolite); and (2) rims, which represent perturbed- and/or altered-amphibole compositions, that are compatible with the secondary deuteritic (Laverne et

al. 1995) paragenesis common to all rocks; these compositions are interpreted as resulting from partial deuteritic alteration of primary amphiboles related to late-emplacement hydrothermal activity (inducing the many varieties comprising those cited above) and not as a new generation of crysts, as inferred from their high mean Al-contents, which are not compatible with a low-*P*, high-level of crystallization.

### *P***-***T* **records and magma-chamber processes**

In all rock types, primary amphiboles were equilibrated at about the same pressure (0.6–0.7  $\pm$  0.1 GPa), although their temperatures of crystallization range from 1000–900 °C for basaltic andesites to 700–600 °C for dacites and rhyolites. Magma differentiation of the MDS is thus characterized by progressive cooling. This cooling occurred during fractional crystallization in deep magma chambers emplaced at the same level within the lower crust  $(\sim 20 \text{ km})$ . In rhyolites, edenite to magnesiohornblende grains reflect a continuous range of *P*-*T* conditions from 700 °C/0.6 GPa to 600 °C/0.1 GPa in agreement with their change of habit from euhedral to subhedral. These amphiboles record the different stages of crystallization in continuum: they began to grow in a deep environment and developed during magma ascent from chambers to the surface. Trachyandesitic lithologies, which are less common in the MDS, may correspond to hot magma trapped in shallower chambers. Their specific whole-rock compositions do not affect the nature of the amphibole, which



**FIGURE 6.** Zoning of a representative calcic-amphibole in andesite. (**a**) BSE image. (**b**) Sketch representation of the same amphibole crystal illustrating the mean observed internal boundaries and schematic growth history in 10 principal steps alternating growth-dominated stages (1, 3, 5, 7, 9) and resorption-dominated stages (2, 4, 6, 8). (c) Ti<sup>4+</sup>, Al<sub>tot</sub>, Fe<sup>2+</sup>/Fe<sub>tot</sub>, and  $X_{Mg}$  variations along the profile. Symbols as in Figure 3.

integrates the mineralogical sequence of the MDS described before. According to the higher level of their magma chambers, the chemical particularity of these rocks (higher  $Na<sub>2</sub>O + K<sub>2</sub>O$ content) may be due to moderate crustal assimilation.

Pargasite and magnesiohastingsite display complex zoning and resorption features that have been correlated tentatively by combining BSE images with electron microprobe analyses. In any given sample, each amphibole crystal, regardless of its size, displays a similar chemical composition and internal type-features. These internal features have been detailed and visually divided into two types (types 1 and 2, depending on whether or not they are associated with resorption features, respectively). Although internal zoning boundaries of type 2 display significant changes in the amphibole composition, they are not really different from boundaries of type 1. All internal limits may be considered as growth disruptions, more or less marked by crystal resorption, and they do separate numerous domains within each amphibole. These domains do not show complete evidence for a single oscillatory process, although chemical zonations may be locally considered as "normal" (i.e., decreasing  $X_{\text{Mg}}$ ). All these zoned grains display a duality between roughly constant *P*, *T*-dependent element-contents (i.e., Ti and Al) and variable system-dependent element-contents (i.e.,  $Mg$ ,  $Fe^{2+}$ ,  $Fe^{3+}$ ). Oscillatory zoning and repetitive resorptions cannot be the result of new pulses of magma in the chamber because: (1) the number of pulses (up to 10 in some grains) is unrealistic given the enormous difference between the time necessary to homogenize the magma and the time required to grow millimeter-sized phenocrysts (i.e., Kuo and Kirkpatrick 1982; Cashman 1992; 1993; Wörner 1996);  $(2)$  the resorption features are not linked to a significant change in thermo-dependent element-content and, thus, to some significant heating of the amphibole [e.g., as in the model developed for similar calc-alkaline dikes by Pudlo and Franz (1995)]; and (3) resorptions are not related to spongy textures, as is currently observed in magma mixing models (Tsuchiyama 1985). According to experimental results on oscillation processes in mineral phases, this oscillatory zoning and the more-or-less pronounced associated resorption may be interpreted in terms of self-organized oscillatory zoning (Wang and Merino 1993; Wang and Wu 1995) in a convective magma chamber without external contribution of heat and/or magma mixing processes (L'Heureux and Fowler 1996; L'Heureux and Jamtveit 2002).

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