

## Galaxite, $\text{MnAl}_2\text{O}_4$ , a spectroscopic standard for tetrahedrally coordinated $\text{Mn}^{2+}$ in oxygen-based mineral structures

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### ABSTRACT

Chemical analyses, crystal structure refinement, cation order determination, and single-crystal optical absorption spectrum of synthetic galaxite are presented. New optical absorption spectra of natural Mn-bearing willemite, rhodochrosite, Mn-rich forsterite, and tephroite are reported for comparative purposes.

The structure of a synthetic galaxite end-member is characterized by a relatively large unit-cell edge,  $a_0 = 8.2104(3)$  Å, a  $u$ -parameter equal to 0.26588(7), a T-O distance of 2.0034(6) Å, and an M-O distance of 1.9310(5) Å.  $\text{Mn}^{2+}$  is strongly ordered at the tetrahedral T-site as demonstrated by the refined structural formula  $^{\text{T}}(\text{Mn}_{0.90}^{2+}\text{Al}_{0.10})^{\text{M}}(\text{Mn}_{0.10}^{2+}\text{Al}_{1.90})\text{O}_4$ .

The optical absorption spectrum of galaxite in the range 300–800 nm shows a set of five relatively sharp bands at 20300, 22250, 23390, 25970, and 27780  $\text{cm}^{-1}$  marking spin-forbidden transitions in  $\text{Mn}^{2+}$  at the tetrahedral site. The molar absorption coefficient of the field-independent  ${}^6\text{A}_1(\text{S}) \rightarrow {}^4\text{E}_g^4\text{A}_1(\text{G})$  absorption band at 23390  $\text{cm}^{-1}$  equals 1.90 L/(mol·cm), which is approximately an order of magnitude higher than for corresponding bands in spectra of compounds that contain isolated  $\text{Mn}^{2+}$ -centered octahedra.

The calculated crystal field splitting, 10Dq, for  $\text{Mn}^{2+}$  at the T-site in galaxite equals 5290  $\text{cm}^{-1}$ . This compares well with derived 10Dq-values of 5860 and 5510  $\text{cm}^{-1}$  for  $\text{Mn}^{2+}$  at the tetrahedral T1- and T2-site in Mn-bearing willemite. In agreement with theory, the 10Dq for  $\text{Mn}^{2+}$  in  $\text{MnO}_4$  tetrahedra is ca. 30% smaller than corresponding values in  $\text{MnO}_6$  octahedra. The lower Racah B-parameters of the spectroscopic data indicate that the degree of covalency of  $\text{Mn}^{2+}$ -O bonds is higher in tetrahedra than in octahedra.

**Keywords:** Crystal structure, galaxite, optical spectroscopy, willemite, tephroite, forsterite

### INTRODUCTION

Manganese, like iron, is a geochemically important element occurring as divalent, trivalent, and tetravalent cations at differently coordinated sites in common silicate and oxide minerals. Methods to distinguish the valence state and coordination of transition elements in mineral structures are essential for many branches of the earth sciences. Optical absorption spectroscopy is one of several methods that have the capability to extract such information at high spatial resolution.

The optical-absorption spectra of  $\text{Fe}^{2+}$  and  $\text{Fe}^{3+}$  in four-, five-, and sixfold coordination in minerals have been explored extensively (e.g., Rossman and Taran 2001; Taran and Langer 2001; Waychunas and Rossman 1983), but comparable data for manganese show some omissions. Many optical spectra of  $\text{Mn}^{2+}$ - and  $\text{Mn}^{3+}$ -bearing minerals have been published during the last four decades, but complete optical spectra of manganese at tetrahedral sites in oxygen-based mineral structures have not been reported. In addition to this, the vast majority of the published spectra deals with mineral structures in which  $\text{Mn}^{2+}$  is distributed at high concentrations over several sites that show

different mean Mn-O bond lengths and variable distortion from octahedral symmetry. In fact, presence of several non-equivalent Mn-centered sites is characteristic for all the silicate minerals containing  $\text{MnO}_6$ -clusters, which are included in the comprehensive list of spectroscopic data on  $\text{Mn}^{2+}$ -bearing minerals presented by Burns (1993). This structural complexity makes spectral band assignments difficult. Consequently, published parameters (e.g., crystal field splitting and Racah) on  $\text{Mn}^{2+}$  at octahedral sites in minerals derived from optical absorption spectra must be regarded with some caution.

For these reasons we present in this paper optical absorption spectra of  $\text{Mn}^{2+}$  in tetrahedral oxygen coordination in spinel (galaxite) and phenacite (willemite) group minerals. To relate the spectra of divalent manganese at tetrahedral sites with those of  $\text{MnO}_6^{0-}$ -cluster we have recorded spectra of three structurally simple minerals; Mn-rich forsterite, tephroite, and rhodochrosite at relatively high spectral resolution.

### SAMPLES AND METHODS

#### Synthetic galaxite

Single crystals of end-member galaxite,  $\text{MnAl}_2\text{O}_4$ , were synthesized by use of a flux-growth method. Nutrient components

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were weighed out in pre-dried oxide form (analytical grade) with a surplus of Al<sub>2</sub>O<sub>3</sub> to minimize potential stabilization of trivalent manganese in the products. The nutrient oxides were mixed with Na<sub>2</sub>B<sub>4</sub>O<sub>7</sub> acting as flux compound, using a flux/oxide ratio of 0.8, and were homogenized under acetone in an agate mortar. Approximately 7 g of the mixture was placed in a 12 cc crucible of yttria-stabilized Pt/Au(5%), covered by a lid, and placed in a vertical furnace equipped with Tylan CO<sub>2</sub>:H<sub>2</sub> flow controllers. A continuous flow of the two gases in the ratio 1:2 was adopted to obtain oxygen fugacity ranging from 10<sup>-12</sup> to 10<sup>-17</sup> bars when the temperature was varied from 1200 to 900 °C. Thermal runs consisted of an initial increase to 1200 °C where the temperature was kept constant for 24 h to allow for complete dissolution of the oxides and melt homogenization. Subsequently, the temperature was decreased by 4 °C/h to 900 °C, followed by a fast cooling down to ambient temperature. The synthesis product consisted of yellow-colored galaxite and colorless corundum crystals dispersed in sodium borate glass. The glass was dissolved in a warm diluted HCl solution.

### NATURAL SAMPLES

Single crystals of Mn-bearing willemite ([Zn, Mn]<sub>2</sub>SiO<sub>4</sub>), Mn-rich forsterite ([Mg, Mn]<sub>2</sub>SiO<sub>4</sub>), tephroite (Mn<sub>2</sub>SiO<sub>4</sub>), and rhodochrosite (MnCO<sub>3</sub>) were selected from specimens in the mineral collections at the Swedish Museum of Natural History. Information on provenience and collection numbers of these specimens is summarized in Table 1. Single crystal absorbers of these optically anisotropic minerals were prepared from euhedral grains oriented with the aid of crystal morphology and optical microscopy utilizing conoscopic interference figures.

### ELECTRON MICROPROBE ANALYSIS

Electron microprobe analysis of the present samples were obtained with a Cameca SX50 instrument at the University of Uppsala operating at an accelerating potential of 15 kV and a sample current of 12 nA. Standard samples were synthetic Al<sub>2</sub>O<sub>3</sub> (for Al), MgO (Mg), CaSiO<sub>3</sub> (Ca and Si), MnTiO<sub>3</sub> (Mn and Ti), Fe<sub>2</sub>O<sub>3</sub> (Fe), ZnS (Zn), and natural albite (Na), orthoclase (K), and vanadinite (Pb). For raw data

reduction, the PAP computer program was applied (Pouchou and Pichoir 1984). The results are summarized in Table 1.

### Single-crystal structural refinement

A single crystal (ca. 250 × 250 × 250 μm) of synthetic galaxite was mounted on a Siemens P4 automated four-circle, single-crystal diffractometer. Details of the data collection are summarized in Table 2. For collection of diffraction intensity data, one-eighth of the reciprocal space was examined with the ω scan method at a fixed scan range. The scan speed was variable, depending on reflection intensity, as estimated through pre-scans. The background was measured with a stationary counter and crystal at the beginning and end of each scan, in both cases for half the scan time. Three standard reflections were monitored every 47 measurements. Further details of the experimental conditions can be found in Lucchesi et al. (1997).

The SHELXTL-PC program package was used for reduction of X-ray diffraction data. Intensities were corrected for polarization and Lorentz effects. Absorption corrections were accomplished with a semi-empirical method (North et al. 1968). All reflections were used in the refinement. No significant deviations from *Fd* $\bar{3}m$  symmetry were noted. Appearance of forbidden space-group reflections such as (200) was attributed, on the basis of ψ-scan checks, to double reflection (Tokonami and Horiuchi 1980). Variable parameters during the structural refinement were scale factor, oxygen coordinate, mean atomic number of tetrahedral (T) and octahedral (M) sites, displacement parameters, and isotropic secondary extinction coefficients. The starting oxygen coordinate proposed by Lucchesi et al. (1997), setting the origin at  $\bar{3}m$ , was used. No chemical constraints were applied during the refinement. Partially ionized (for O and <sup>M</sup>Al) and fully ionized (for <sup>T</sup>Mn<sup>2+</sup>) scattering curves were applied as these resulted, as observed by Della Giusta et al. (1996), in the best conventional agreement factors over all sinθ/λ intervals and the best correspondence between observed and calculated F(222). Off-diagonal displacement parameters were extremely small and of the same magnitude as their standard deviation, when not forced to zero by symmetry. Three isotropic, full-matrix, refinement cycles were followed by anisotropic cycles until convergence was attained. No significant correlation (*r*<sup>2</sup> < 0.40) between parameters was observed

### Optical absorption spectroscopy

Optical absorption spectra in the UV/VIS range, 300–800 nm (33 333 to 12 500 cm<sup>-1</sup>), were recorded at room temperature on double-sided polished single crystals of high optical quality. The absorber thickness of the studied minerals was 103 (galaxite), 206 (tephroite), 389 (Mn-rich forsterite), 415 (rhodochrosite), and 893 μm (willemite). Spectra were recorded with a Zeiss MPM800 single beam microscope-spectrometer using Zeiss Ultrafluor 10× lenses as condenser and objective. A 75W Xenon arc lamp served as a light source and a photomultiplier as detector. Polarized spectra of the anisotropic phases were recorded using a Glan-Thompson prism. The accuracy of wavelength and absorbance readings of the spectrometer was

TABLE 1. Sample origin and composition

Mineral	Galaxite	Willemite	Forsterite	Tephroite	Rhodochrosite
Locality	Synthetic	Franklin, NJ, U.S.A.	Långban, Sweden	Franklin, NJ, U.S.A.	CO, U.S.A.
Collection no.		NRM no. 333857	NRM no. 15934	NRM no. 24308	NRM no. 881868
SiO <sub>2</sub> (wt%)	nd	27.06	34.94	29.30	nd
Al <sub>2</sub> O <sub>3</sub>	58.73	nd	0.03	0.00	nd
TiO <sub>2</sub>	nd	nd	nd	nd	nd
MgO	nd	0.18	25.14	0.43	0.67
CaO	nd	nd	0.23	0.66	0.15
MnO	41.35	6.64	38.66	67.65	60.24
FeO	nd	nd	0.05	0.03	0.05
ZnO	nd	66.72	0.12	1.59	0.02
PbO	nd	nd	nd	nd	0.41
CO <sub>2calc</sub>					38.27
SUM	100.08	100.60	99.16	99.66	99.81

### Formula based on 3 cations (galaxite, willemite, forsterite, and tephroite) and 1 cation (rhodochrosite)

Si (apfu)	nd	0.987	0.993	0.986	nd
Al	1.992	nd	0.001	0.000	nd
Ti	nd	nd	nd	nd	nd
Mg	nd	0.010	1.065	0.022	0.019
Ca	nd	nd	0.007	0.024	0.003
Mn	1.008	0.205	0.930	1.928	0.976
Fe	nd	nd	0.001	0.001	0.001
Zn	nd	1.797	0.002	0.040	0.000
Pb	nd	nd	nd	nd	0.001

Notes: nd = not detected. Estimated relative errors for major elements are ±1% in all analyses.

monitored by recording spectra of Ho<sub>2</sub>O<sub>3</sub>-doped and Pr<sub>2</sub>O<sub>3</sub>/Nd<sub>2</sub>O<sub>3</sub>-doped calibration standards (Hellma glass filters 666-F1 and 666-F7). Sample spectra were recorded during three collection cycles in 0.5 nm steps at a spectral resolution of 1 nm. Air served as a reference medium in all measurements. The diameter of the measuring spot equaled 50 μm. To define band positions and intensities, fits of the recorded spectra were performed with the Jandel PeakFit software under the assumption of Gaussian peak shapes. The precision of the energies (positions) of the field-independent absorption bands is estimated to be better than 60 cm<sup>-1</sup>.

## RESULTS

### Crystal chemistry

The chemical analyses of our synthetic galaxite show, within analytical errors, an ideal MnAl<sub>2</sub>O<sub>4</sub> end-member composition (Table 1). The result of the structural refinement of this galaxite is summarized in Table 3. On the basis of scattering powers at the T and M sites, the following structural formula was refined: <sup>T</sup>(Mn<sub>0.90</sub>Al<sub>0.10</sub>)<sup>M</sup>(Mn<sub>0.10</sub>Al<sub>1.90</sub>)O<sub>4</sub>. This cation distribution explains why the observed tetrahedral (2.003 Å) and octahedral (1.931 Å) bond distances differ from the optimized bond distances of <sup>T</sup>Mn<sup>2+</sup> (2.036 Å) and <sup>M</sup>Al (1.908 Å) in spinel group minerals (Lavina et al. 2002). In our sample, disordering of Mn<sup>2+</sup> and Al over the T and M site equals 0.10 apfu. This value is much lower than 0.29–0.34 apfu reported by Greenwald et al. (1954) for synthetic galaxite. However, it is comparable with <sup>M</sup>Mn<sup>2+</sup> and <sup>T</sup>Al contents

in natural galaxites reported by Lucchesi et al. (1997) and Essene and Peacor (1983), respectively.

Chemical analyses (Table 1) of the Mn-bearing willemite shows a Mn<sup>2+</sup> content equal to 0.21 apfu, which may be disordered over the two irregular non-equivalent tetrahedra T1 and T2 (Sreekanth Chakradhar et al. 2004). Mn-rich forsterite shows comparable contents of Mn<sup>2+</sup> (0.93 apfu) and Mg (1.07 apfu). As demonstrated by Francis and Ribbe (1980), Mn<sup>2+</sup> orders strongly at the larger octahedral site (M2), while the Mg occupies the other non-equivalent octahedron (M1). In our tephroite, the Mn<sup>2+</sup> content is 1.93 apfu, which is expected to result in comparable Mn<sup>2+</sup> occupancies at the M1 and M2 octahedra. The chemical analyses of our rhodochrosite sample show a nearly complete Mn<sup>2+</sup> occupancy (0.98 apfu) at the octahedral M-site.

### Optical absorption spectra of tetrahedrally coordinated Mn<sup>2+</sup> in minerals

The optical absorption spectrum of a galaxite single crystal (Fig. 1) reveals a set of five relatively narrow absorption bands typical of spin-forbidden electronic d-d transitions, with the sharp and relatively strong field-independent d-d bands due to the <sup>6</sup>A<sub>1</sub>(S) → <sup>4</sup>E<sub>g</sub>(D) and <sup>6</sup>A<sub>1</sub>(S) → <sup>4</sup>E<sub>g</sub><sup>4</sup>A<sub>1g</sub>(G) transitions in Mn<sup>2+</sup> at 27780 and 23390 cm<sup>-1</sup>, respectively. The molar absorption coefficient of the band caused by the latter transition equals 1.90 L/(mol·cm). The field-dependent transitions <sup>6</sup>A<sub>1</sub>(S) → <sup>4</sup>T<sub>2g</sub>(D), <sup>6</sup>A<sub>1</sub>(S) → <sup>4</sup>T<sub>2g</sub>(G), and <sup>6</sup>A<sub>1</sub>(S) → <sup>4</sup>T<sub>1g</sub>(G) are observed at 25970, 22250, and 20300 cm<sup>-1</sup>, respectively. The spectrum of our end-member galaxite is, with the exception of more intense bands due to a higher Mn-concentration, very similar to that obtained for a synthetic high-temperature spinel sensu stricto doped with low concentrations, 0.001–0.050 apfu, of Mn<sup>2+</sup> (Jouini et al. 2006). However, their proposed assignments for the recorded Mn<sup>2+</sup>-bands in these spectra are not reasonable as they lead to an unrealistically low Racah parameter B (≈370 cm<sup>-1</sup>). Our spectrum shows no traces of Mn<sup>3+</sup>, which in spinel give rise to very strong

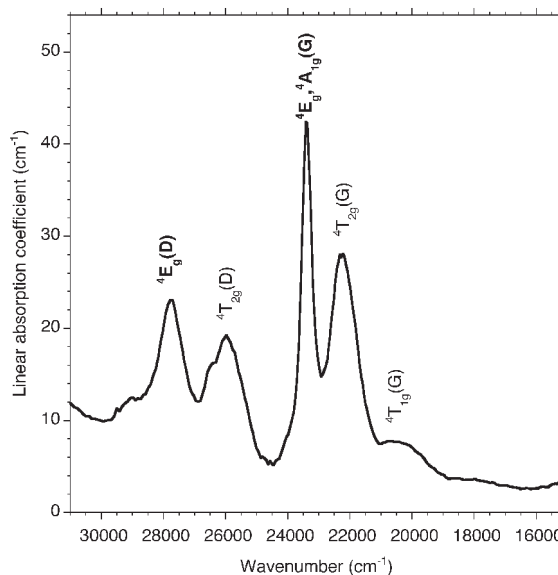
**TABLE 2.** Parameters for X-ray data collection

Determination of unit-cell parameters	
Radiation	MoKα <sub>1</sub> (0.70930 Å)
Reflections used	13 (Friedel pairs on both +2θ and -2θ)
Range	85–95 °2θ
Temperature	296 K
Diffraction intensity collection	
Radiation	MoKα (0.71073 Å)
Monochromator	High crystallinity graphite crystal
Range	3–95 °2θ
Reciprocal space range	0 ≤ h ≤ 17; 0 ≤ k ≤ 17; 0 ≤ l ≤ 17
Scan method	ω
Scan range	2.4°
Scan speed	Variable 1.502–14.648 °/min
Temperature	296 K
Data reduction	
Refinement	SHELXTL-PC
Corrections	Lorentz, Polarization
Absorption correction	Semi-empirical, 13ψ scans (10–95 °2θ)
Set of measured reflections	723
Set of unique reflections	151 [I > 2σ(I)]

**TABLE 3.** Results of the structural refinement of galaxite

a <sub>0</sub> (Å)	8.2104(3)
u	0.26588(7)
T-O (Å)	2.0034(6)
M-O (Å)	1.9310(5)
m.a.n. T	23.81(9)
m.a.n. M	13.60(6)
U <sub>11</sub> (T)	0.0060(2)
U <sub>11</sub> (M)	0.0053(2)
U <sub>11</sub> (O)	0.0090(2)
U <sub>12</sub> (M)	-0.0004(1)
U <sub>12</sub> (O)	-0.0003(2)
Ext.	0.0053(2)
R1(F) (%)	1.97

Notes: T-O and M-O = tetrahedral and octahedral bond distances, respectively. m.a.n. = mean atomic number; U<sub>11</sub> = displacement parameters (Å<sup>2</sup>); U<sub>11</sub> = U<sub>22</sub> = U<sub>33</sub> and U<sub>12</sub> = U<sub>13</sub> = U<sub>23</sub> (=0 for the T-site due to symmetry reasons); Ext. = isotropic secondary extinction coefficient, according to Larson's algorithm (1970); R1(F) in the form: (Σ|F<sub>obs</sub> - F<sub>calc</sub>|)/(ΣF<sub>obs</sub>). Space group: Fd3̄m, Z = 8.



**FIGURE 1.** Optical absorption spectrum of synthetic end-member galaxite with excited state terms for each band indicated.

and broad spin-allowed d-d bands in the range 20 000–25 000  $cm^{-1}$  (Bosi et al. 2007).

The polarized absorption spectra of willemite single crystals (Fig. 2) reveal band structures comparable to the one recorded in the galaxite spectrum. The sharp and intense field-independent  ${}^6A_1(S) \rightarrow {}^4E_g(D)$  and  ${}^6A_1(S) \rightarrow {}^4E_g^4A_{1g}(G)$  transitions in  $Mn^{2+}$  occur at 28 010 and 23 700  $cm^{-1}$ , respectively. These bands show molar absorption coefficients of 0.80 L/(mol·cm). The bands caused by the field-dependent transitions  ${}^6A_1(S) \rightarrow {}^4T_{2g}(D)$ ,  ${}^6A_1(S) \rightarrow {}^4T_{2g}(G)$ , and  ${}^6A_1(S) \rightarrow {}^4T_{1g}(G)$  are distinctly split. This feature, which we ascribe to the fact that  $Mn^{2+}$  is distributed over two non-equivalent tetrahedral sites (e.g., Sreekanth Chakradhar et al. 2004; Patra et al. 2005), was not noted in the unpolarized partial (18 000–25 000  $cm^{-1}$ ) absorption spectrum of willemite published by Germer et al. (2004). The assignment of the field-dependent bands is not unequivocal, but with regard to the difference in mean bond lengths of the two T-sites (Klaska et al. 1978) it is most likely that the bands at 21 230, 22 700, and 27 070  $cm^{-1}$  are caused by  $Mn^{2+}$  at the larger T2-tetrahedron, while the bands at 20 370, 23 120, and 26 320  $cm^{-1}$  mark transitions in  $Mn^{2+}$  at the smaller T1-tetrahedron. Alternative site assignments of these field-dependent bands yield, in view of the small size differences in bond distances for the two tetrahedra, unrealistically large differences in crystal field splitting parameters ( $Dq$ ).

#### Comparative optical absorption spectra of octahedrally coordinated $Mn^{2+}$ in minerals

Polarized optical absorption spectra of the carbonate rhodochrosite (Fig. 3) show simple structures comprising two intense and sharp bands at 29 160 and 24 540  $cm^{-1}$ , which mark the field-independent spin-forbidden transitions  ${}^6A_1(S) \rightarrow {}^4E_g(D)$  and  ${}^6A_1(S) \rightarrow {}^4E_g^4A_{1g}(G)$  in octahedrally coordinated  $Mn^{2+}$ . In addition, the three field-dependent bands caused by the  ${}^6A_1(S) \rightarrow {}^4T_{2g}(D)$ ,  ${}^6A_1(S) \rightarrow {}^4T_{2g}(G)$ , and  ${}^6A_1(S) \rightarrow {}^4T_{1g}(G)$  transitions in octahedrally coordinated  $Mn^{2+}$  occur at 27 620, 22 540, and

18 240  $cm^{-1}$ . This simple  $Mn^{2+}$ -band spectrum is in keeping with the simple structure of rhodochrosite, which comprises only one regular octahedral site with all Mn-O distances equal to 2.193 Å (Maslen et al. 1995). The recorded band energies are in very good agreement with those determined from high-resolution spectra on  $MnCO_3$  (Lohr and McClure 1968), but are distinctly lower (up to 200  $cm^{-1}$ ) than those determined by Keester and White (1968). Reasons for the higher band energies noted by the latter authors are not obvious, but may be related to different sample chemistry or spectral calibrations.

Distinct differences are observed in the present polarized absorption spectra of single crystals of Mn-rich forsterite and tephroite (Fig. 4). The spectra of Mn-rich forsterite show sharp bands caused by the spin-forbidden field-independent transi-

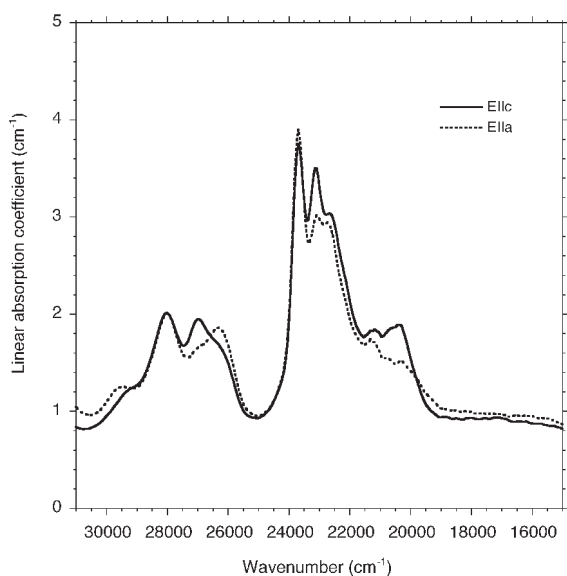


FIGURE 2. Polarized optical absorption spectra of Mn-bearing willemite.

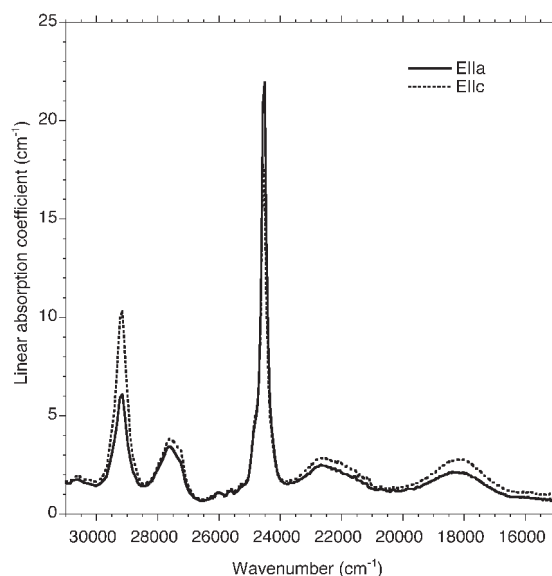


FIGURE 3. Polarized optical absorption spectra of rhodochrosite.

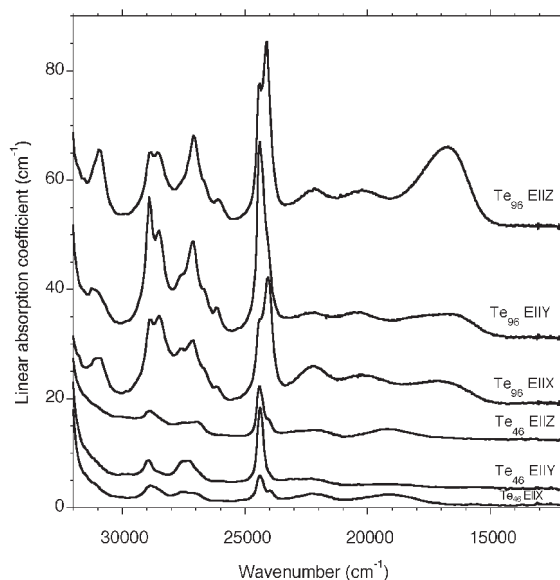


FIGURE 4. Polarized optical absorption spectra of a Mn-rich forsterite ( $Te_{46}$ ) and tephroite ( $Te_{96}$ ). For the sake of clarity, spectra are offset.

tions  ${}^6A_1(S) \rightarrow {}^4E_g(D)$  and  ${}^6A_1(S) \rightarrow {}^4E_g{}^4A_{1g}(G)$  in octahedrally coordinated  $Mn^{2+}$  at 28880 and 24400  $cm^{-1}$ , respectively. Weak shoulders are present ca 300  $cm^{-1}$  below these band maxima. In spectra of tephroite, these band shoulders are intensified and appear as strong sharp bands at 28520 and 24090  $cm^{-1}$ . The field-independent bands noted at higher energies in the spectra of the Mn-rich forsterite are retained in the tephroite spectra, which results in an apparent splitting of these bands (Fig. 5). The three field-dependent  $Mn^{2+}$ -bands are observed at 19120, 22180, and 27530  $cm^{-1}$  in the spectra of the Mn-rich forsterite, while three additional field-dependent bands occur at 16760, 20240, and 27100  $cm^{-1}$  in the tephroite spectra. Consequently, the strong Mn-ordering in olivine group minerals is nicely reflected in the spectra by the two fairly intense absorption bands caused by spin-forbidden transitions in  $Mn^{2+}$  at M2-octahedra in the Mn-rich forsterite and the evolution of two additional absorption bands related to  $Mn^{2+}$  at the M1-octahedron in tephroite spectra. The distinct shift of the field-independent spin-forbidden  $Mn^{2+}$ -bands toward lower energies in spectra with increasing Mn content in samples on the join forsterite-tephroite reported by Burns (1970) is not reproduced in the present spectra. This is explained by the lower spectral resolution in his spectra, which resulted in an apparent energy-shift of unresolved bands due to variations in the intensity ratio of the two band components related to  $Mn^{2+}$  at the M2 and M1-site. The very small composition driven shifts (ca. 60  $cm^{-1}$ ) of the  $Mn^{2+}$  bands related to the M1-site noted in the present spectra indicate strong structural relaxation in the olivine structure.

## DISCUSSION

### Crystal field splitting and Mn-O bond covalency

Information on crystal field splitting in  $Mn^{2+}$  and the degree of covalency of  $Mn^{2+}$ -L bonds can be derived from absorption spectra by solving the Tanabe-Sugano equations (Lever 1968) for the five spin-forbidden  $Mn^{2+}$ -bands ( $v_1$ - $v_5$ ) normally observed in the UV-Visual spectral range:

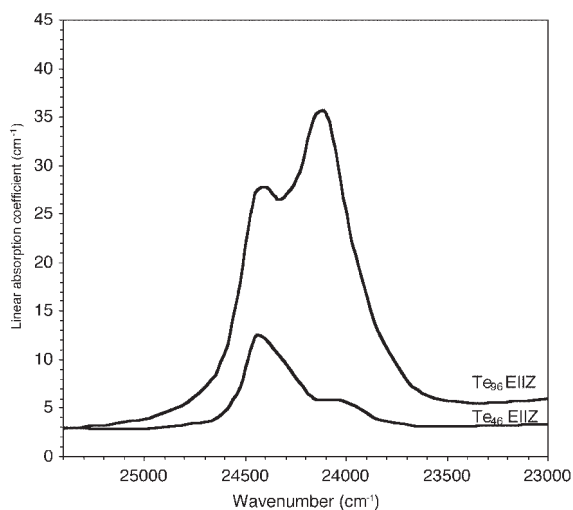


FIGURE 5. E||Z-spectra of the  ${}^6A_1(S) \rightarrow {}^4E_g{}^4A_{1g}(G)$  transition in octahedrally coordinated  $Mn^{2+}$  in Mn-rich forsterite ( $Te_{46}$ ) and tephroite ( $Te_{96}$ ).

$v_1$	${}^6A_1(S) \rightarrow {}^4T_{1g}(G)$	$-10Dq + 10B + 6C - 26B^2/10Dq$
$v_2$	${}^6A_1(S) \rightarrow {}^4T_{2g}(G)$	$-10Dq + 18B + 6C - 38B^2/10Dq$
$v_3$	${}^6A_1(S) \rightarrow {}^4E_g{}^4A_{1g}(G)$	$10B + 5C$
$v_4$	${}^6A_1(S) \rightarrow {}^4T_{2g}(D)$	$13B + 5C + x$
$v_5$	${}^6A_1(S) \rightarrow {}^4E_g(D)$	$17B + 5C$

where  $10Dq$  is the crystal field splitting and  $B$  and  $C$  are the interelectronic repulsion Racah parameters. To fully account for electronic interactions, additional terms, as the Trees ( $\alpha$ ) and the Seniority ( $\beta$ ) correction parameters, need to be introduced (e.g., Curie et al. 1974). In addition, access to low-temperature (below ca. 50 K) spectra providing information on the purely electronic zero phonon transitions would further improve the calculations. However, room temperature spectra in combination with the simplified expressions above may be applied for comparative semi-quantitative analyses of Me-L bonding.

Lower  $Dq$ - and  $B$ -values for tetrahedrally coordinated  $Mn^{2+}$  as compared to octahedrally coordinated  $Mn^{2+}$  is evident from Table 4, which summarizes the energies of the five  $Mn^{2+}$ -bands ( $v_1$ - $v_5$ ) in spectra of the present minerals and in several selected compounds from literature together with  $Dq$ -,  $B$ -, and  $C$ -values calculated by using the above expressions. The lower  $Dq$ - and  $B$ -values are consequences of lower coordination number, and thus smaller electronic Me-L repulsion, and a higher degree of bond covalency due to shorter Me-O distances in the tetrahedra. The differences are directly observed in the absorption spectra by the lower energies of the field-independent transitions ( $v_5$  and  $v_3$ ) and a smaller energy difference between these two bands ( $\Delta$ ) for  $Mn^{2+}$  in tetrahedral coordination (Fig. 6).

The crystal field splitting,  $10Dq$ , for  $Mn^{2+}$  at the tetrahedral site in galaxite equals ca. 5920  $cm^{-1}$ , which is ca. 67% of the corresponding value for octahedrally coordinated  $Mn^{2+}$  in rhodochrosite. This is in very good agreement with a theoretical value

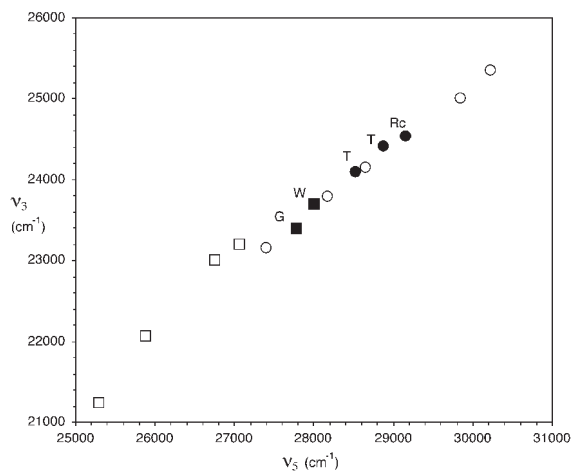


FIGURE 6. Plot of the energy of the  ${}^6A_1(S) \rightarrow {}^4E_g{}^4A_{1g}(G)$  band ( $v_3$ ) vs. the energy of the  ${}^6A_1(S) \rightarrow {}^4E_g(D)$  band ( $v_5$ ) in the  $Mn^{2+}$ -rich compounds and minerals summarized in Table 4. Squares show data for structures containing tetrahedrally coordinated  $Mn^{2+}$  and circles represent data for octahedrally coordinated  $Mn^{2+}$ . Filled symbols indicate data for the present samples, where G = galaxite, W = willemite, T = tephroite, and Rc = rhodochrosite.

**TABLE 4.**  $Mn^{2+}$ -bands, assignments, Racah (B and C), and crystal field splitting (Dq) parameters for divalent manganese in the present samples and in selected  $Mn^{2+}$ -rich compounds

Mineral/compound	$Mn^{2+}$ -cluster	$\nu_1$ ( $cm^{-1}$ )	$\nu_2$ ( $cm^{-1}$ )	$\nu_3$ ( $cm^{-1}$ )	$\nu_4$ ( $cm^{-1}$ )	$\nu_5$ ( $cm^{-1}$ )	Reference	B ( $cm^{-1}$ )	C ( $cm^{-1}$ )	Dq ( $cm^{-1}$ )
Galaxite (T-site)	$MnO_4$	20300	22250	23390	25970	27780	This study	627	3424	592
Willemite (T2-site)	$MnO_4$	21230	22700	23700	27070	28010	This study	616	3509	551
Willemite (T1-site)	$MnO_4$	20370	23120	23700	26320	28010	This study	616	3509	586
$MnSiN_2$	$MnN_4$	18380	21030	22070	24100	25890	Esmailzadeh et al. 2006	546	3323	638
$ZnS:Mn$	$MnS_4$	17891	19683	21237	22638	25297	Langer 1965	580	3087	609
$[(Me_4)N]_3[MnCl_4]$	$MnCl_4$	21200	22400	23200	26300	27070	Cotton et al. 1962	553	3534	527
$[(Et_4)N]_3[MnBr_4]$	$MnBr_4$	21350	22180	23000	25900	26750	Cotton et al. 1962	536	3529	511
Rhodochrosite	$MnO_6$	18240	22540	24540	27620	29160	This study	660	3588	879
Tephroite (M2-site)	$MnO_6$	19120	22180	24400	27530	28880	This study	640	3600	834
Tephroite (M1-site)	$MnO_6$	16760	20240	24090	27100	28520	This study	633	3552	1045
$MnF_2$	$MnF_6$	19440	23500	25350	28370	30230	Curie et al. 1974	697	3676	890
$Mn(ClO_4)_2 \cdot 6H_2O$	$Mn(H_2O)_6$	18690	22990	25000	27970	29850	Lohr and McClure 1968	693	3614	880
$MnSiF_6 \cdot 6H_2O$	$Mn(H_2O)_6$	18870	22990	25000	28090	29850	Lohr and McClure 1968	693	3614	869
$MnCl_2 \cdot 2H_2O$	$Mn(OH_2)_2Cl_4$	18980	21050	24130	26800	28650	Lohr and McClure 1968	646	3535	870
$[(CH_3)_4N][MnCl_3]$	$MnCl_6$	18540	22495	23785	26960	28170	Ono and Fuchikami 1977	626	3504	763
$CsMnBr_3$	$MnBr_6$	18180	21750	23150	26450	27400	McPherson et al. 1974	607	3416	745

of ca. 70%, which is based on the differences in Me-O distances in the two structures and a decreased 3d-electron repulsion (4/9 of the octahedral case) in the tetrahedron. The calculated slightly lower B-parameter,  $627\text{ cm}^{-1}$ , for  $Mn^{2+}$  at the tetrahedral site in galaxite is in keeping with shorter  $Mn^{2+}$ -O bonds and thus a higher degree of bond covalency than for  $Mn^{2+}$  in octahedra.

### Band intensity

Absorption bands caused by spin-forbidden transitions in  $Mn^{2+}$  in unbridged octahedra are generally very weak. The molar absorption coefficients ( $\epsilon$ ) recorded for such absorption bands are normally  $\leq 0.25\text{ L}/(\text{mol}\cdot\text{cm})$  (e.g., Lohr and McClure 1968). However, magnetic coupling between structurally adjacent  $Mn^{2+}$  pairs in corner-, edge- or face-sharing octahedra has been demonstrated to result in intensified spin-forbidden absorption bands in various salts (e.g., Ferguson et al. 1965, 1966; Lohr and McClure 1968). Similar effects have been observed for spin-forbidden absorption bands in spectra of other  $3d^5$ -cations as, e.g., in  $Fe^{3+}$ -sulphate minerals (Rossman 1975). Absorption bands caused by spin-forbidden transitions in  $3d^5$ -cations in tetrahedra, which lack a center of symmetry, are expected to intensify by one to two orders of magnitude due to relaxation of the Laporte selection rule caused by mixing of p- and d-orbitals (e.g., Burns 1993).

The  $\epsilon$ -values recorded in the present spectra for the relatively strong and sharp field-independent band caused by the  ${}^6A_1(S) \rightarrow {}^4E_g{}^4A_{1g}(G)$  transition in octahedrally coordinated  $Mn^{2+}$  are invariably higher than those expected (Table 5). This may be caused by band intensification due to magnetically coupled  $Mn^{2+}$  pairs in corner and/or edge-sharing polyhedra in rhodochrosite and olivine group minerals (Table 6). In addition, the recorded  $\epsilon$ -values for the M2-related field-independent band increases from  $0.65\text{ L}/(\text{mol}\cdot\text{cm})$  in spectra of Mn-rich forsterite to  $1.50\text{ L}/(\text{mol}\cdot\text{cm})$  in tephroite spectra (Table 5). This intensity enhancement is most likely related to transitions in coupled  $Mn^{2+}$  pairs in edge-sharing M1-M2 octahedra in tephroite. The number of such pairs is on the basis of cation occupancy determinations (Francis and Ribbe 1980) expected to be approximately one order of magnitude higher in the tephroite than in the Mn-rich forsterite.

Molar absorption coefficients for the field-independent band caused by the  ${}^6A_1(S) \rightarrow {}^4E_g{}^4A_{1g}(G)$  transition in tetrahedrally coordinated  $Mn^{2+}$  in galaxite and Mn-bearing willemite equals

**TABLE 5.** Molar absorption coefficients for the  ${}^6A_1(S) \rightarrow {}^4E_g{}^4A_{1g}(G)$  band in the present spectra

Mineral	site	$\epsilon$ [ $L/(\text{mol}\cdot\text{cm})$ ]
Galaxite	T	1.90
Willemite	T1+T2	0.80
Rhodochrosite	M	0.68
Forsterite ( $Te_{48}$ )	M2	0.65
Tephroite ( $Te_{96}$ )	M1	1.60
Tephroite ( $Te_{96}$ )	M2	1.50

**TABLE 6.** Cation site sharing in the investigated minerals

Mineral	Sites	Sharing	Distance ( $\text{\AA}$ )	Reference
Galaxite	T-M	corner	3.404	This study
Willemite	T1-T1	corner	3.216	Klaska et al. 1978
	T2-T2	corner	3.230	
	T1-T2	corner	3.112	
	T1-T2	corner	3.239	
Rhodochrosite	M-M	corner	3.793	Maslen et al. 1995
Forsterite ( $Fe_{52}Te_{48}$ )	M1-M1	edge	3.062	Francis and Ribbe 1980
	M1-M2	edge	3.304	
	M1-M2	corner	3.629	
	M1-M2	corner	3.710	
	M2-M2	corner	3.936	
Tephroite ( $Fe_{99}Te_{91}$ )	M1-M1	edge	3.117	Francis and Ribbe 1980
	M1-M2	edge	3.351	
	M1-M2	corner	3.676	
	M1-M2	corner	3.755	
	M2-M2	corner	4.009	

$1.90$  and  $0.80\text{ L}/(\text{mol}\cdot\text{cm})$ , respectively, which is within the range expected for normal single ion transition bands due to tetrahedrally coordinated  $Mn^{2+}$ . Considerably higher  $\epsilon$ -values,  $\sim 50\text{ L}/(\text{mol}\cdot\text{cm})$ , for the corresponding absorption band have been observed in spectra of  $MnSiN_2$  (Esmailzadeh et al. 2006), which contain a network of magnetically ordered corner-sharing  $MnN_4$ -tetrahedra and consequently possesses a structural framework for band intensification through coupling effects. Formation of  $Mn^{2+}$  pairs in corner-sharing tetrahedra in willemite has been suggested from time-resolved photoluminescence spectra (e.g., Patra et al. 2005). Strongly intensified absorption bands caused by electronic transitions in coupled  $Fe^{3+}$  pairs in corner-sharing octahedra and tetrahedra have been observed in spectra of spinel-magnesioferite solution crystals (Andreozzi et al. 2001). The crystal structures of galaxite and willemite contain corner-sharing cation sites (Table 6) that may facilitate pair coupling. Consequently, electronic transitions in coupled

Mn<sup>2+</sup>-pairs in galaxite and willemite as the cause for intensified optical absorption bands cannot be ruled out.

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